

Sterile Neutrino Searches with MINOS+

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Birmingham HEP Seminar
29/04/20



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Outline

- Introduction
- MINOS+ Experiment
- Three flavour oscillation results
- Four flavour oscillation results

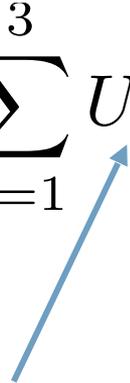
Neutrinos

- Neutrino oscillations have become a well-established and well-described phenomenon over the last 20 years.
 - The Nobel Prize in Physics 2015 was awarded jointly to Takaaki Kajita and Arthur B. McDonald "for the discovery of neutrino oscillations, which shows that neutrinos have mass"
- Oscillations arise from the quantum mechanical interference between the neutrino mass states.
 - At least two of the neutrinos must be massive!
- The neutrino eigenstates of the weak interaction are not the same as the mass eigenstates.

Neutrinos

- For three neutrino flavours:

$$\nu_\alpha = \sum_{i=1}^3 U_{\alpha i} \nu_i$$

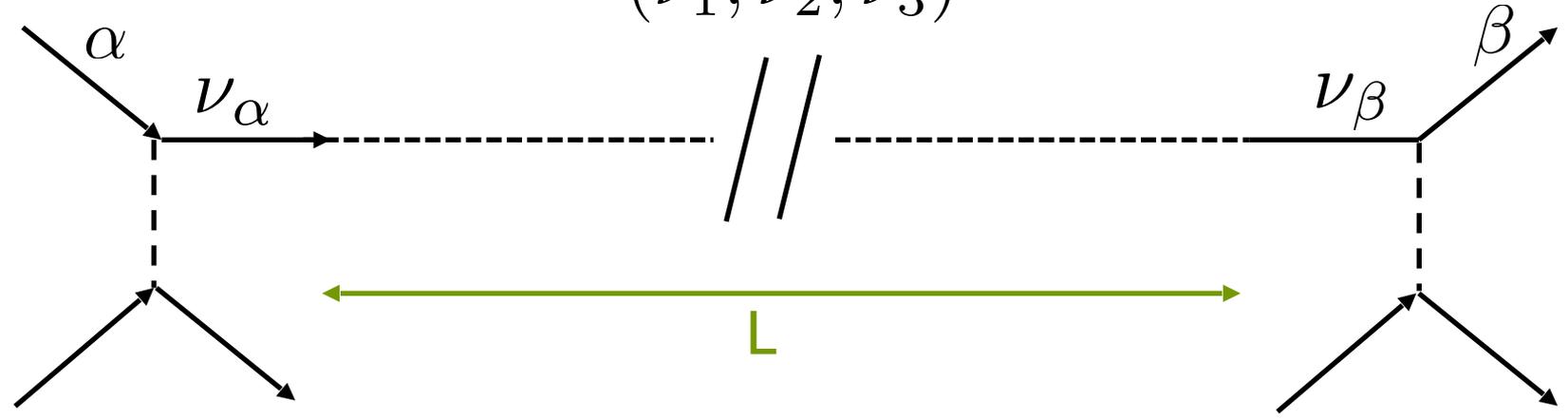


Flavour eigenstates: ν_e, ν_μ, ν_τ

Mass eigenstates: ν_1, ν_2, ν_3

3x3 unitary matrix - the PMNS matrix

(ν_1, ν_2, ν_3)



Neutrinos

- For three neutrino flavours: $c_{jk} = \cos \theta_{jk}, s_{jk} = \sin \theta_{jk}$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Muon neutrino disappearance
(accelerator and atmospheric)

Electron antineutrino
disappearance (reactor)
Electron neutrino appearance
(accelerator)

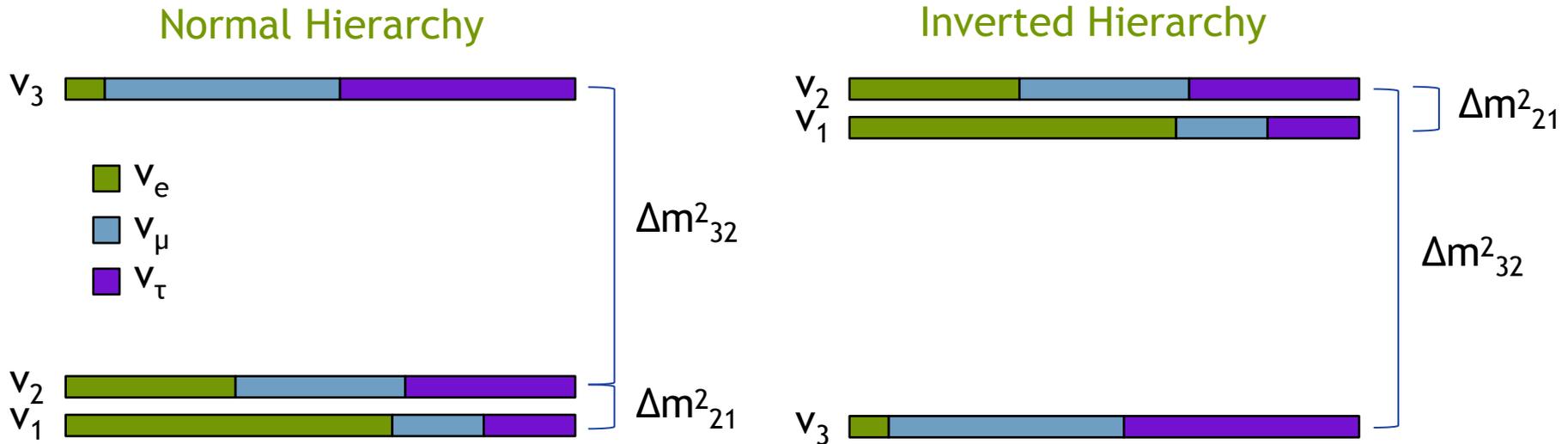
(Anti)electron neutrino
disappearance
(solar and reactor)

- Three mixing angles and a CP violating phase.
- Oscillations are driven by mass-squared splittings

$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \sin^2 2\theta_{23} \sin^2 \left(\frac{1.27 \Delta m_{31}^2 L}{E} \right)$$

Mass Hierarchy

- The order of all the mass states isn't completely known.



- The sign of Δm^2_{21} is known from matter effects in the Sun and from the definition of ν_1 having the largest ν_e component.
- The sign of Δm^2_{32} is still unknown.

Current State of Measurements

- Very successful programme of measurements.

Nu-Fit v4.0

- The remaining unknowns:

- Is the mass-hierarchy

- Normal $\Delta m_{32}^2 > 0$?
- Inverted $\Delta m_{32}^2 < 0$?

- Is $\theta_{23} = 45^\circ$?

- If not, is it higher or lower?

- What is the value of δ ?

- Is there CP violation in the neutrino sector?

- How many neutrinos are there?

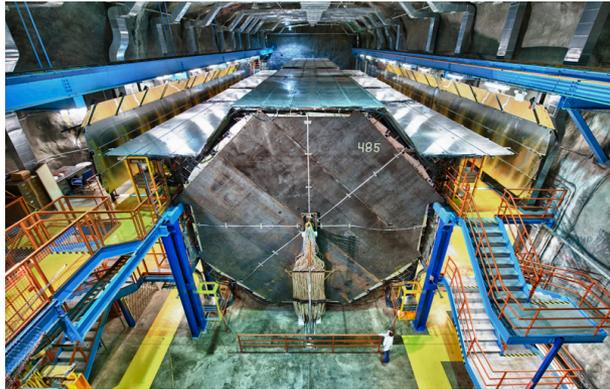
		Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 4.7$)	
		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range
without SK-atm	$\sin^2 \theta_{12}$	$0.310^{+0.013}_{-0.012}$	0.275 \rightarrow 0.350	$0.310^{+0.013}_{-0.012}$	0.275 \rightarrow 0.350
	$\theta_{12}/^\circ$	$33.82^{+0.78}_{-0.76}$	31.61 \rightarrow 36.27	$33.82^{+0.78}_{-0.76}$	31.61 \rightarrow 36.27
	$\sin^2 \theta_{23}$	$0.580^{+0.017}_{-0.021}$	0.418 \rightarrow 0.627	$0.584^{+0.016}_{-0.020}$	0.423 \rightarrow 0.629
	$\theta_{23}/^\circ$	$49.6^{+1.0}_{-1.2}$	40.3 \rightarrow 52.4	$49.8^{+1.0}_{-1.1}$	40.6 \rightarrow 52.5
	$\sin^2 \theta_{13}$	$0.02241^{+0.00065}_{-0.00065}$	0.02045 \rightarrow 0.02439	$0.02264^{+0.00066}_{-0.00066}$	0.02068 \rightarrow 0.02463
	$\theta_{13}/^\circ$	$8.61^{+0.13}_{-0.13}$	8.22 \rightarrow 8.99	$8.65^{+0.13}_{-0.13}$	8.27 \rightarrow 9.03
	$\delta_{CP}/^\circ$	215^{+40}_{-29}	125 \rightarrow 392	284^{+27}_{-29}	196 \rightarrow 360
	$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.39^{+0.21}_{-0.20}$	6.79 \rightarrow 8.01	$7.39^{+0.21}_{-0.20}$	6.79 \rightarrow 8.01
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.525^{+0.033}_{-0.032}$	+2.427 \rightarrow +2.625	$-2.512^{+0.034}_{-0.032}$	-2.611 \rightarrow -2.412

Esteban, I., Gonzalez-Garcia, M.C., Hernandez-Cabezudo, A. et al. J. High Energ. Phys. (2019) 2019: 106. [https://doi.org/10.1007/JHEP01\(2019\)106](https://doi.org/10.1007/JHEP01(2019)106)

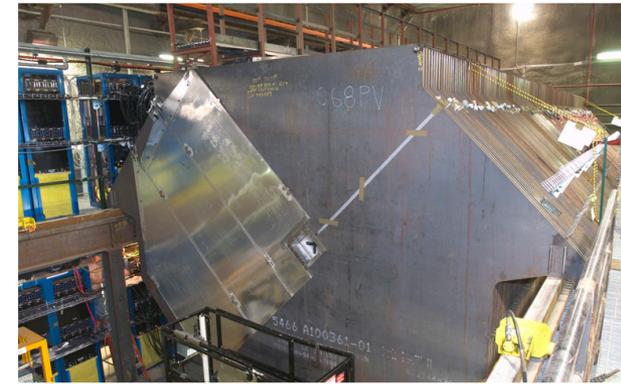


MINOS and MINOS+

The MINOS+ Experiment



Far Detector
735 km from beam target
5.4 kton mass



Near Detector
1 km from beam target
1 kton mass

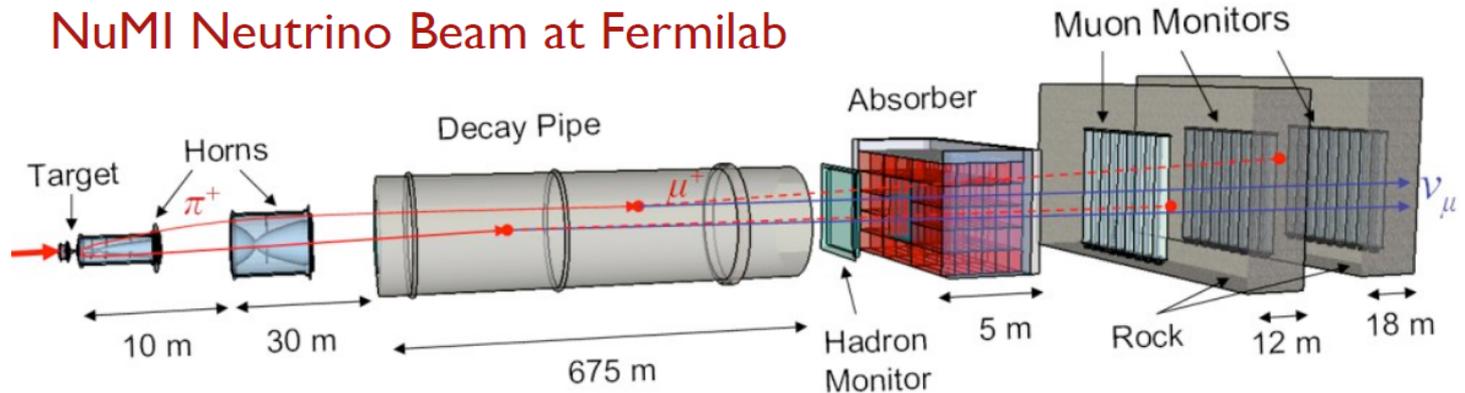
- MINOS/MINOS+ had two functionally identical, magnetised, tracking, sampling calorimeters.
 - Can distinguish muon charge from the curvature.
- Exposed by the NuMI beam at Fermilab.
- MINOS+ is the continuation of MINOS into the NOvA era at FNAL.

The NuMI Beam

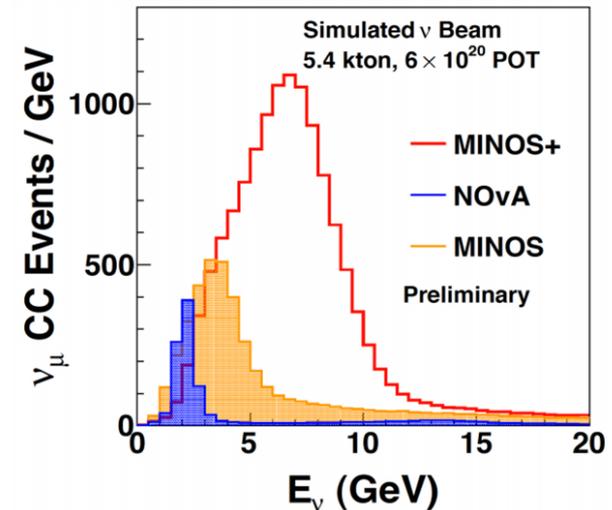
- MINOS+ collected neutrinos from the NuMI beam at Fermilab.

NuMI Neutrino Beam at Fermilab

120 GeV
protons from
Main Injector

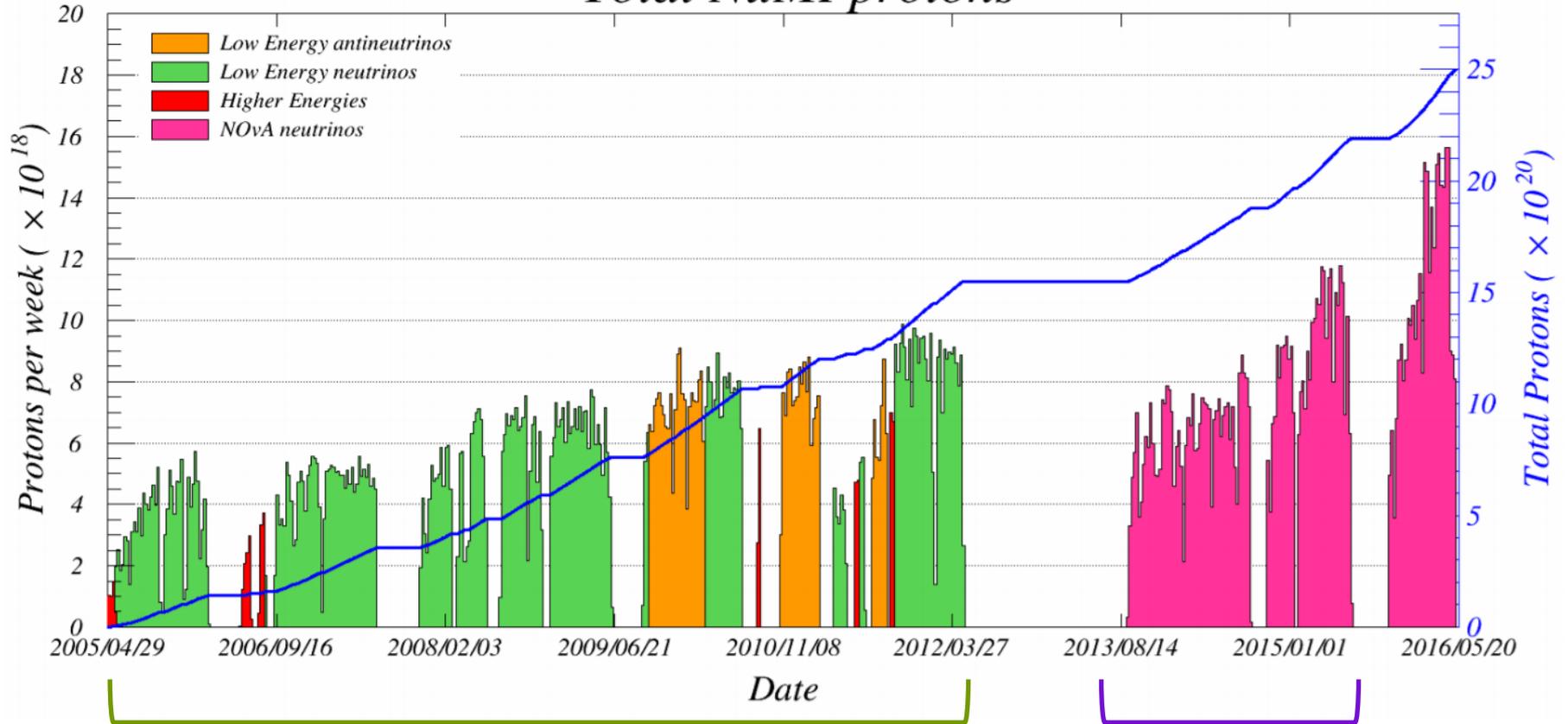


- Neutrinos produced by decay of focused mesons produced in the target.
- Polarity of the horns can be reversed to produce an antineutrino beam.



