

Search for Hidden Particles



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ATLAS SUSY Searches* - 95% CL Lower Limits				ATLAS Preliminary	
Model	\sqrt{s} [TeV]	\mathcal{L}_{int} [fb ⁻¹]	σ [fb]	Mass limit [TeV]	Reference
Neutralino	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Chargino	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Stau	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Top squark	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Bottom squark	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Glueball	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Gravitino	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]
Other	6.0-7.0	36.1-100.0	100-1000	1.0-1.5	[1]

*This table contains the available mass limits for various models of supersymmetry. All limits quoted are model-dependent and for illustrative purposes only. For more information, see the ATLAS SUSY Searches page.



- **Higgs found!** SM complete and consistent up to Planck scale. **But...**
 - matter-antimatter asymmetry
 - neutrino masses/mixing
 - dark matter
 - flavour anomalies... **New physics?**
- **NO** smoking gun in direct searches up to ~ 5 TeV...

What is the energy scale of new physics?



→ Neutrino masses and oscillations:

Right Handed see-saw neutrino masses from 1 eV to 10^{15} GeV

→ Dark matter:

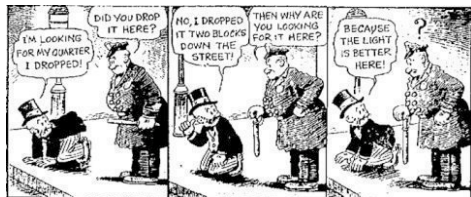
From 10^{-22} eV (super-light scalar) to $\geq 10^{20}$ GeV (wimpzilla, Q-ball)

→ Baryogenesis:

Mass of new particle from 10 MeV to 10^{15} GeV

→ Higgs mass hierarchy:

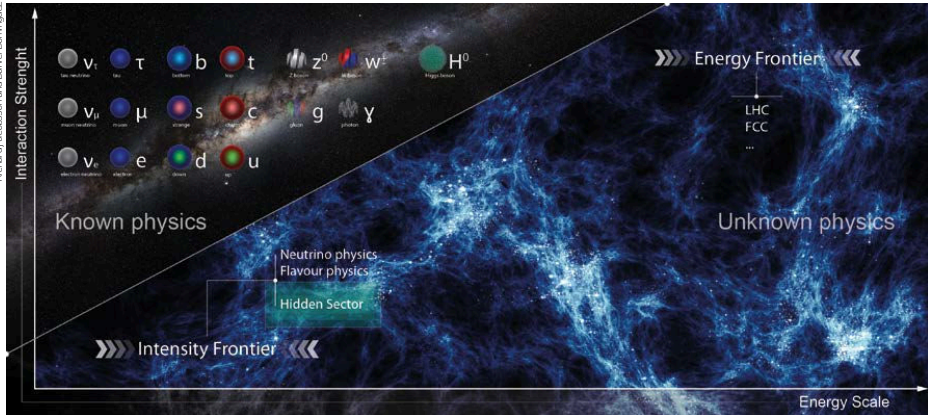
SUSY, GUT, composite Higgs, large extra dimensions theories require the presence of new particles above the Fermi scale.



Where is new physics? Experimental approach



Richard J. Coles and Daniel Dominguez



→ Unsolved **problems** \implies new particles

→ Why didn't we detect them? **Too heavy** or **too weakly interacting**



$$\mathcal{L}_{\text{world}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{portal}} + \mathcal{L}_{\text{HS}}$$

- **Hidden Sector (HS)** naturally accomodates Dark Matter
 - it may have a **rich** structure
- Interaction with visible sector (SM) proceeds through **mediators**
- HS processes very **strongly suppressed** relative to SM
 - production BRs $\sim 10^{-10}$
 - very weak interaction with matter
 - very long-lived objects!
- Can search HS through **decays to visible particles**

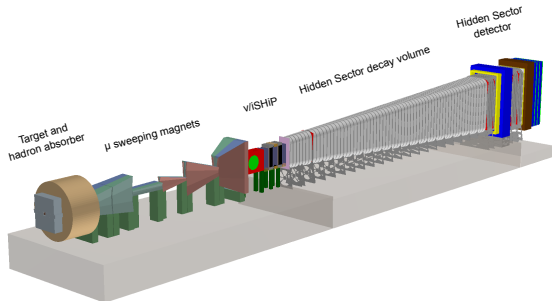
SHiP: Search for Hidden Particles



SHiP is a new proposed intensity-frontier experiment aiming to search for HNLs and other neutral hidden particles with mass up to $\mathcal{O}(10)$ GeV and **extremely weak** couplings.