Data Samples

Total NuMI protons



10.56 x 10²⁰ POT MINOS

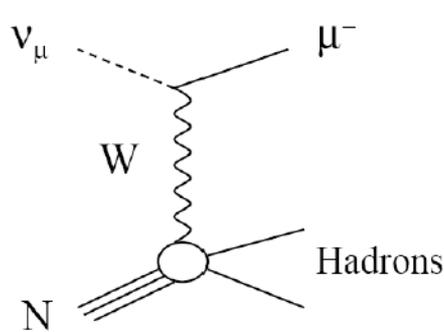
5.80 x 10²⁰ POT MINOS+

- Results shown today use all MINOS and 2/3 years of MINOS+ data

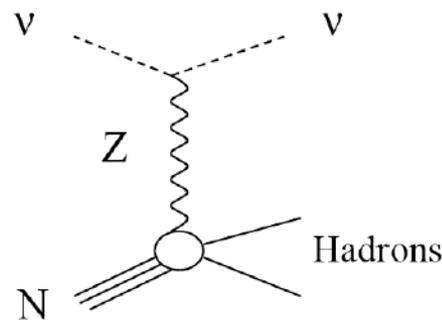
Neutrino Interactions in MINOS+

- There are three main types of interactions seen in MINOS+

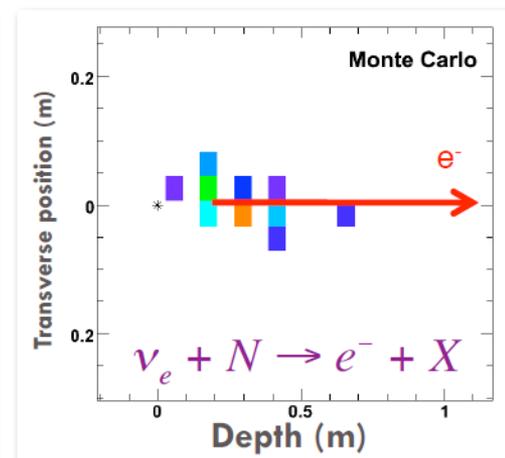
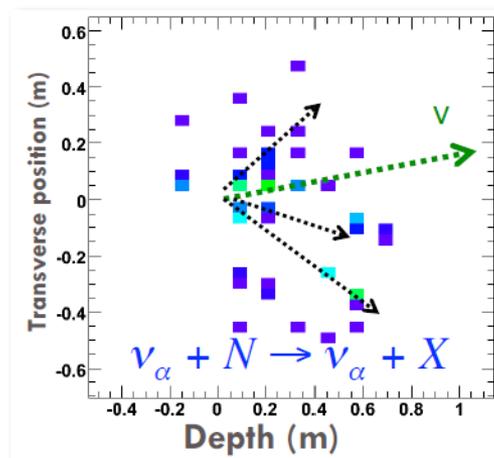
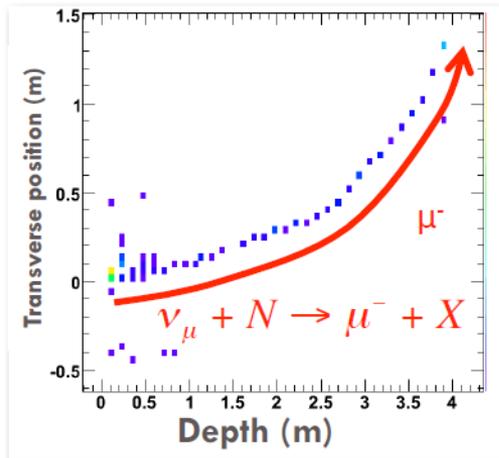
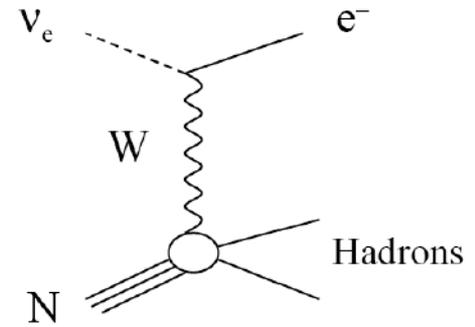
ν_μ charged-current



ν neutral-current

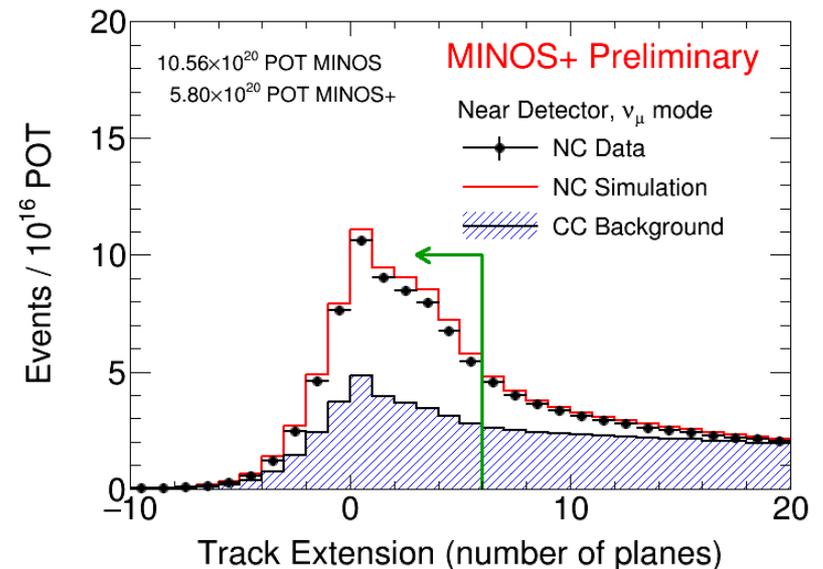
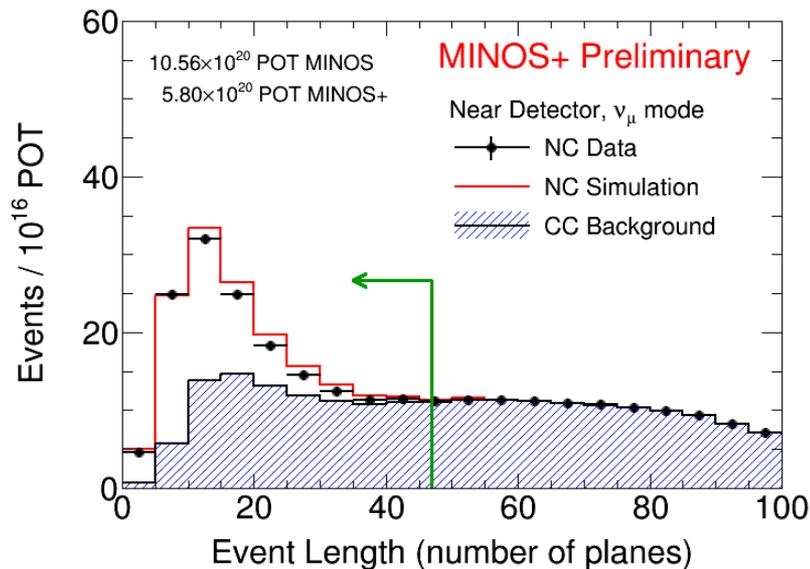


ν_e charged-current



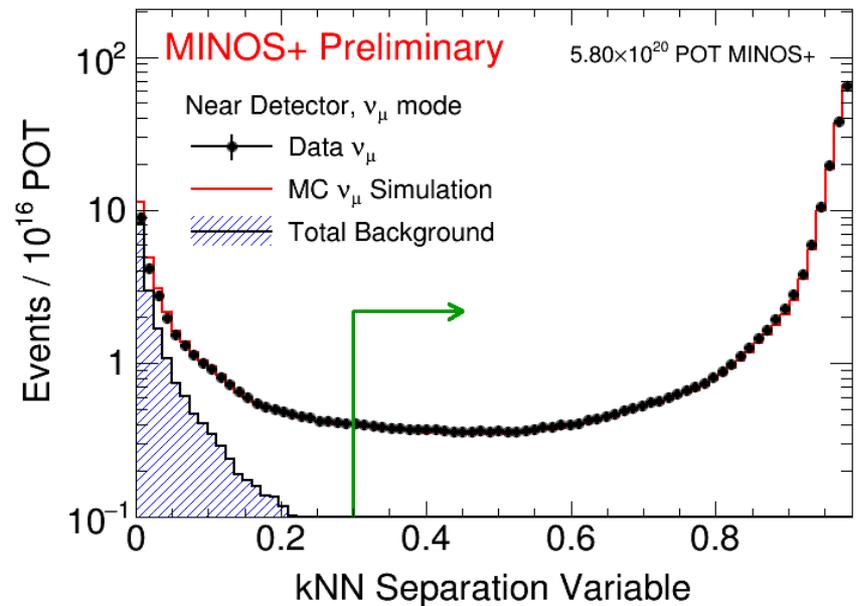
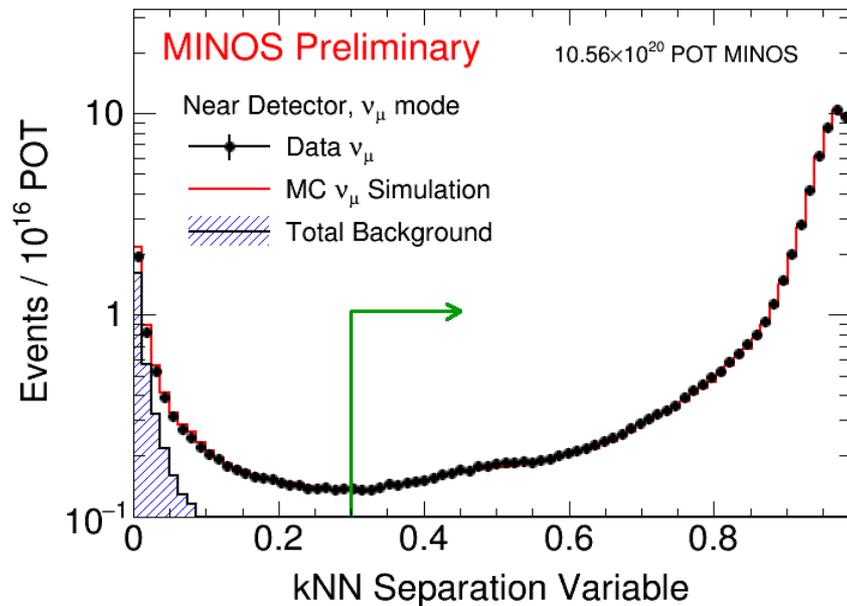
NC Event Selection

- The first step is to select the neutral current interactions.
- Two main selection criteria:
 - Event length and the extension of the track beyond the hadronic shower.



CC Event Selection

- Charged current interactions are selected from those that do not pass the neutral current selection.
 - Use a kNN to select CC interactions from the backgrounds.
 - Uses four topological and energy deposition variables as input.

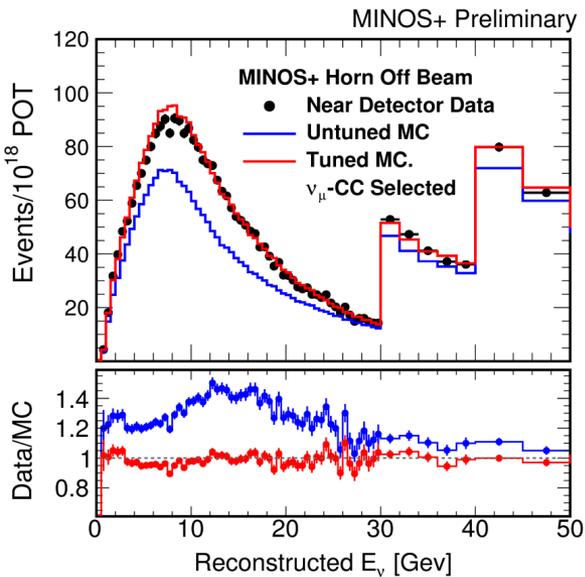




Three Flavour Oscillation Analysis

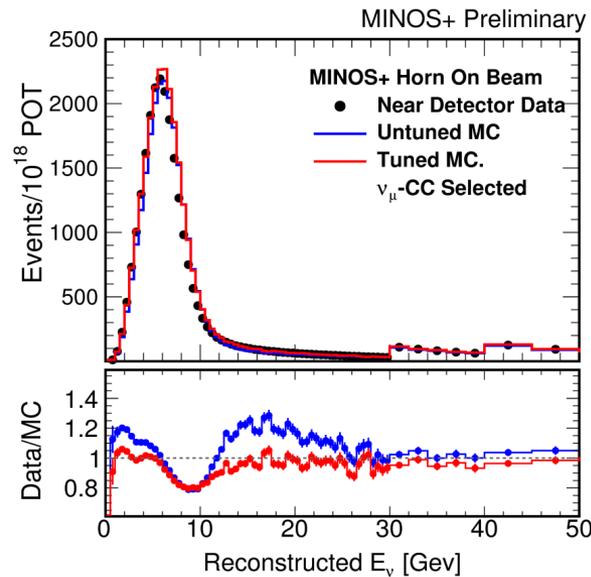
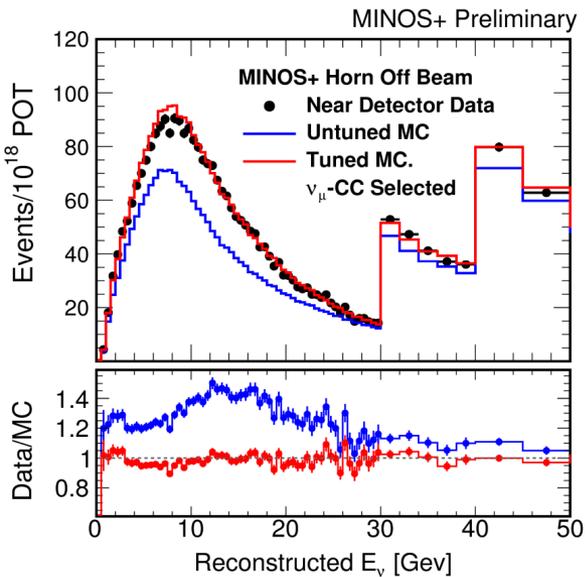
Beam Neutrinos - Flux Prediction

- In our three-flavour analysis we use the ND to tune the MC
- A special sample with the magnetic horns switched off allows us to probe hadron production effects



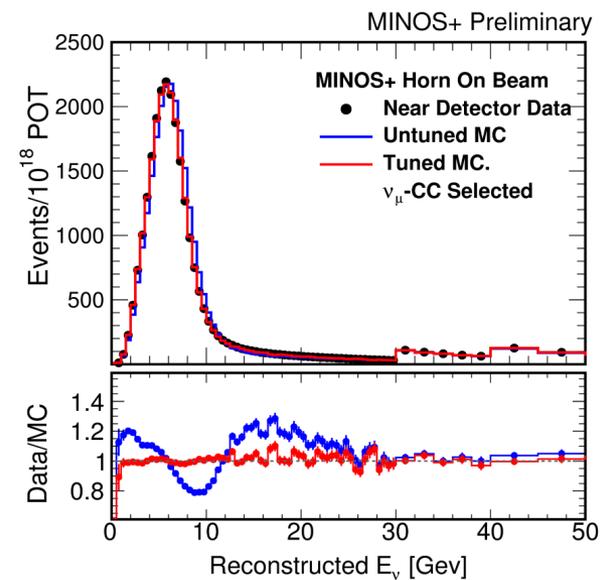
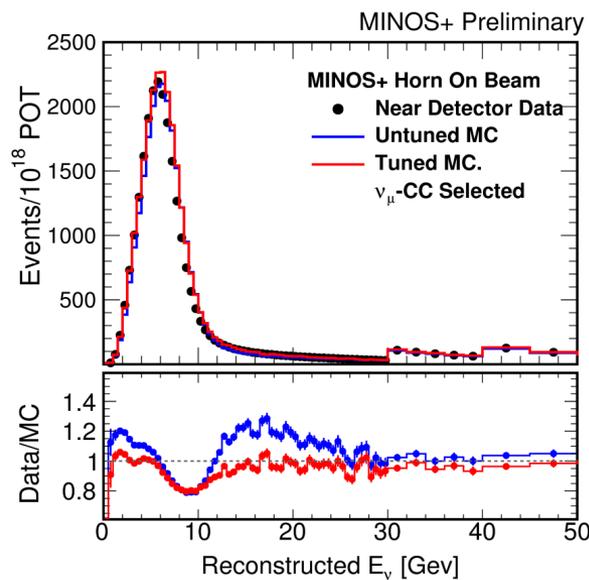
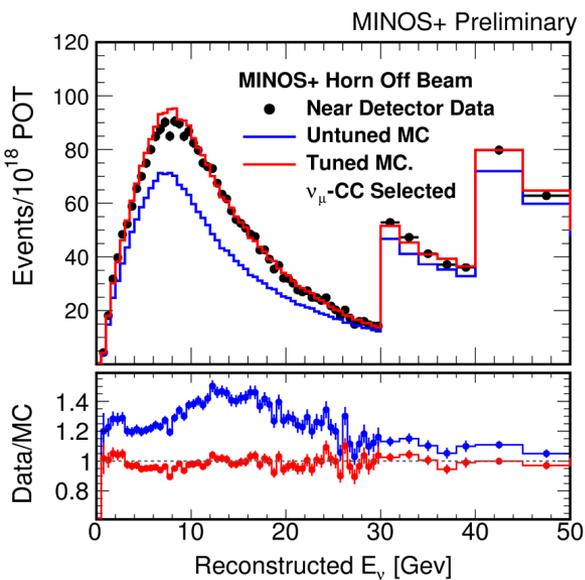
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- We then apply these hadron production weights to the standard horn on MC



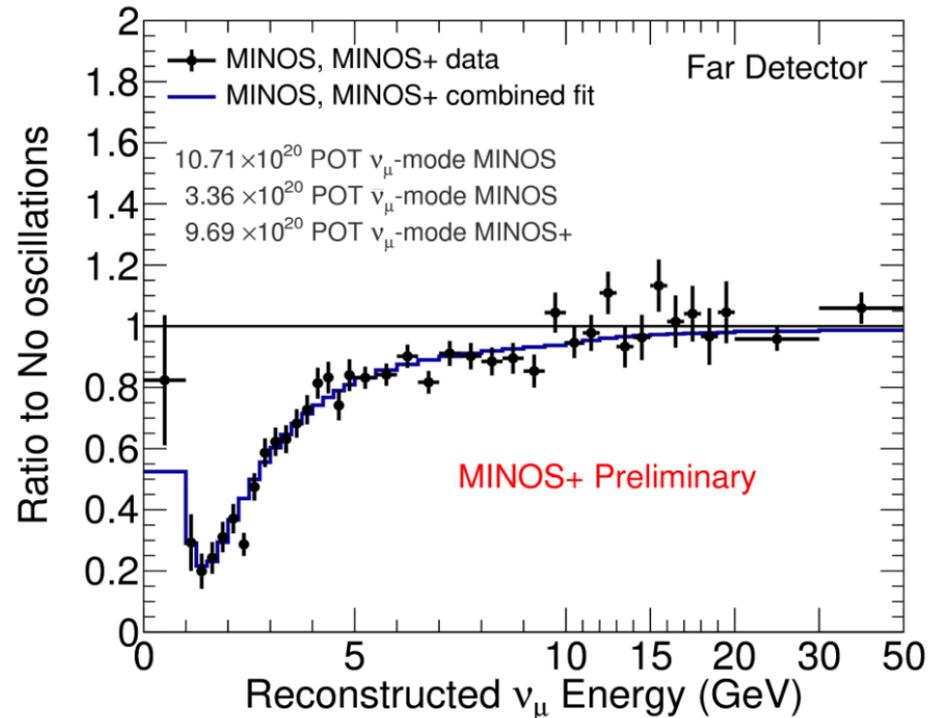
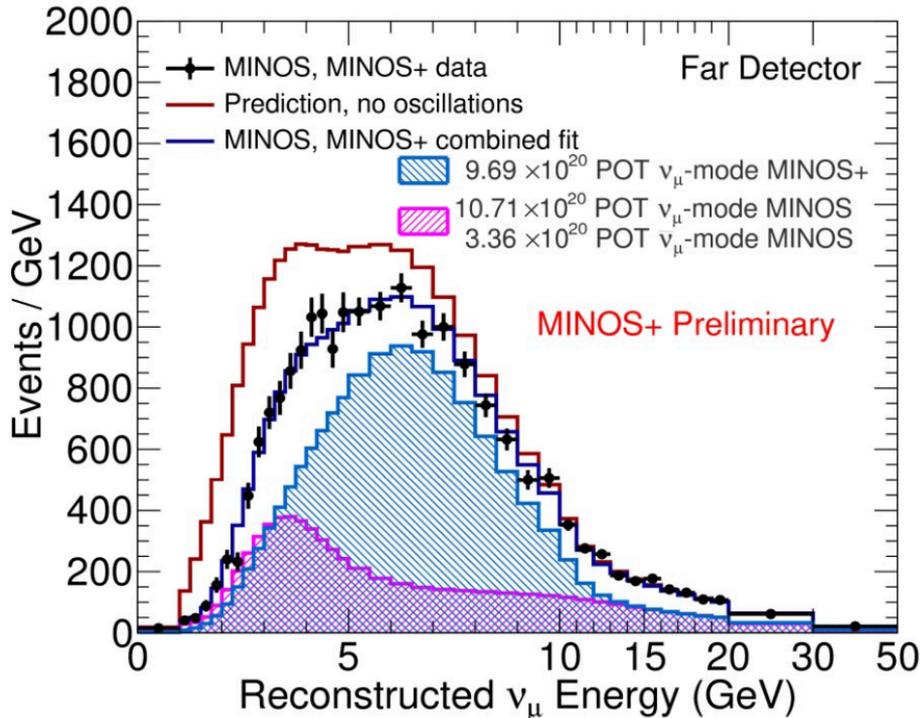
Beam Neutrinos - Flux Prediction

- In our three-flavour analysis we use the ND to tune the MC
- Finally, we fit the standard hour on MC to tune the beam focussing component of the flux prediction



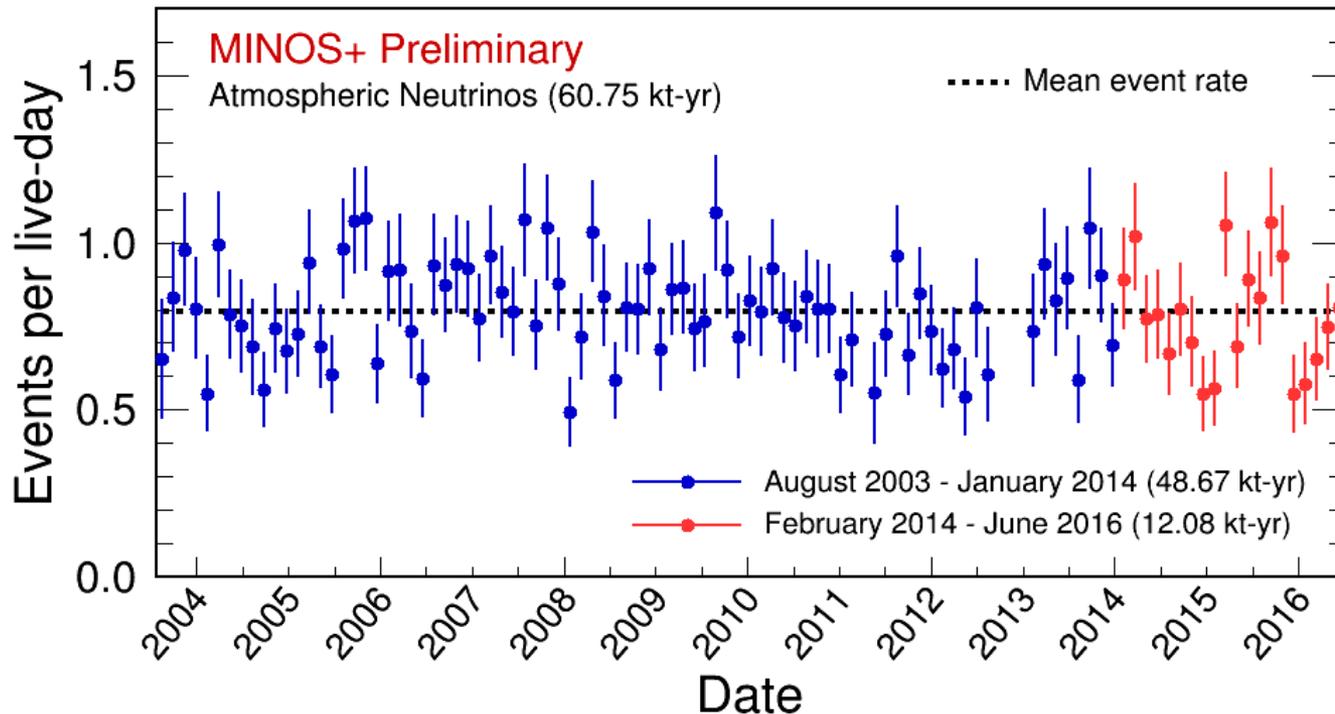
Beam Neutrinos

- MINOS was designed to measure the atmospheric scale oscillation parameters.
 - Look for disappearance of muon neutrinos in the FD relative to ND.
 - Measure muon neutrinos through charged current interactions.



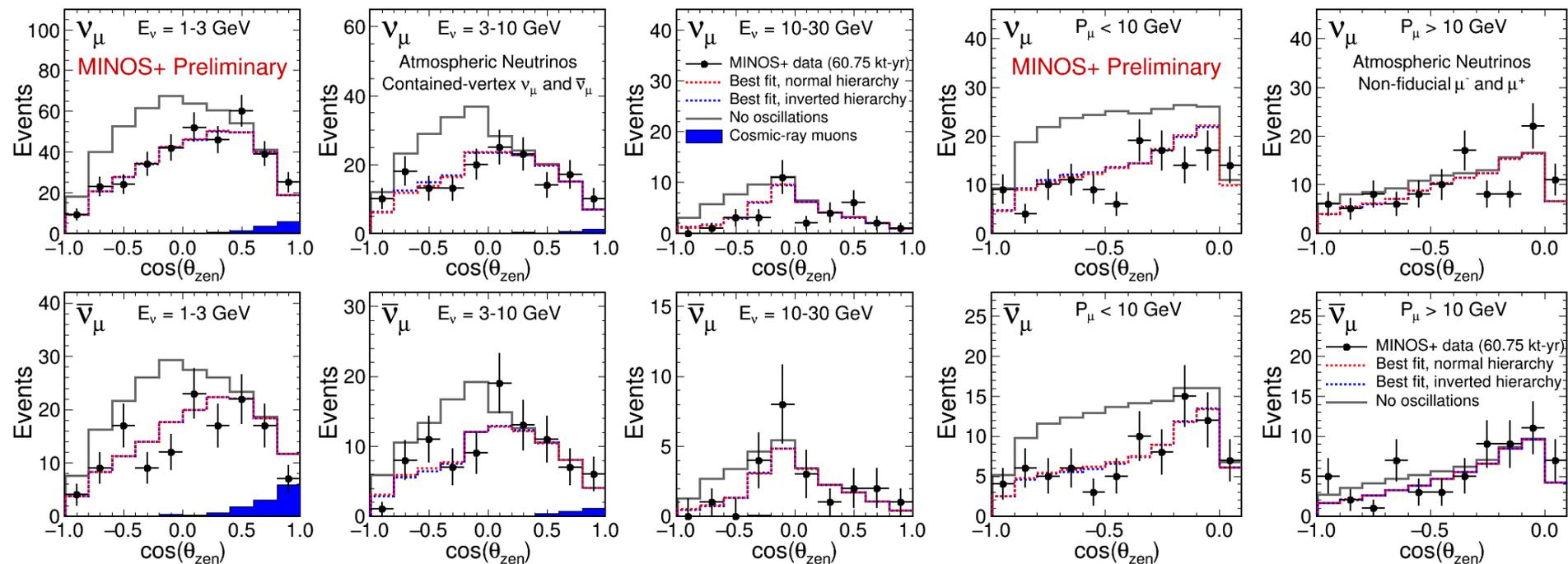
Atmospheric Neutrinos

- The MINOS+ Far Detector has collected a large number of atmospheric neutrinos over 12 years
 - Neutrinos and anti-neutrinos separated by curvature in the magnetic field



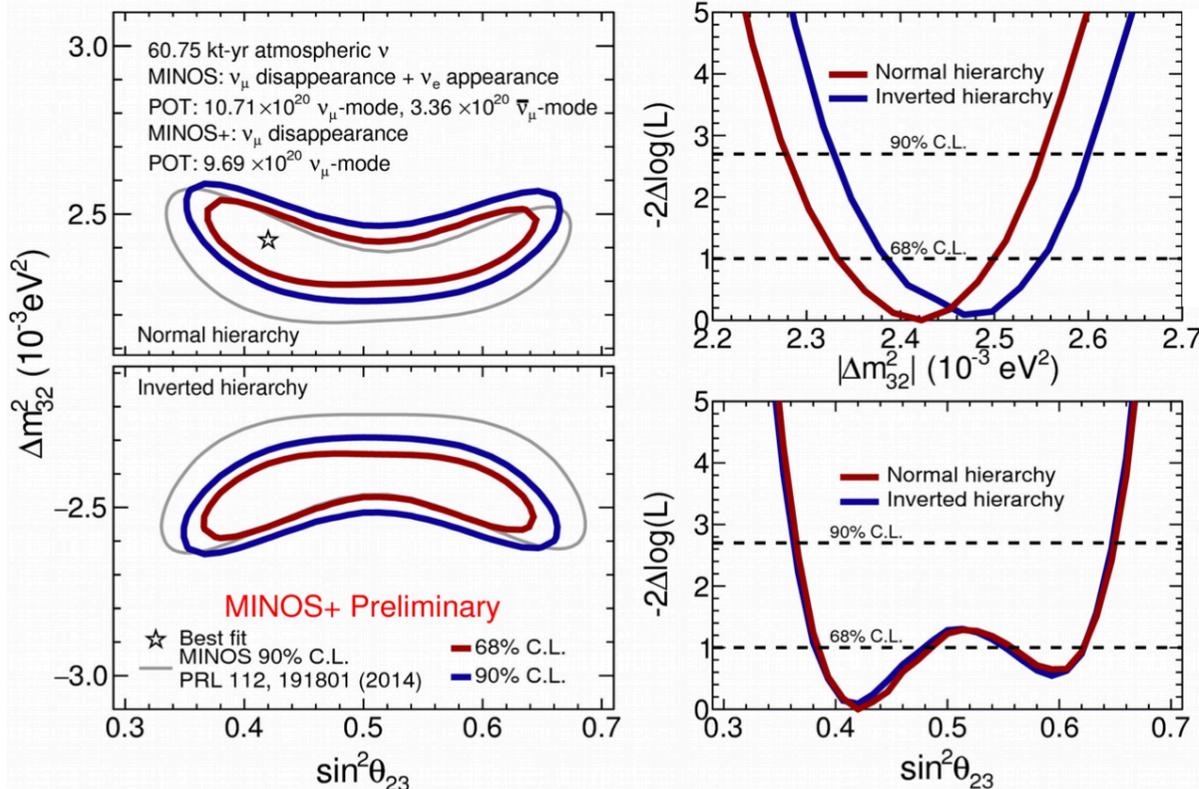
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Three Flavour Oscillations

- Fit gives 1D and 2D contours



Best fit

$$\Delta m_{32}^2 = 2.42 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} = 0.42$$

$|\Delta m_{32}^2|$ 90% C.L. intervals

NH: $(2.28 - 2.55) \times 10^{-3} \text{ eV}^2$

IH: $(2.33 - 2.60) \times 10^{-3} \text{ eV}^2$

Measured to $\sim 3.5\%$ at 68% C.L.

$\sin^2 \theta_{23}$ 90% C.L. interval

0.36 - 0.65

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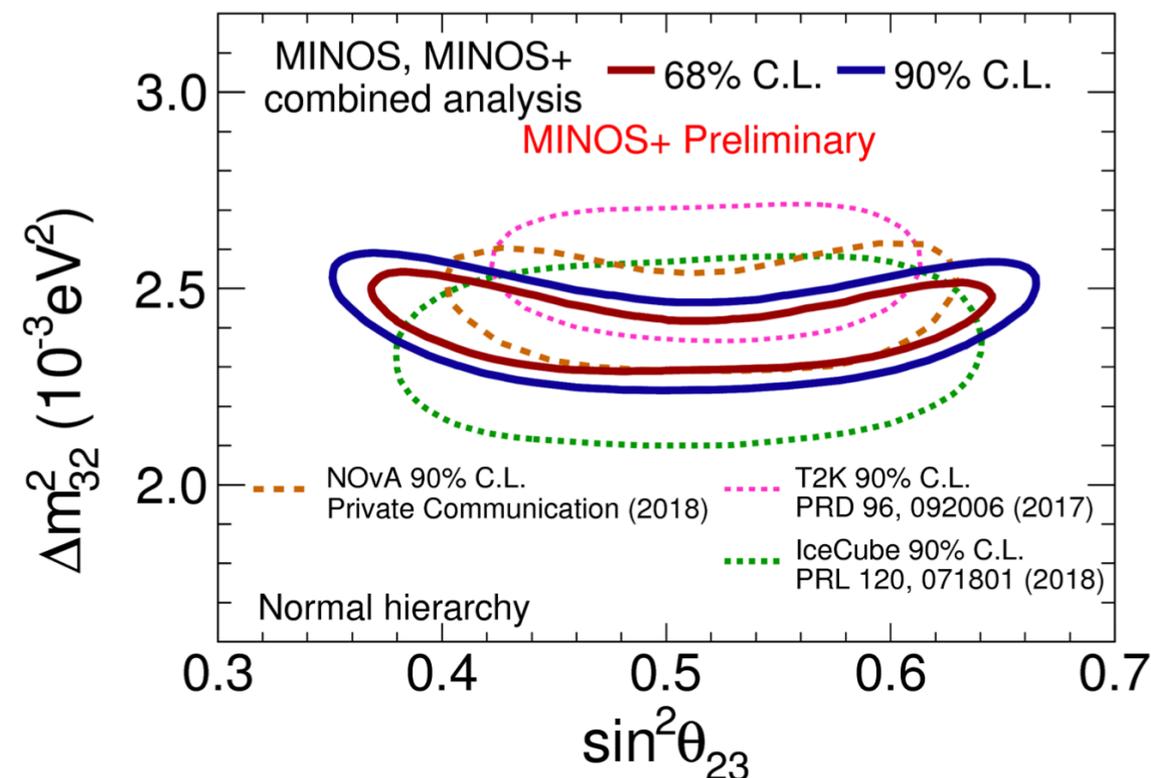
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Beyond Three Neutrino Flavours

How Many Neutrinos?