Aims to be a **zero background** experiment!

Facility also ideally suited for studying ν_τ and $\bar{\nu}_\tau$ properties and testing lepton flavour universality by comparing interactions of μ and τ neutrinos.





- **The SHiP experiment**
 - Detector system
 - Background strategies

- **Physics with ν_τ**

- **The search for Heavy Neutral Leptons**
 - Evaluating SHiP sensitivity

- **Probing the Hidden Sector**
 - Vector portal
 - Scalar portal
 - Axion-like particles

- **Conclusions**



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Experimental requirements



→ Hidden particles production in **charm** (and beauty) **charm**

- Intense proton beam from the SPS (400 GeV)
- Very dense target of $10 \times \lambda_{\text{int}}$
 - abundant production of heavy flavours
 - reduce neutrino production from π and K decays

→ decay of hidden particles:



- **large decay volume** followed by **spectrometer, calorimeter, PID**
- **shielding** from SM particles: hadron absorber + VETO detectors (*)

→ τ neutrinos:

- $N_\tau = 4N_p (\sigma_{c\bar{c}}/\sigma_{pN}) f_{D_s} \times \text{Br}(D_s \rightarrow \tau) \simeq 6 \times 10^{15}$
- distinguish $\nu_\tau / \bar{\nu}_\tau$: **magnetized emulsion target** + **high-res tracker**

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TEASER: $\tau \rightarrow 3\mu$ sensitivity $\sim 10^{-10} / \sqrt{N_{\text{targets}}}$

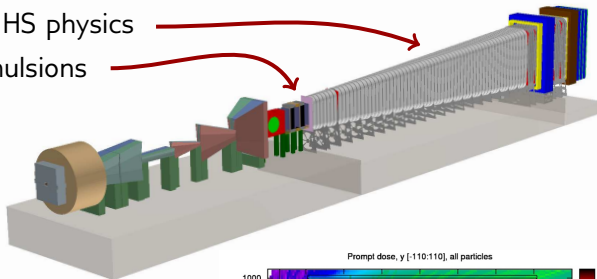
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...and the muons (*)?

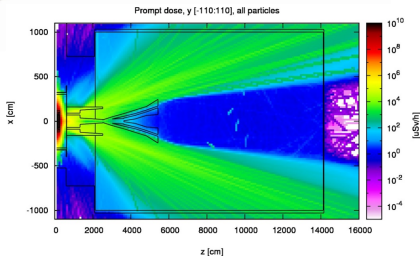


Residual μ flux after the hadron absorber is **dangerous**:

- background for HS physics
- ageing of ν_τ emulsions



- **active muon shield** based on sweeping magnets
- “conical”-shaped vessel



History

2013:

- submission of the EOI (October, 16 authors) → *arXiv:1310.1762*

2014:

- SPSC discusses EOI (January)
- 1st workshop (June, 100 participants)

2015:

- submission of TP (April, 233 authors)
→ *arXiv:1504.04956*
- submission of PP (April, 85 authors)
→ *arXiv:1504.04855*
- discussion with SPSC referees

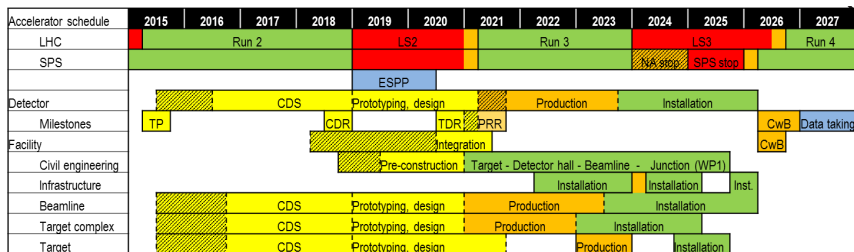
2016:

- endorsement by the SPSC (February)

2014–today:

- 12 collaboration meetings, workshops...



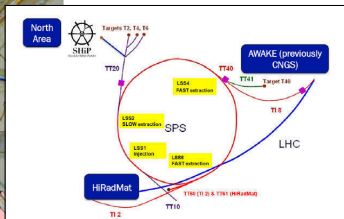
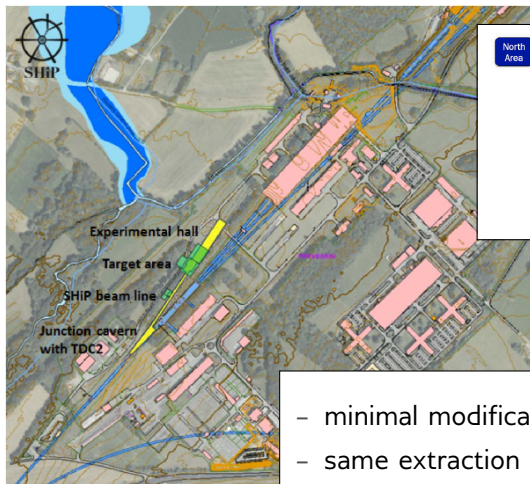


- “Comprehensive Design Study” + Report (2018) to input to the European Strategy consultation in 2019
- **New phase of optimisation**
 - Re-optimize detector design
 - Assess and improve sensitivity to various hidden sector models

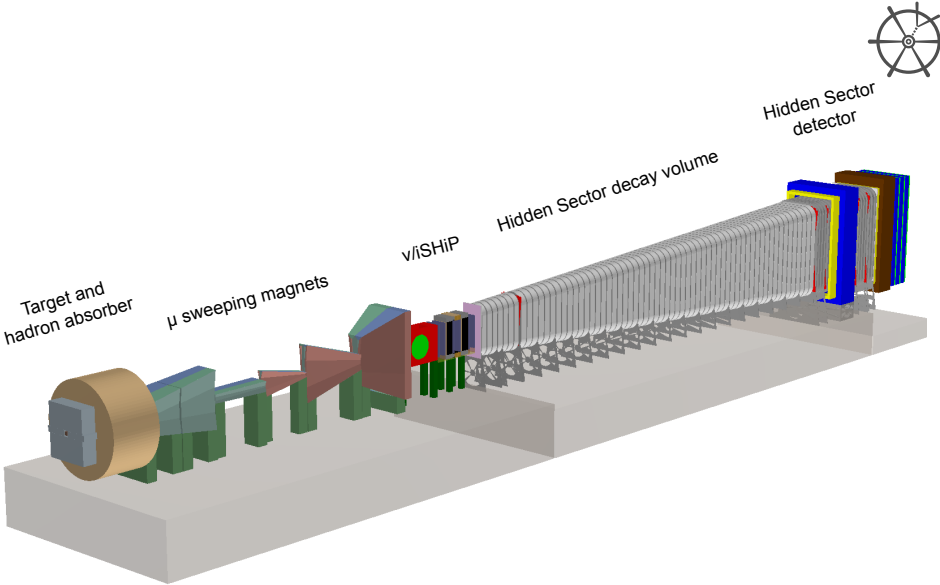


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 - Vector portal
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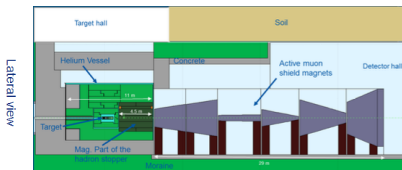
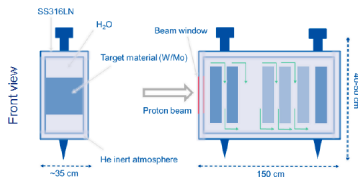
The facility at the SPS



- minimal modification to the SPS complex
- same extraction and transfer line as other NA facilities
- ~120 m long, ~20 m wide hall



Target and hadron absorber

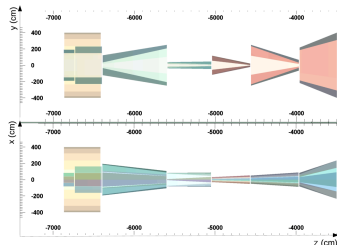
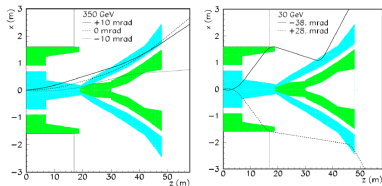