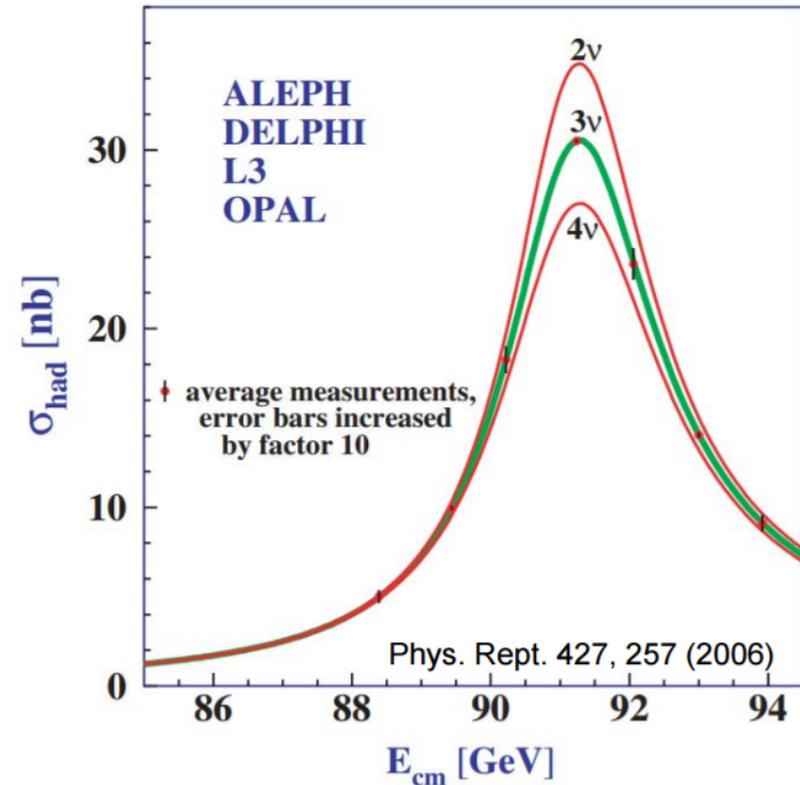
- Invisible width of the Z-boson from LEP very strongly measured that there are 3 neutrinos.

- For $m_\nu < \frac{1}{2}m_Z$ fourth neutrino must not couple to the Z-boson.
 - Hence the name *sterile*.

- Results from Planck:

$$\left. \begin{array}{l} N_{\text{eff}} = 3.2 \pm 0.5 \\ \sum m_\nu < 0.32 \text{ eV} \end{array} \right\} 95\%, \text{ Planck TT+lowP+lensing+BAO.}$$

P. A. R. Ade, et al. (2016) Astron. Astrophys. 594, arXiv 1502.01589

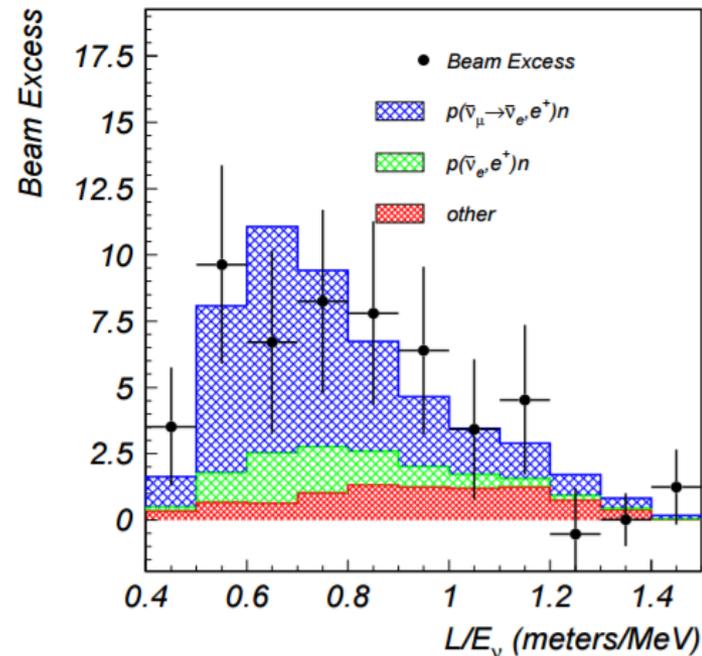
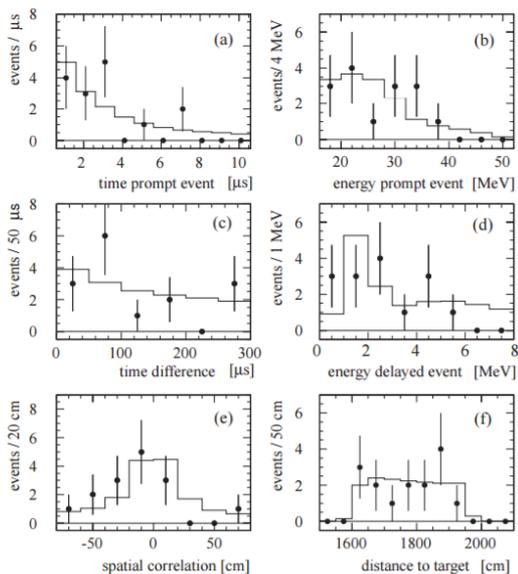


Some Anomalies

- The majority of neutrino oscillation data is well described by the three flavour model.
 - However, there are some outliers.
- Anomalous appearance of ν_e in short-baseline ν_μ beams.
- Gallium experiment calibration sources.
- Reactor neutrino flux deficit.
- The main point is that all three anomalies were consistent with oscillations at a mass-splitting scale of approximately 1 eV^2

Some Anomalies - 1

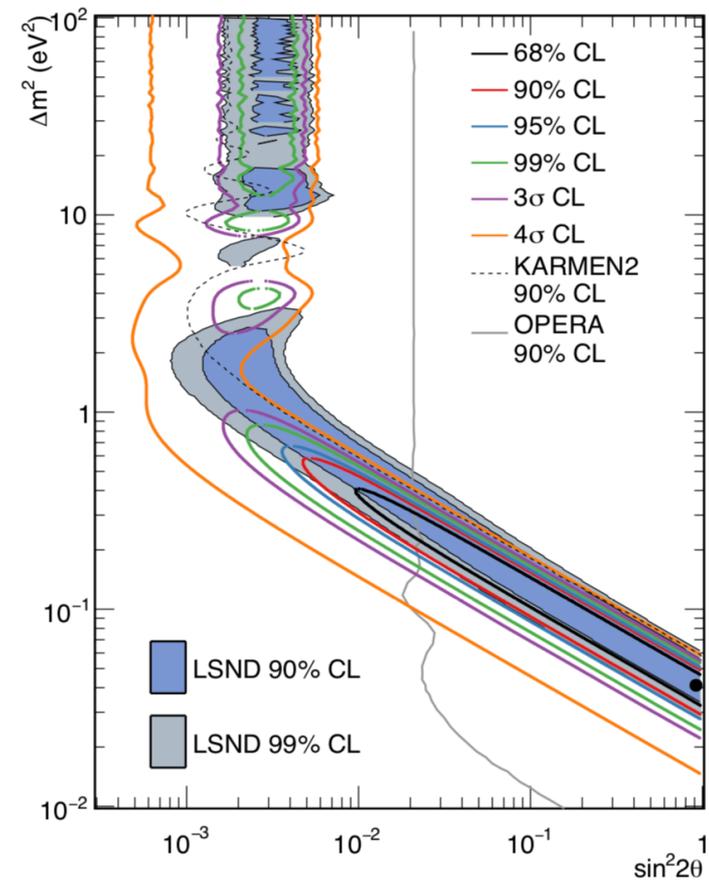
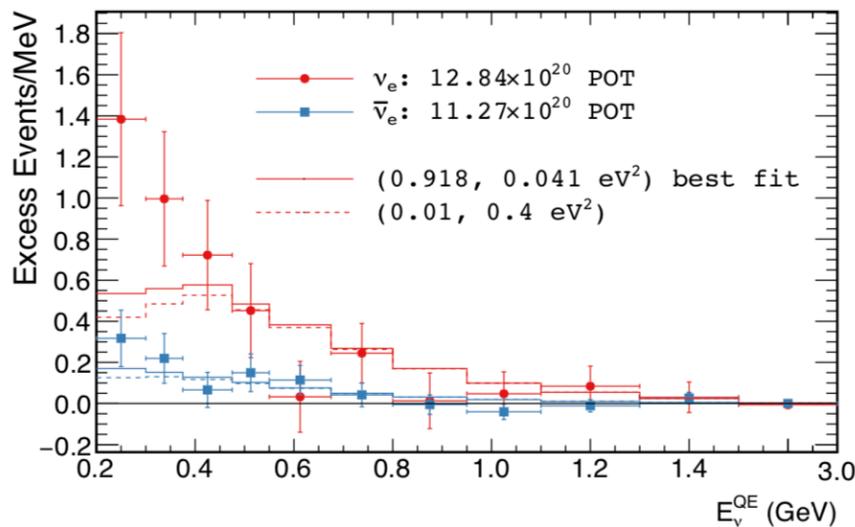
- LSND saw an excess of $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
 - Could be interpreted as oscillations at a mass-splitting scale of approximately 1 eV^2
- However, KARMEN2 saw results consistent with expectation.



- The MiniBooNE experiment was devised to investigate these differing results...
 - Looked at $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ and $\nu_\mu \rightarrow \nu_e$

Some Anomalies - 1

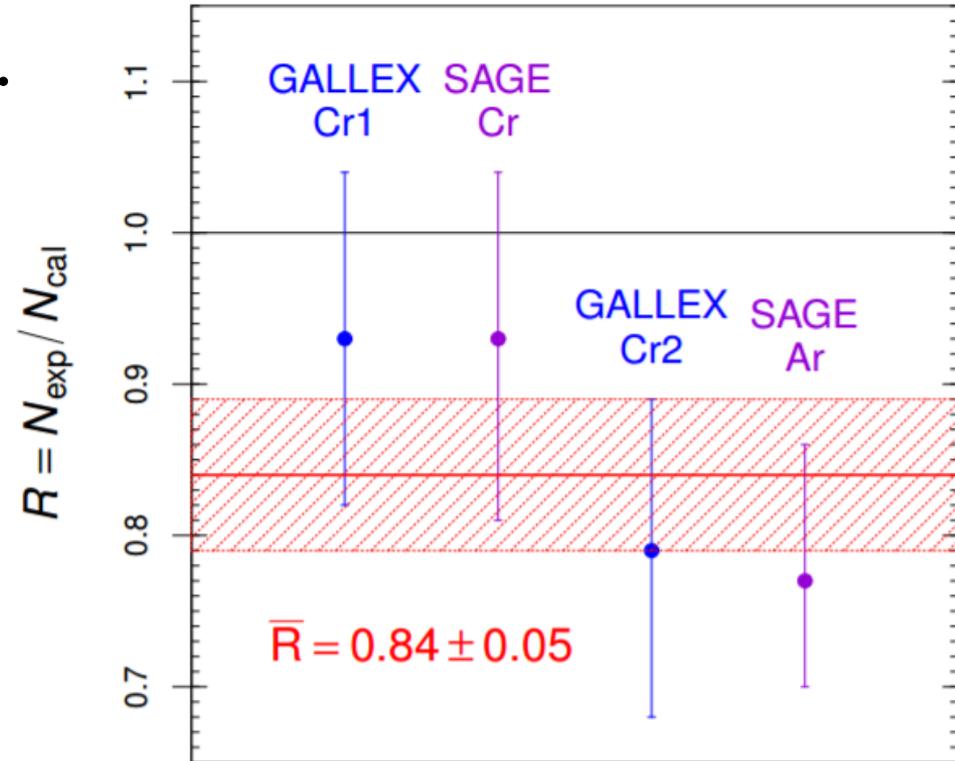
- MiniBooNE saw excess appearance in both neutrino and anti neutrino channels.
- Not identical to LSND, but allowed similar regions of phase-space.



A. Aguilar-Arevalo et al. Phys. Rev. Lett. 121 (2018), p. 221801.

Some Anomalies - 2

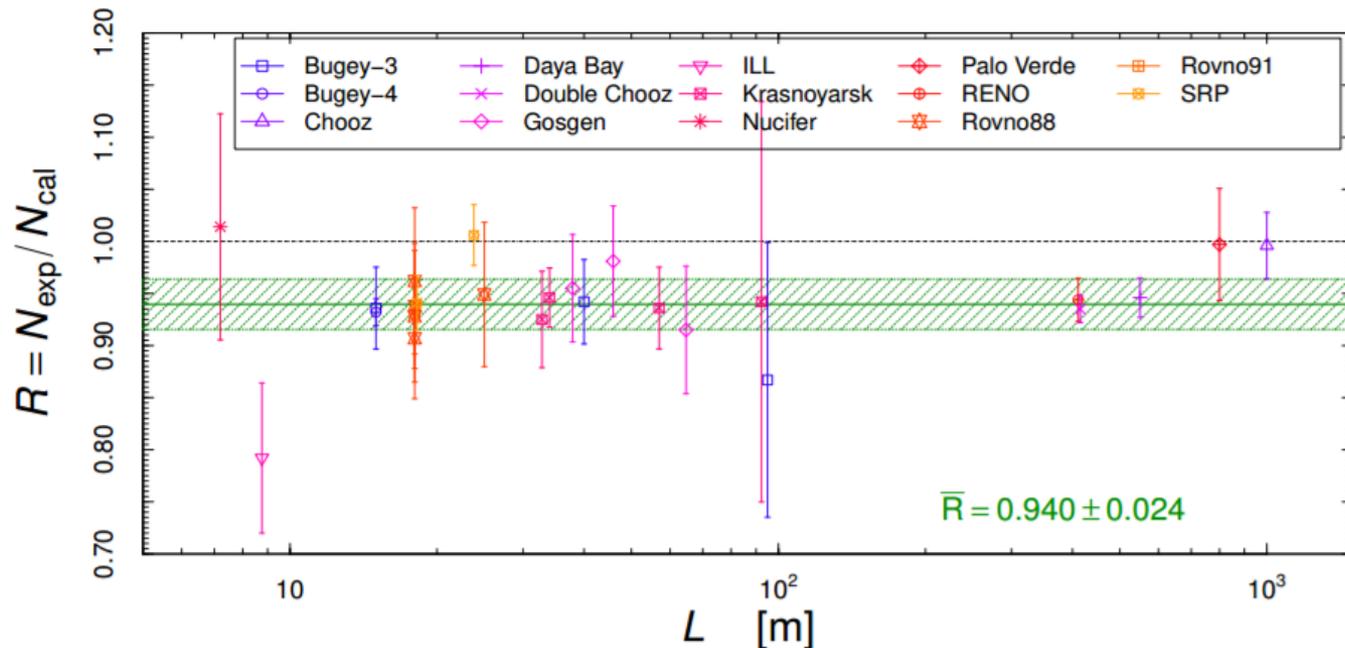
- GALLEX and SAGE were two solar neutrino experiments.
- Calibrated using radioactive sources.
- Measured rates from the calibration sources showed consistent deficits.
- Again, consistent with a 1 eV^2 mass-splitting.



Gariazzo et al. J.Phys. G43 (2016) 033001
DOI:10.1088/0954-3899/43/3/033001

Some Anomalies - 3

- The majority of reactor neutrino experiments have seen a deficit of $\bar{\nu}_e$.

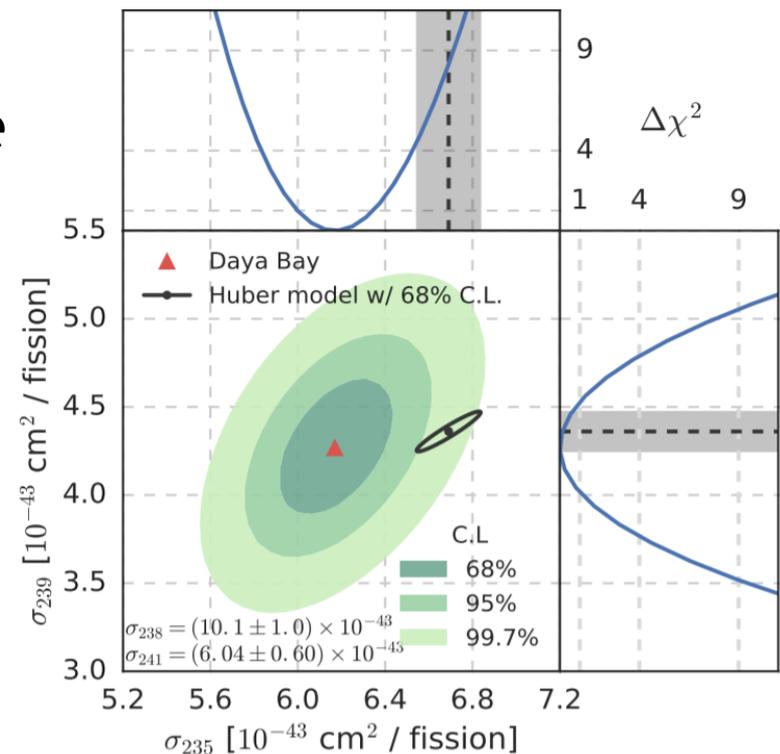


Gariazzo et al. (2017). arXiv: 1703.00860 [hep-ex]

- Again, consistent with a 1 eV^2 mass-splitting, but...

Some Anomalies - 3

- Daya Bay released results from studying their flux as a function of reactor fuel cycles to extract information on the uranium and plutonium components.
- Flux deficit appears to only come from the uranium flux.
- The sterile neutrino hypothesis for the reactor anomaly is: “incompatible with Daya Bay’s observation at 2.6σ ”.



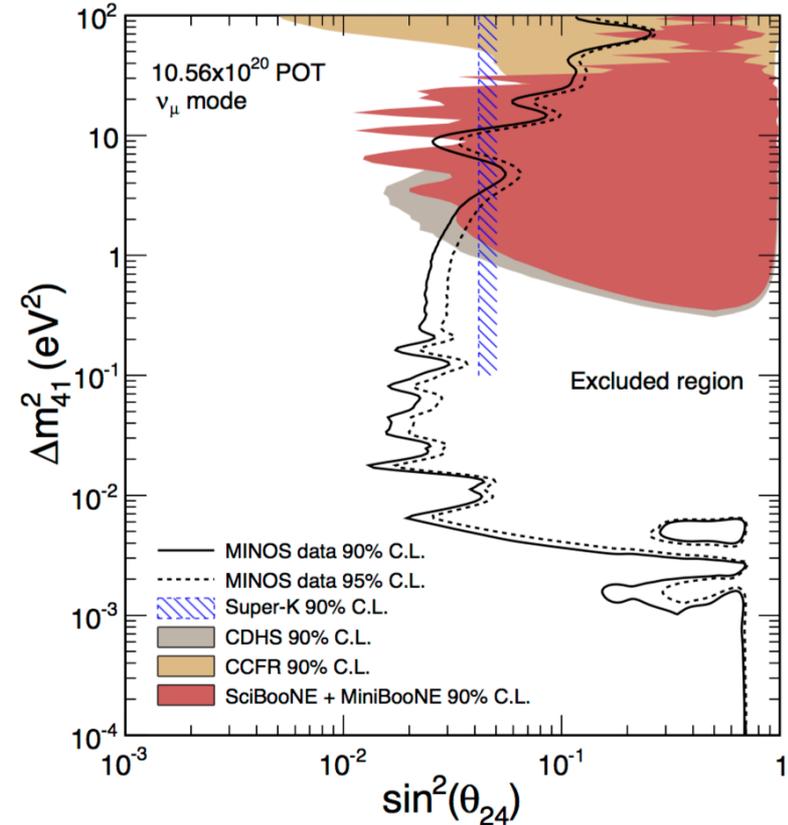
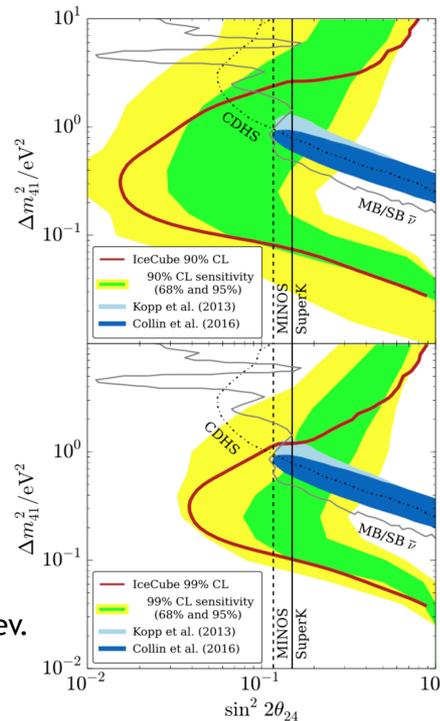
An et al. (2017). arXiv: 1704.01082 [hep-ex]

Null Results

- A number of muon neutrino disappearance experiments see no evidence of a sterile neutrino.

- MiniBooNE + SciBooNE
- MINOS
- IceCube
- CDHS
- CCFR
- Super-K
- ...

M. G. Aartsen et al. Phys. Rev. Lett. 117, 071801 (2016)

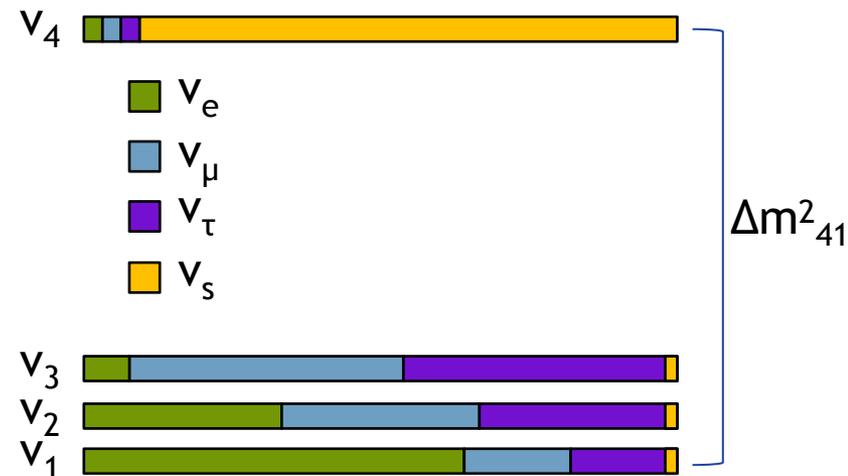


P. Adamson et al., Phys. Rev. Lett. 117, 151803 (2016).

Four Flavour Formalism

- Most common extension to include a 4th neutrino is the 3+1 model.
- PMNS matrix becomes 4 x 4
 - Three new mixing angles: θ_{14} , θ_{24} and θ_{34}
 - Two new CP phases: δ_{14} and δ_{24}
- Three new mass-splittings, but only one is independent.
 - Δm^2_{41}

$$\begin{pmatrix} \boxed{U_{PMNS}} & U_{e4} \\ & U_{\mu4} \\ & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$





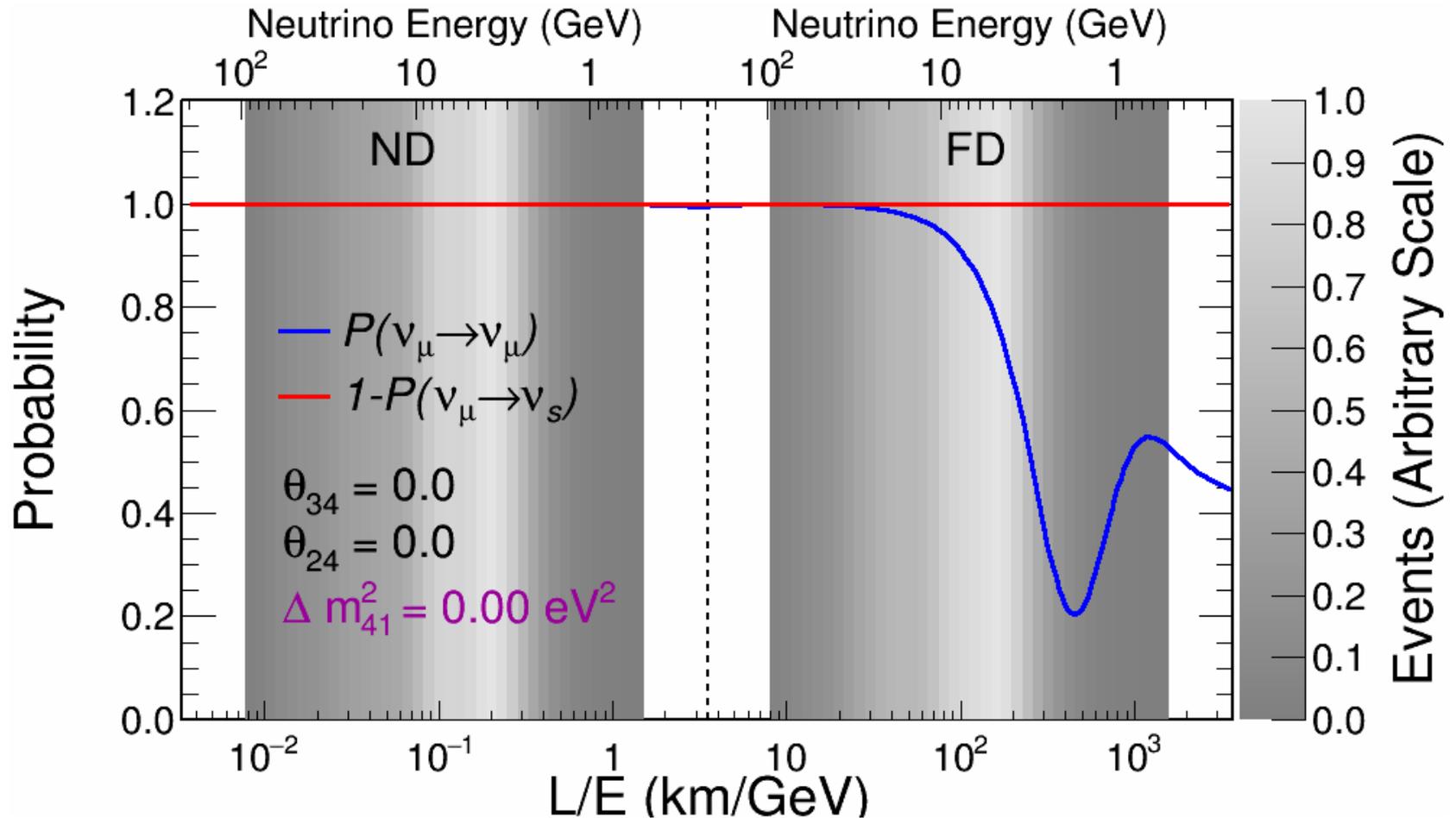
MINOS+ Four Flavour Oscillation Analysis

Sterile Oscillations in MINOS+

- MINOS+ is sensitive to three of the sterile oscillation parameters.
- Muon neutrino disappearance: θ_{24} and Δm^2_{41}
 - Measured with muon neutrino charged-current events.
- Active neutrino disappearance: θ_{24} , θ_{34} and Δm^2_{41}
 - Measured using neutral-current interactions.
 - Sensitivity reduced compared to CC due to worse energy resolution and lower cross-section.
- Oscillations can cause effects in both detectors depending on the value of Δm^2_{41}

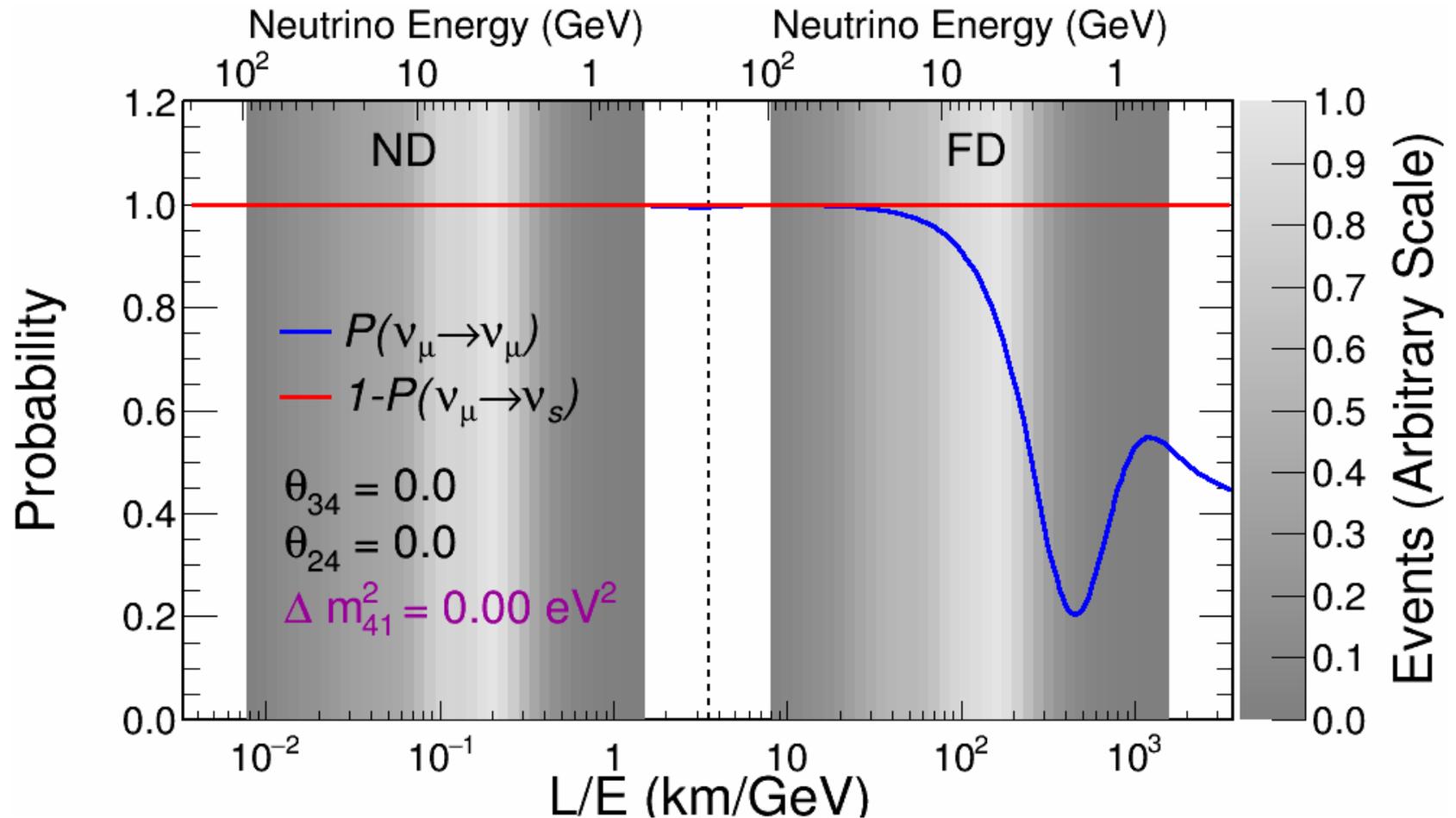
Sterile Oscillations in MINOS+

$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \sin^2 2\theta_{23} \cos 2\theta_{24} \sin^2 \Delta_{31} - \sin^2 2\theta_{24} \sin^2 \Delta_{41}$$