- Layers of titanium/zirconium/molybdenum
 - Heavy target stops hadrons before they decay. Only **muons** emerge
- Peak power **2.5 MW** (average 355 kW)
- **Challenges:**
 - Extremely radioactive environment
 - Many constraints on the integration, access, thermal stress, radioactivity...
- **Milestones:**
 - Reduced scale prototype by 2019 Q2
 - Full-scale prototype with cooling and power connections by 2020 Q3

Active muon shield



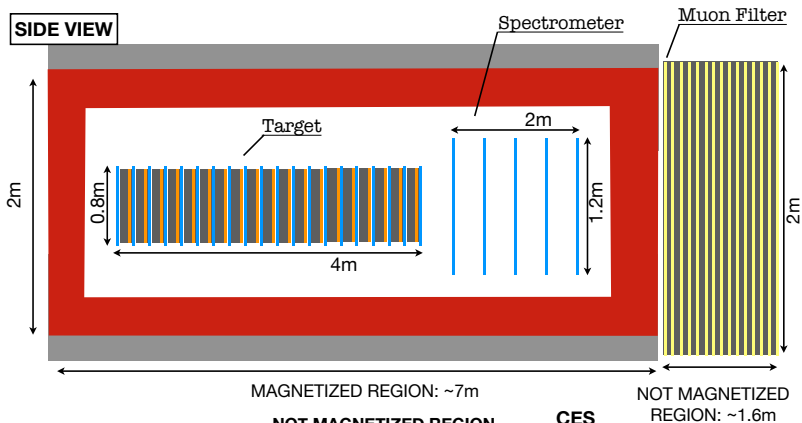
- Has to reduce μ flux by $\leq 10^{-6}$
- Critical component to optimise to maximize the exp. acceptance
 - global optimisation with **machine learning** ongoing
- μ spectrum measurement planned for 2018 at H4



JINST 12 P05011 2017

- **Challenges:**
 - narrow separation between field directions
 - aiming to 1.8 T to minimise length
 - size of the simulated μ sample to optimise with

The ν/i SHiP detector

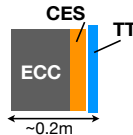


MAGNETIZED REGION

- ▶ Horizontal Magnetic field ≥ 1 T
- ▶ 19 brick walls
- ▶ 20 Target Tracker planes
- ▶ 5 Spectrometer planes

NOT MAGNETIZED REGION

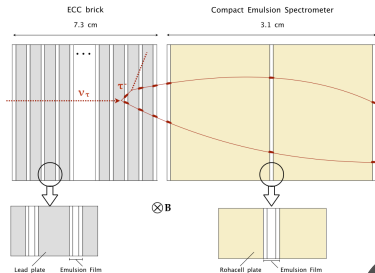
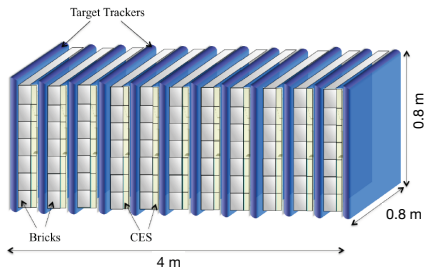
- ▶ 24 Iron 5 cm-thick planes
- ▶ 22 RPC 2 cm-thick planes



The ν/i SHiP detector: technologies



- ECC Emulsion Cloud Chambers composed of lead (passive target) and high-res nuclear emulsion films (tracking device)
 - ν interaction and τ decay vertex with μm resolution
- CES Compact Emulsion Spectrometer: air gaps and nuclear emulsions
 - measure the charge of the τ daughters, separate $\nu_\tau/\bar{\nu}_\tau$
- TT Target Trackers: scintillating fibres or MPGD
 - link emulsion tracks to spectrometer tracks
 - provide time stamp and calorimetry information



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- **Magnetized tracker** for tracking and muon charge: drift tubes
 - **Muon filter**, also BG tagger for HS detector: iron/RPCs
 - **New generation scanning system** based on GPU processing
→ scanning $190 \text{ cm}^2 \text{ h}^{-1}$
 - **ML techniques** for online/offline data processing