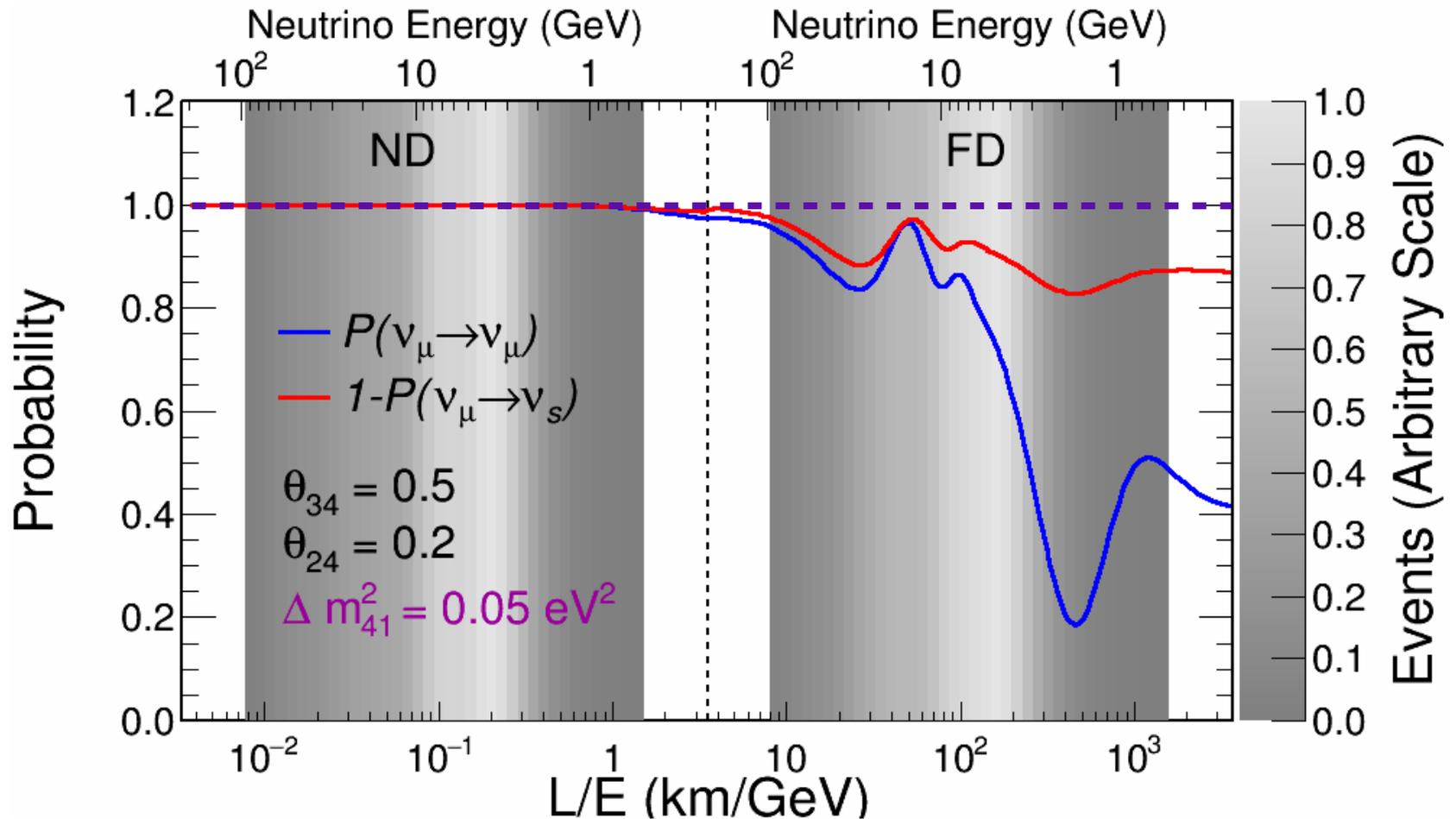
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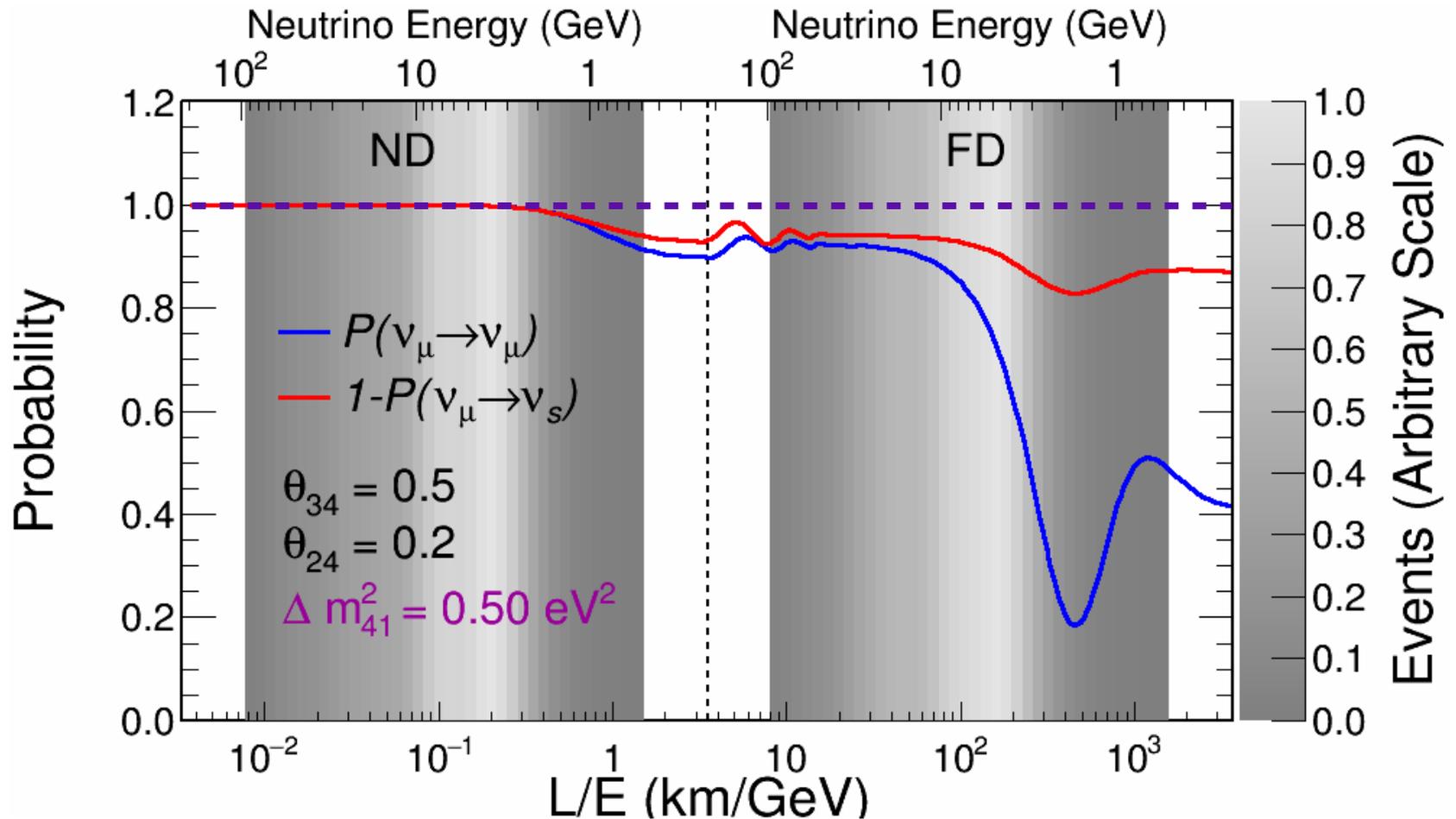
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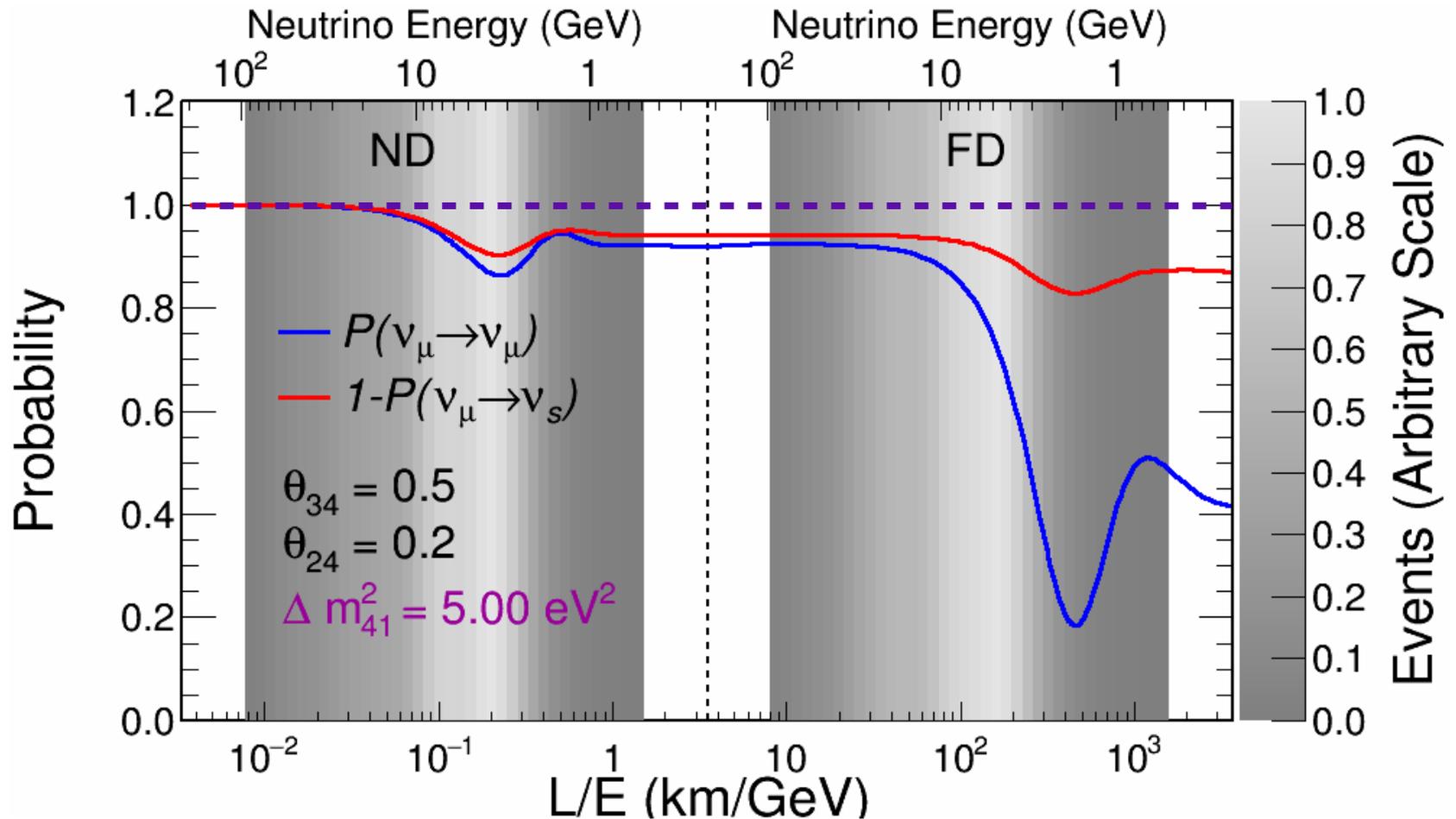
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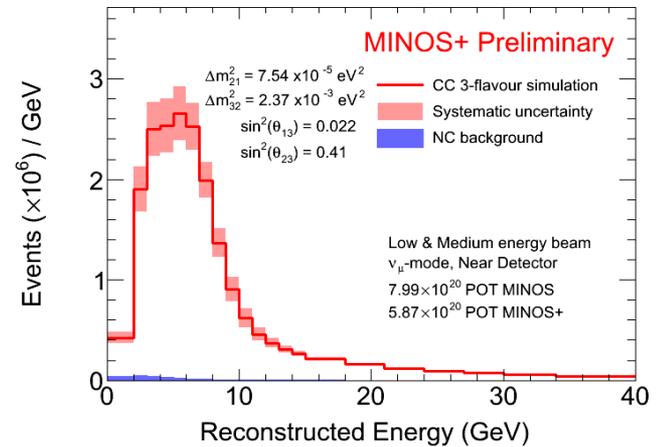
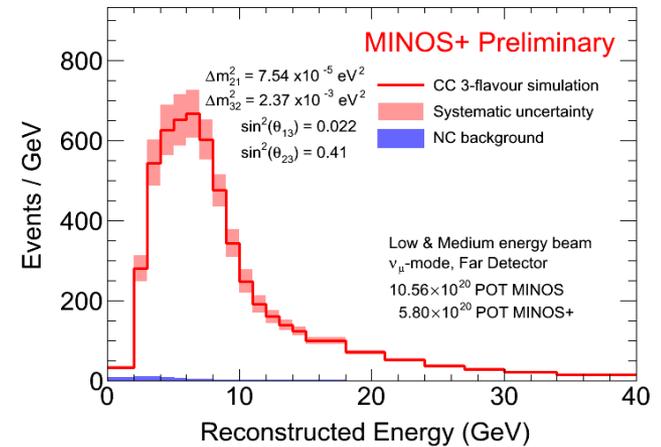
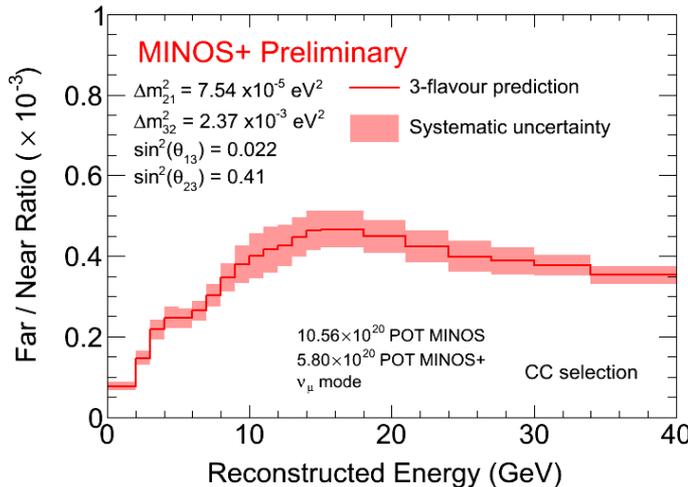
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Analysis Method

- The previous MINOS sterile neutrino analysis used the ratio of the Far and Near spectra.
 - Can't use the ND to tune the MC like in our three-flavour analysis.
 - Many systematics cancel in the ratio.



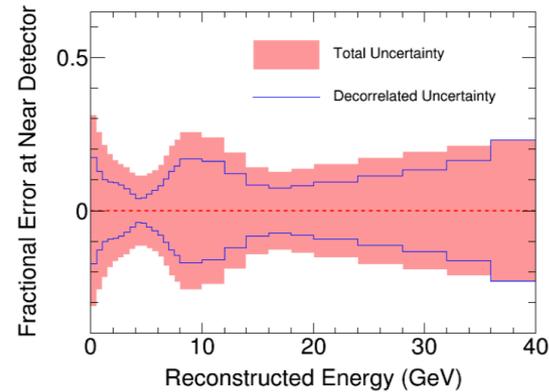
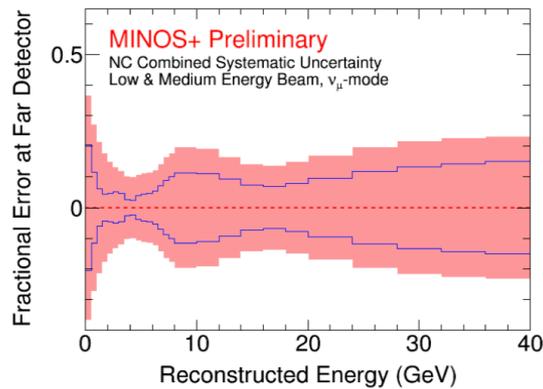
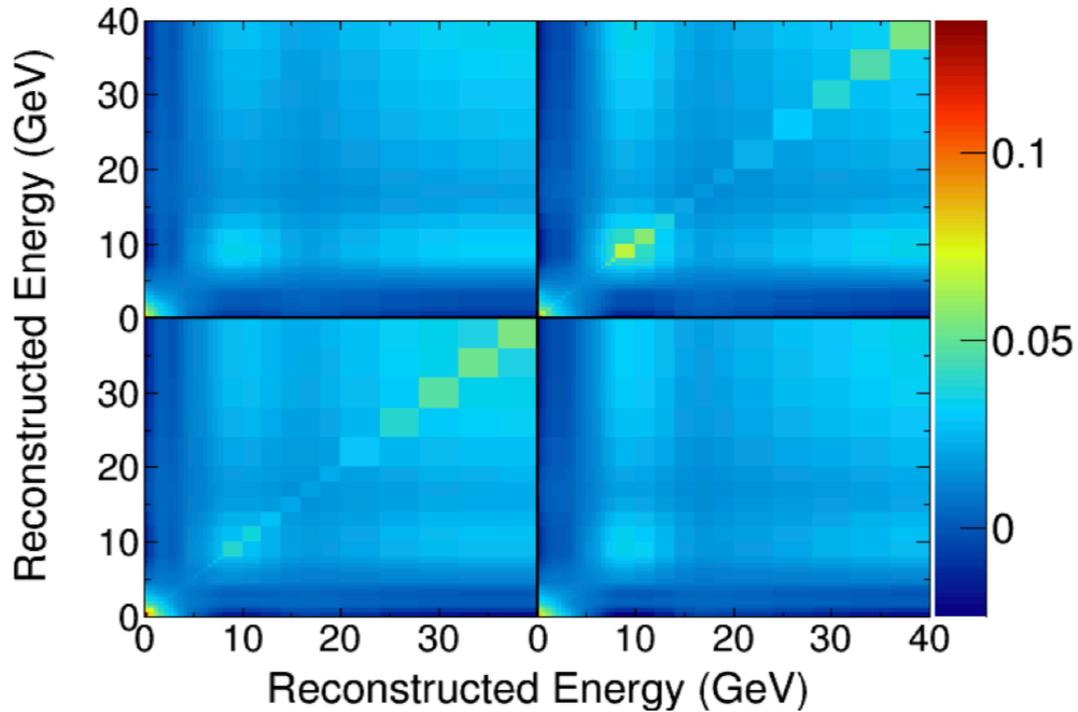
- Uncertainty in the ratio was dominated by FD statistics.

The Two Detector Fit

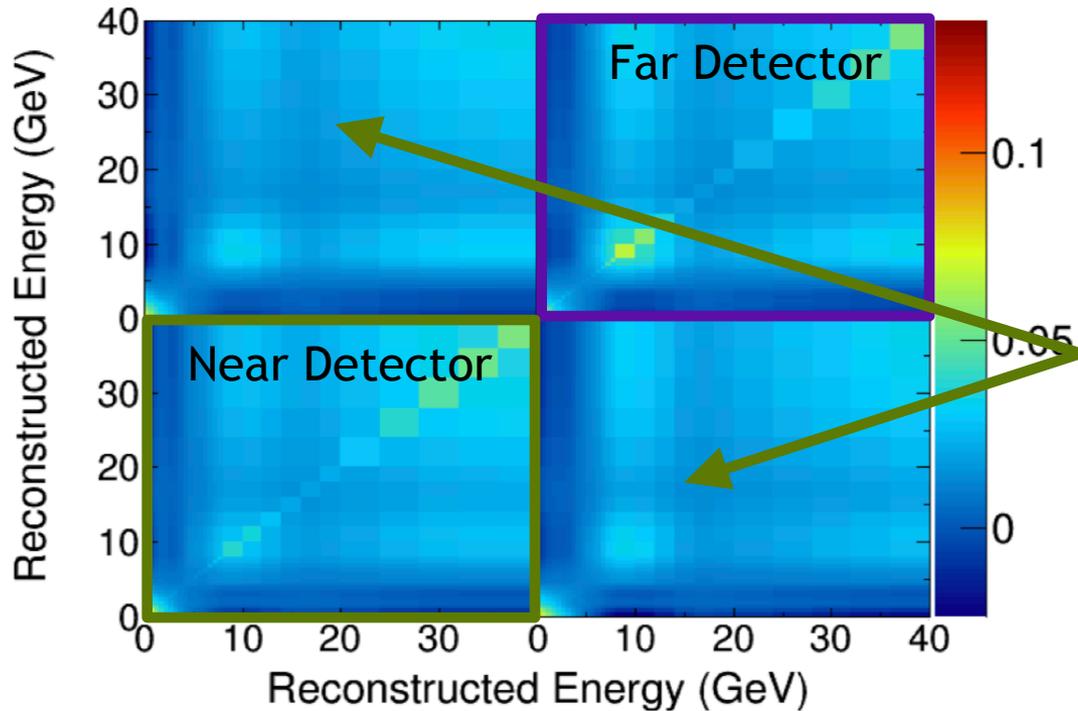
- We have now moved to a simultaneous two detector fit.
 - Use the a-priori flux prediction from MINERvA [1]
- We use a single covariance matrix that encapsulates the correlations between the systematic uncertainties.
 - This still enables us to have some cancellation of the systematic uncertainties without using the Far-over-Near ratio.
- Consider a total of 44 systematic uncertainties across the different event selections.