The Hidden Sector detector



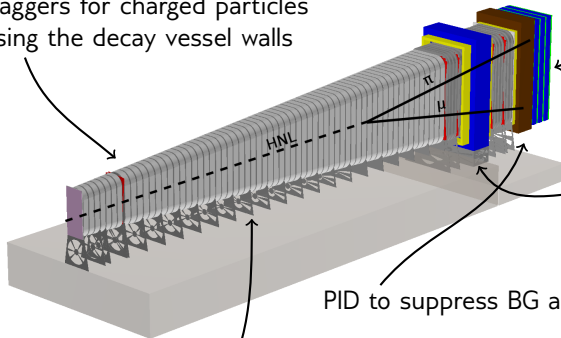
BG taggers for charged particles crossing the decay vessel walls

Timing to suppress combinatorial background

Tracking for vertexing and IP measurement

PID to suppress BG and distinguish final states:

Evacuated decay vessel to suppress ν interactions



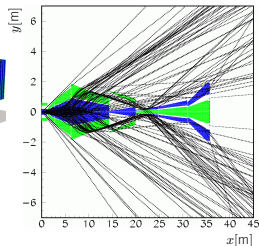
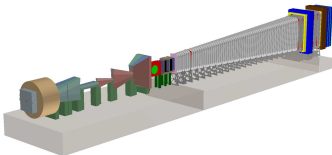
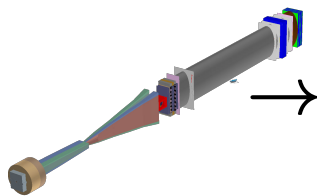
Hidden particle	Final states
HNL, neutralino	$l^\pm h^\mp$
Vector, scalar, axion, goldstino	$l^\pm l^\mp$
HNL, neutralino, axino	$l^\pm l^\mp \nu$
Axion, sgoldstino	$\gamma\gamma$
Sgoldstino	$\pi^0 \pi^0$

Hidden Sector detector optimisation



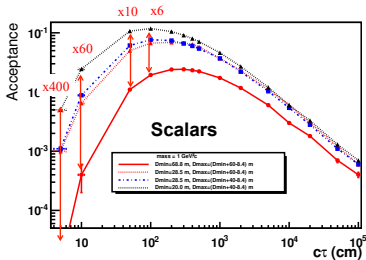
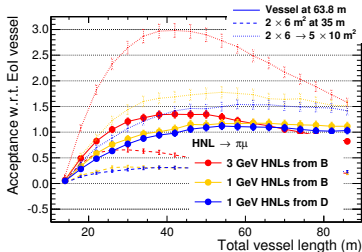
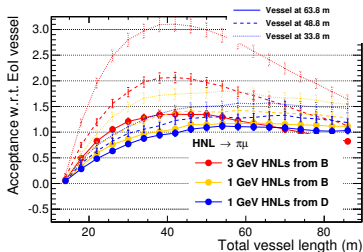
- **Optimisation:**

- **magnetize** hadron stopper → muon shield **shorter** and **lighter**
- **lower μ flux** entering shield → decay volume **closer** to target → **increased signal acceptance and less background**
- decay vessel has now a **pyramidal frustum** shape → **from ~ 150 m to ~ 107 m**



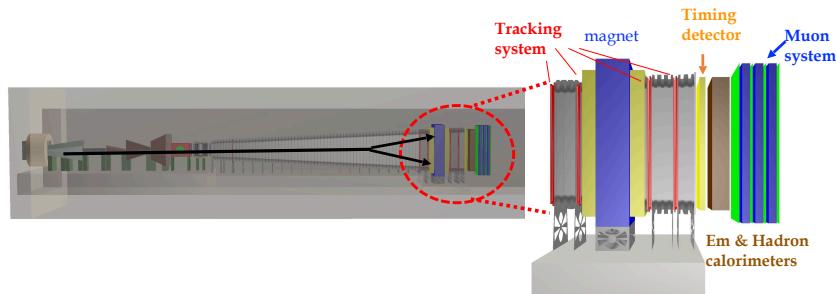
- $\leq 10^{-3}$ **mbar** to suppress ν interactions
- **Veto systems** surrounding the whole 60 m fiducial volume

Optimization of the decay volume



- studied cylindrical, conical solutions in vacuum or He
- might be able to release the vacuum requirement
- acceptance depends on the hidden particle's lifetime

Hidden Sector spectrometer



- Full reconstruction and PID of e^\pm , μ^\pm , π^\pm , $\gamma/\pi^0/\rho^\pm$ final states
- NA62-like straw tracker with 5 m long straw tubes
- Timing detector: scintillating bars with SiPM arrays, ≤ 100 ps
- Shashlik-type ECAL (HCAL) with 11504 (1512) channels
- Muon detector: 4 xy layers of scintillating bars with WLS fibres (23 km!) and SiPM readout



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- Evaluating SHiP sensitivity

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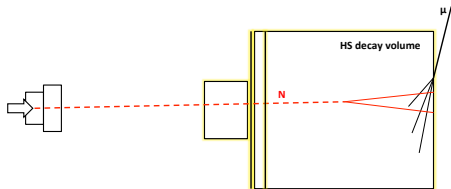
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→ Conclusions

Background sources



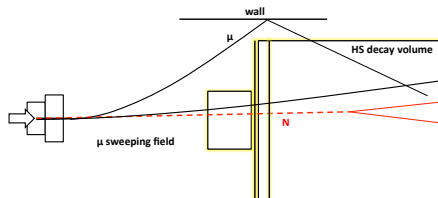
→ **cosmic μ** can scatter on the cavern/vessel walls



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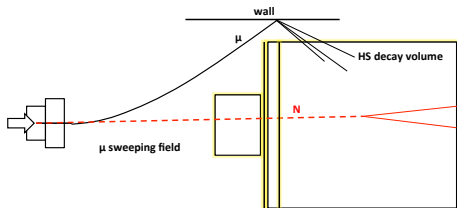
- **cosmic μ** can scatter on the cavern/vessel walls
- **combinatorial combinations** of tracks from different events/vertices



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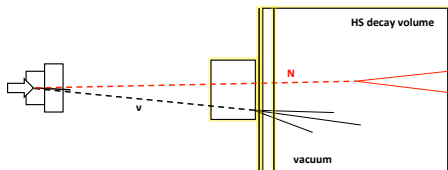
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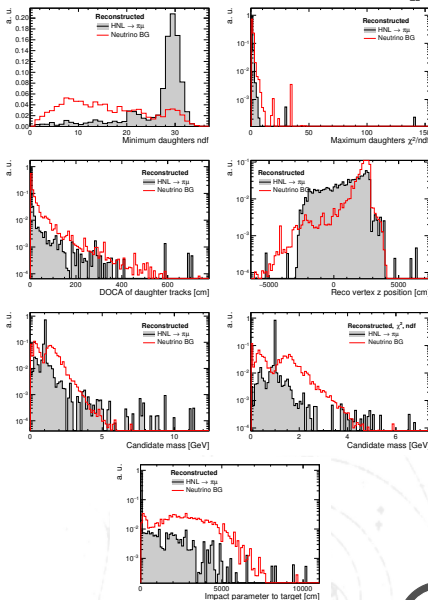
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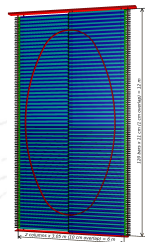
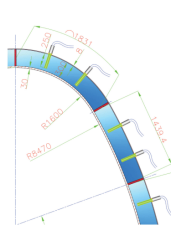
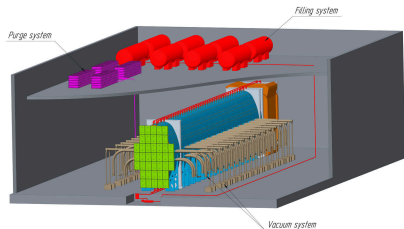
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Shields and background taggers



- **Hadron stopper** after the target
- **Magnetic μ sweeper** creates a ≥ 5 m wide fiducial area
- **ν detector** precedes HS detector and tags upstream particles
- **Upstream VETO** complements its acceptance
- **Straw VETO** tags decays of K_L produced in the ν detector
- **Liquid scintillator** tags interactions crossing the vessel walls
- **Timing detector** reduces combinatorial background





- discard events with activity in the VETO detectors
- select candidates based on the reconstructed direction (must point back to the target)
- require good quality tracks & reconstructed vertex
- event must be fully contained in the fiducial volume, with margins
- we expect < 1 candidate per event \rightarrow cut on multiplicity

Selection efficiency

Sample	Multiplicity	Fiducial vol	Track q.	BG cuts/VETO
$HNL \rightarrow \pi\mu$	97.5 %	76.1 %	87.0 %	94.2 %
$\gamma' \rightarrow \mu\mu$	99.6 %	85.2 %	94.4 %	94.0 %
ν background	79.1 %	21.0 %	6.5 %	0.0 %



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Overall $\lesssim 0.1$ background events / 5 years is attainable!