[1] L. Aliagia, et al, Phys. Rev. D 94, 092005, 2016

Systematic Uncertainties: NC



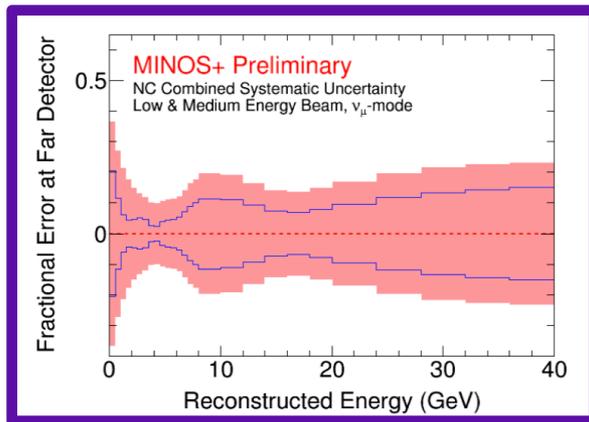
Systematic Uncertainties: NC



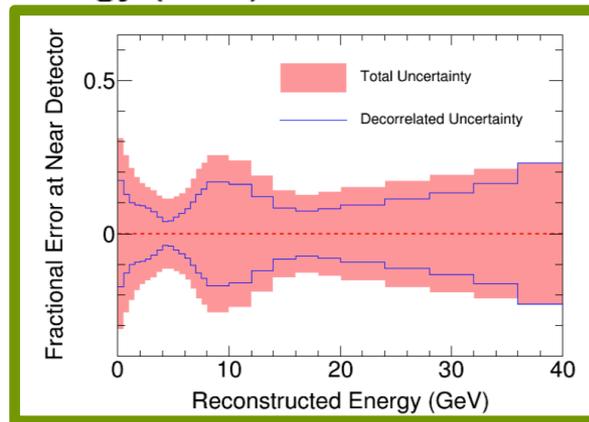
Diagonal components form the bands below

Correlations between the two detectors

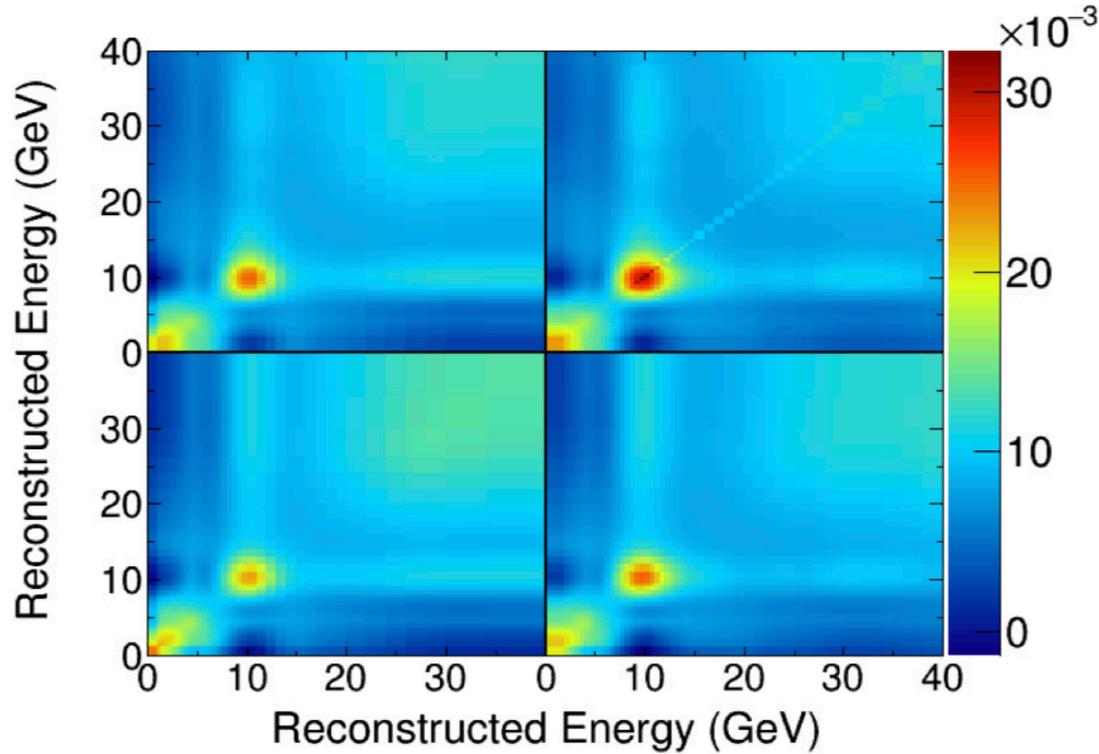
Far Detector



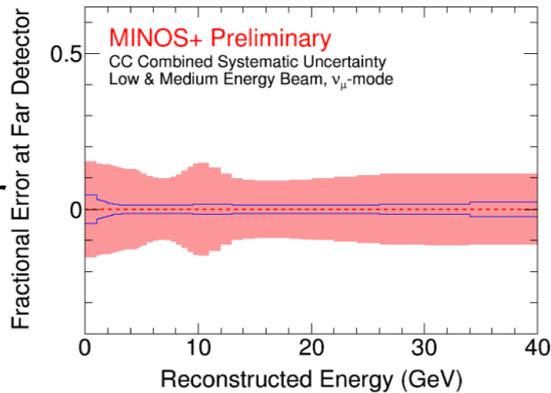
Near Detector



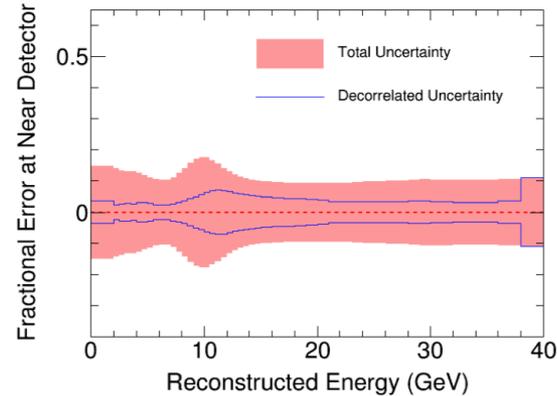
Systematic Uncertainties: CC



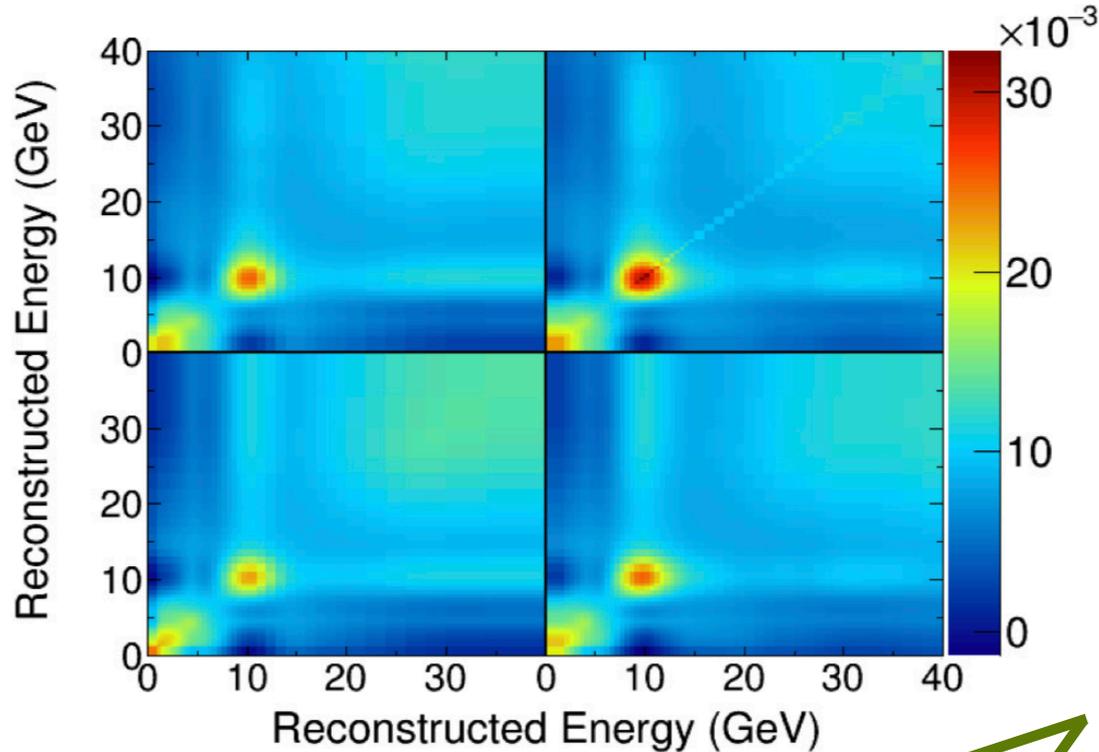
Far Detector



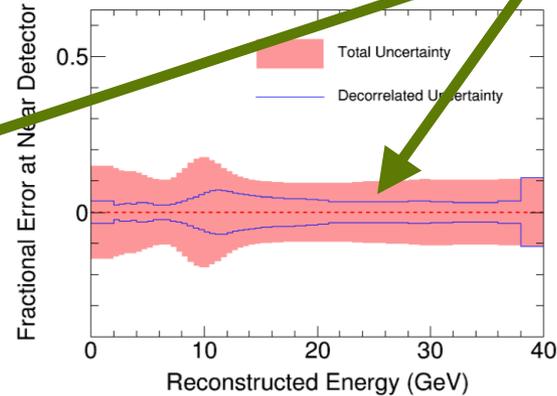
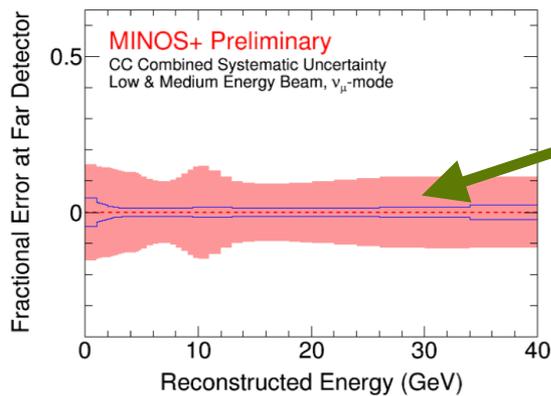
Near Detector



Systematic Uncertainties: CC



Strong correlations between the detectors gives strong cancellation of systematic uncertainties



The Fit Procedure

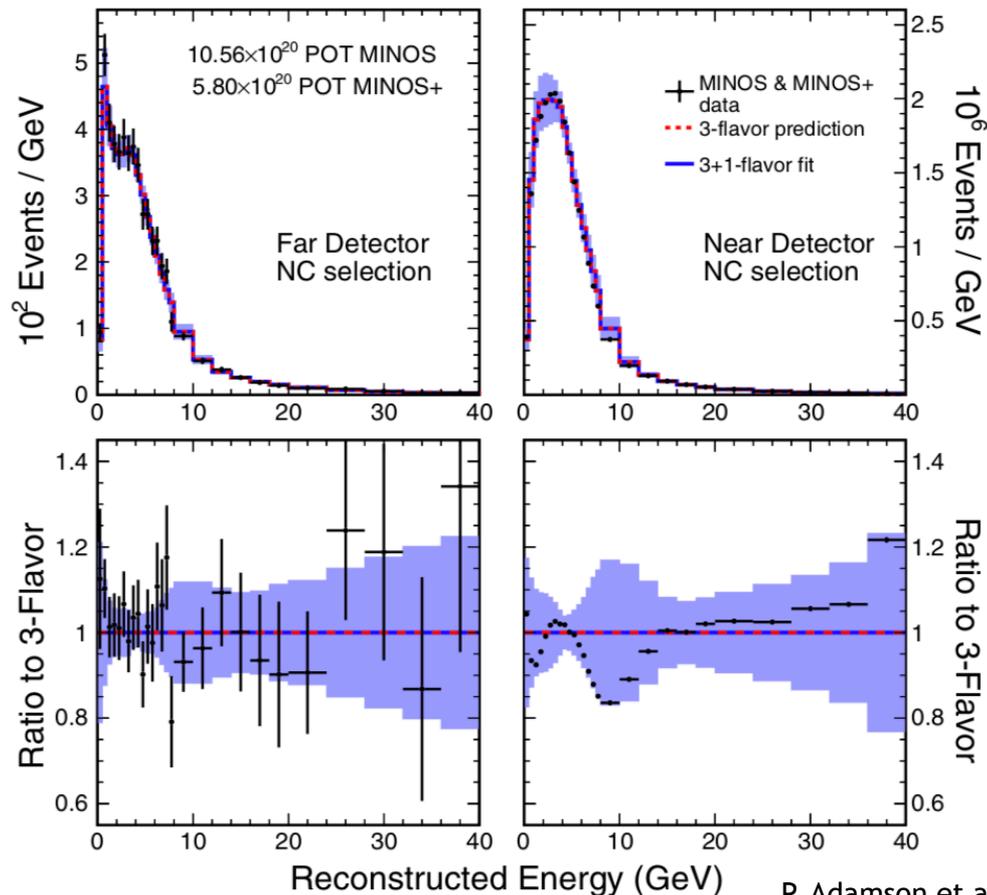
- Perform a fit to minimise the following for both the CC and NC samples

$$\chi^2 = \sum_{i=1}^N \sum_{j=1}^N (x_i - \mu_i) [V^{-1}]_{ij} (x_j - \mu_j) + \text{penalty}$$

- We fit for Δm^2_{41} , Δm^2_{32} , θ_{23} , θ_{24} and θ_{34}
- Global best fit values are used for Δm^2_{21} , θ_{12} and θ_{13}
- The other parameters have a negligible effect on the analysis and are set to zero: θ_{14} , δ_{13} , δ_{14} and δ_{24}
- Penalty term prevents from Δm^2_{32} becoming degenerate with Δm^2_{41}

Event Spectra: NC Selection

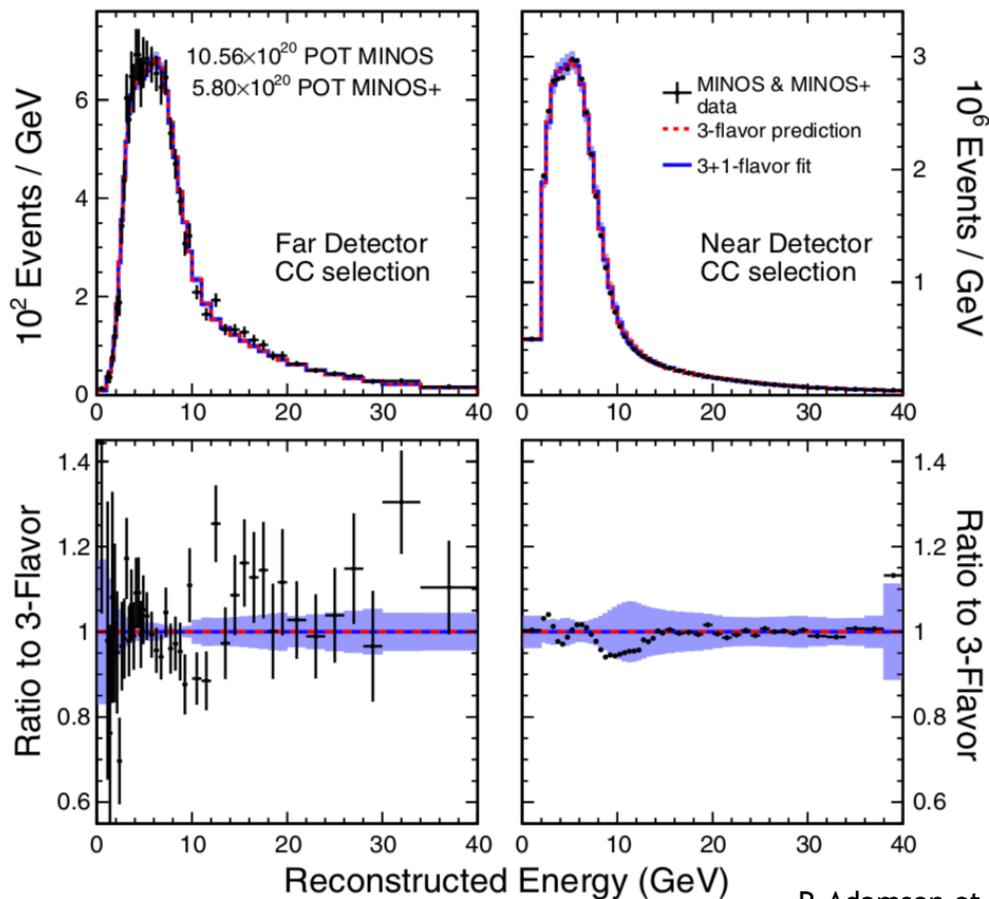
- Spectra shown after correcting the prediction after decorrelation of the covariance matrix



P. Adamson et al., Phys. Rev. Lett. 122, 091803 (2019).

Event Spectra: CC Selection

- Spectra shown after correcting the prediction after decorrelation of the covariance matrix



P. Adamson et al., Phys. Rev. Lett. 122, 091803 (2019).

The Fit Result

- Best fit point:

$$\Delta m_{41}^2 = 2.33 \times 10^{-3} \text{ eV}^2$$

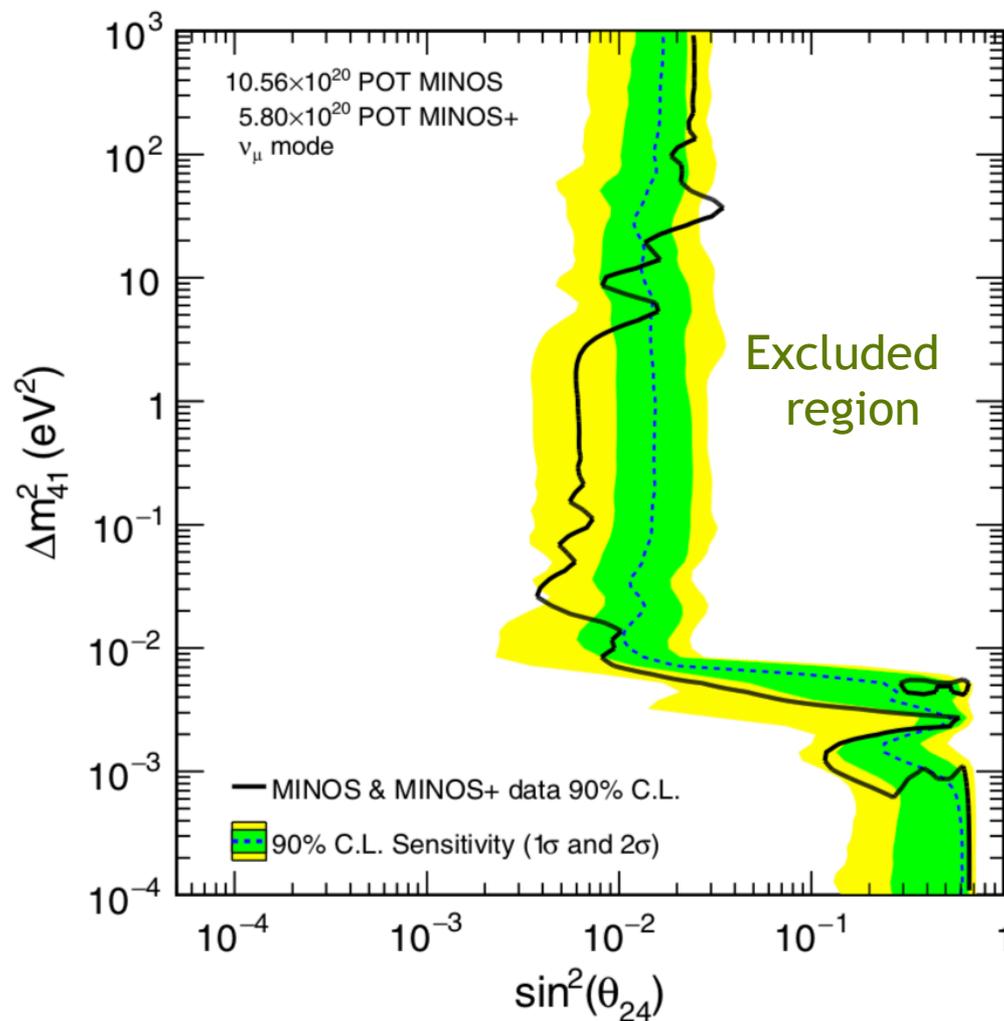
$$\sin^2 \theta_{24} = 1.1 \times 10^{-4}$$

$$\theta_{34} < 8.4 \times 10^{-3}$$

$$\sin^2 2\theta_{23} = 0.92$$

$$\chi^2_{\min} / \text{dof} = 99.3 / 140$$

$$\chi^2(4\nu) - \chi^2(3\nu) = 0.01$$



P. Adamson et al., Phys. Rev. Lett. 122, 091803 (2019).

Comparison with Other Results

- Best fit point:

$$\Delta m_{41}^2 = 2.33 \times 10^{-3} \text{ eV}^2$$

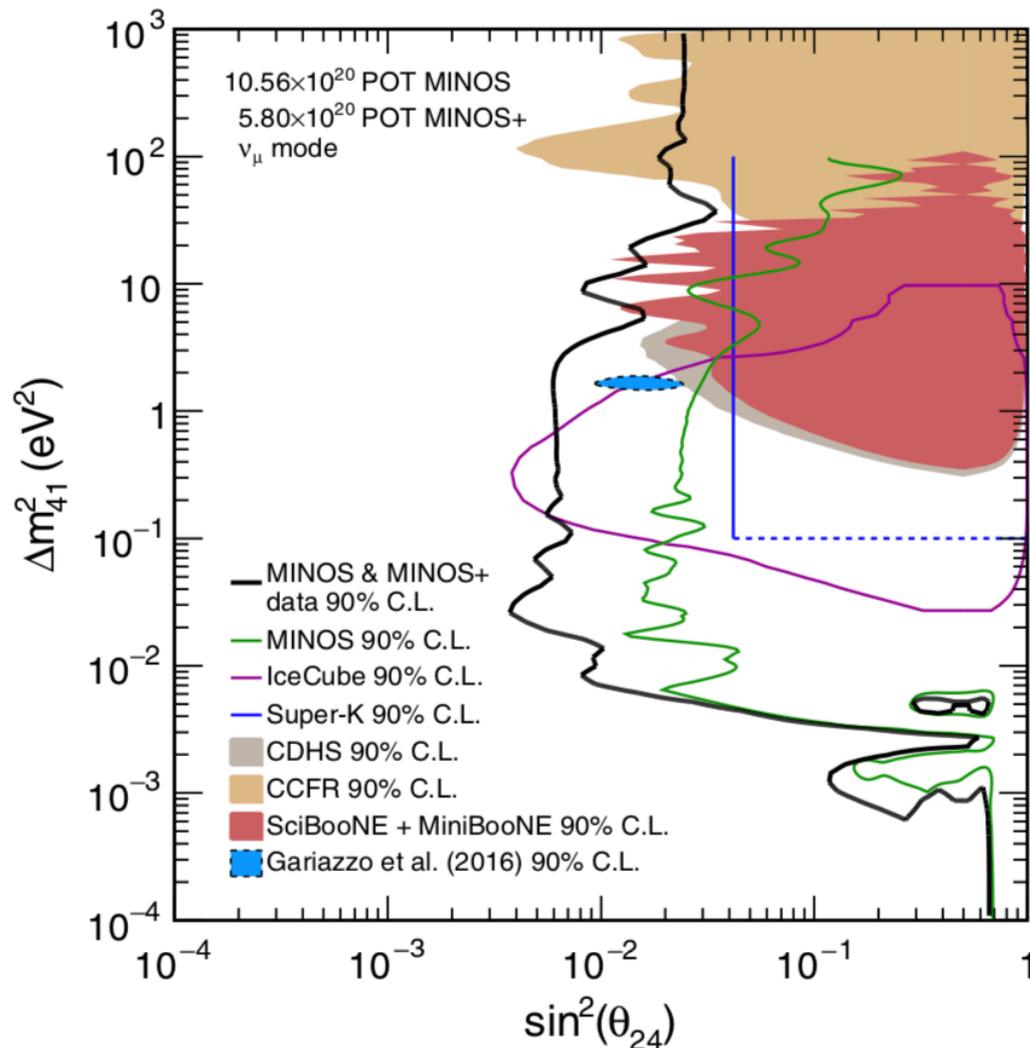
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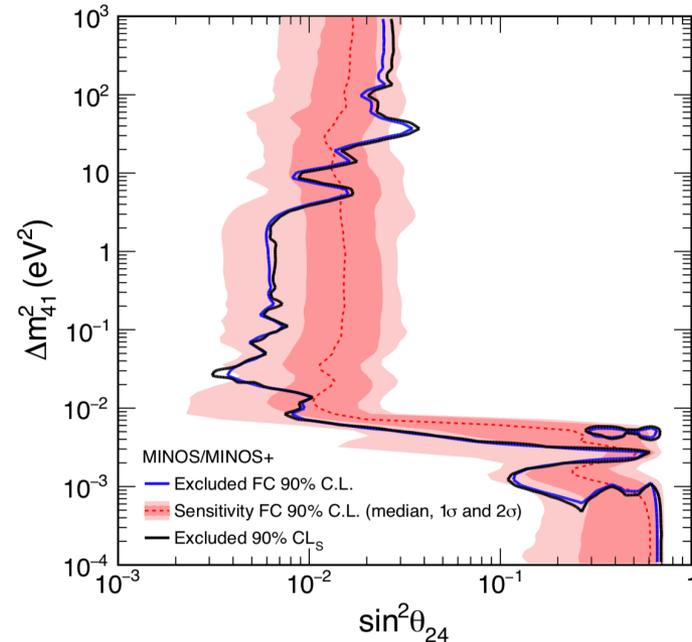
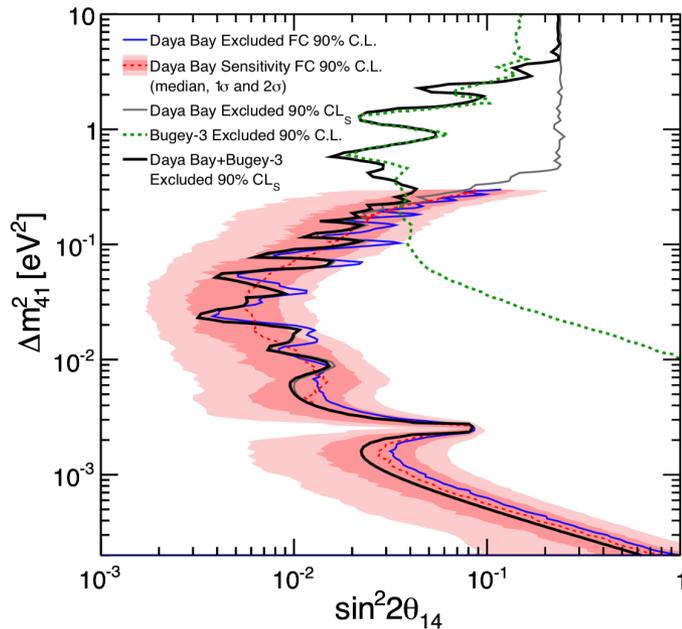
$$\chi^2(4\nu) - \chi^2(3\nu) = 0.01$$



P. Adamson et al., Phys. Rev. Lett. 122, 091803 (2019).

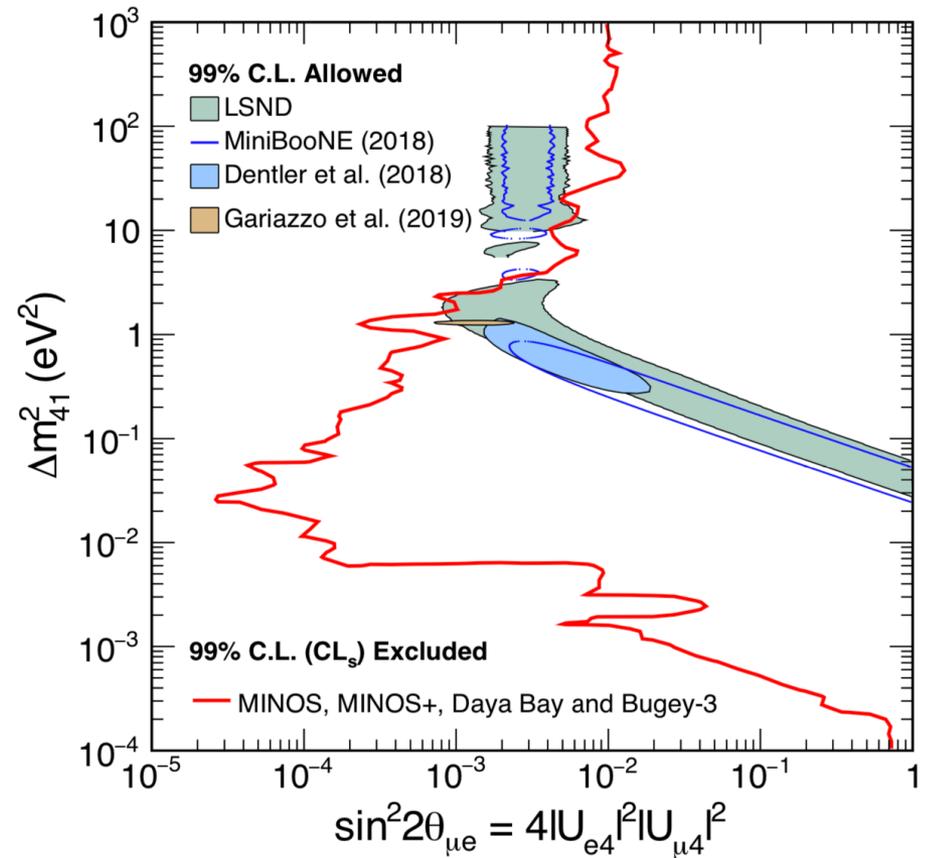
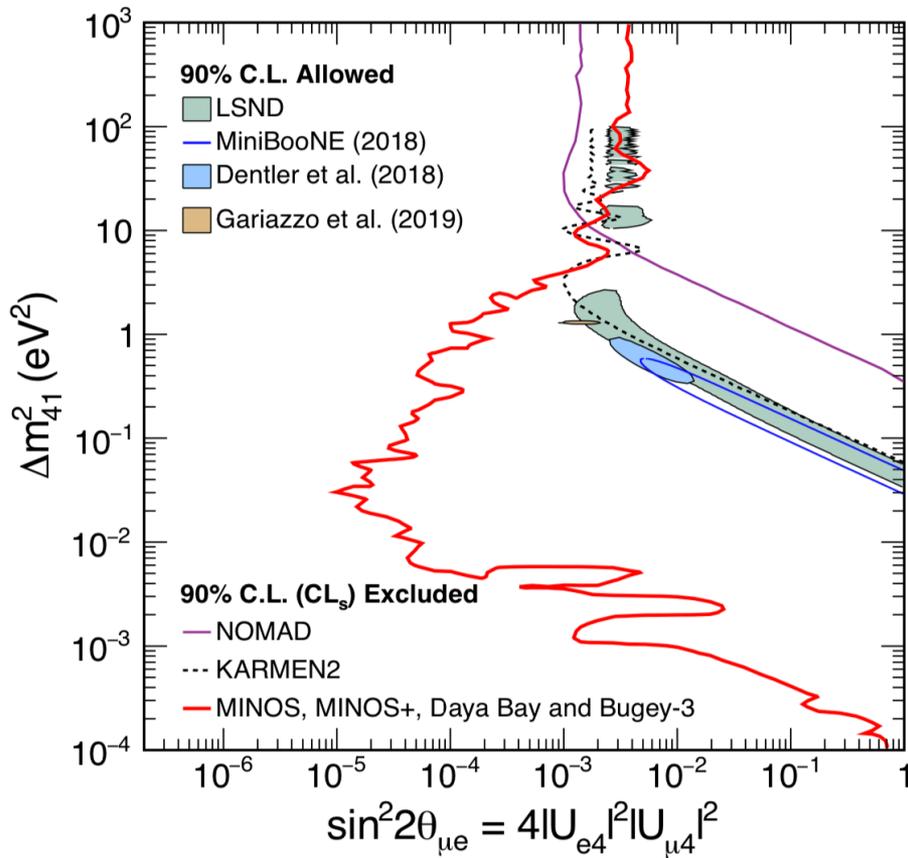
Combination with Daya Bay

- MINOS+ sensitive to $|U_{\mu 4}|^2 = \sin^2 \theta_{24} \cos^2 \theta_{14}$
- Daya Bay ν_e disappearance sensitive to $|U_{e 4}|^2 = \sin^2 \theta_{14}$



- Combine to probe the same parameter-space as LSND and MiniBooNE: $4|U_{e 4}|^2|U_{\mu 4}|^2 = \sin^2 2\theta_{14} \sin^2 \theta_{24} \equiv \sin^2 2\theta_{\mu e}$

Combination with Daya Bay

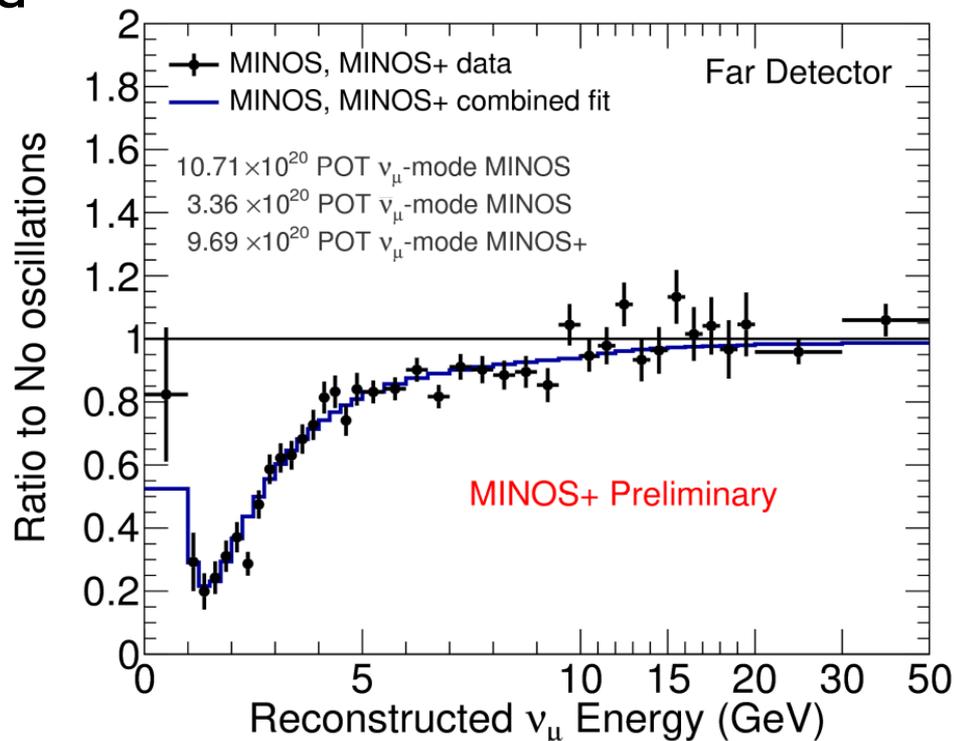


P. Adamson, et al., <https://arxiv.org/abs/2002.00301>

Summary

- MINOS/MINOS+ has produced its final three flavour muon neutrino disappearance result

- Very high statistics covering the entire oscillation dip
- Measured Δm^2_{32} to 3.5%



- The four-flavour analysis gives a leading exclusion on the sterile neutrino hypothesis over many orders of magnitude in Δm^2_{41}



Thank You!
Any Questions?