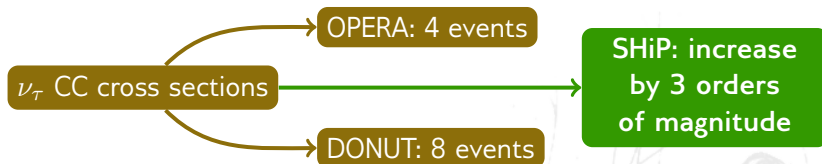
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ν SHiP: a unique opportunity for neutrino physics



High intensity beam dump \implies high flux of neutrinos (all species).

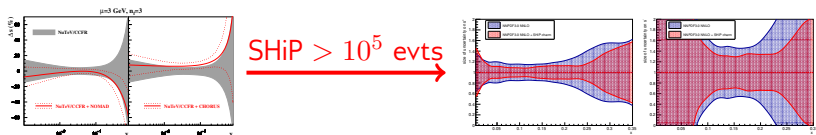
- ν_e tag: electron shower in the ECC brick
- ν_μ tag: μ in the muon filter
- ν_τ tag: separation of τ production and decay vertices
- first observation of $\bar{\nu}_\tau$
- $\nu_\tau/\bar{\nu}_\tau$ cross-section measurements
- First evaluation of the F_4 and F_5 structure functions
- strange quark content from charm production in ν scattering



Tests of perturbative QCD and lepton universality



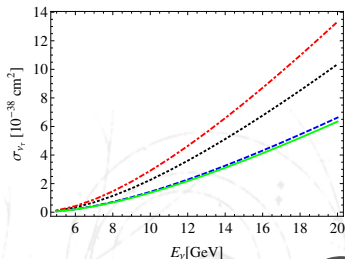
- PDF improvements with ν -nucleon DIS: strange sea quark content currently relies on $\mathcal{O}(5000)$ charm di- μ events:



LHC and SHiP will probe different ranges of x .

- Lepton universality tests:

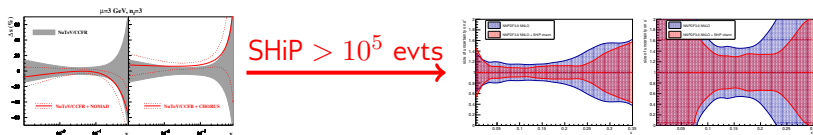
- hints from LHCb, B factories, ...
- DIS σ including BSM: *Liu, Rashed, Datta PRD92(2015)7, 073016*, to compare to σ_{SM}
- results depend on our knowledge of the ν_τ flux!



Tests of perturbative QCD and lepton universality



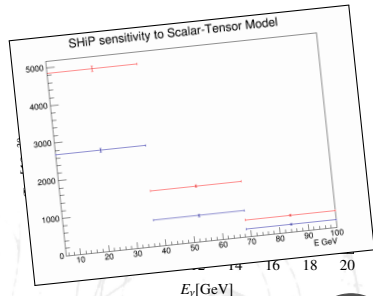
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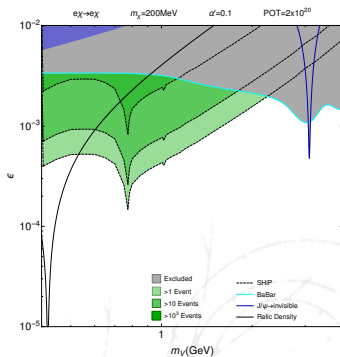
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Detect dark matter from dark photon decay through elastic scattering on electrons: $\chi e^- \rightarrow \chi e^-$. Signature in the emulsion target: a vertex with only e^- coming out. Simulation \implies background from neutrino scattering can be reduced with kinematical selections to 284 events / 5 y.

Dark photon parameter space for $\gamma' \rightarrow$ invisible decays excluded by SHiP at 90% C.L., with such expected background and for $m_\chi = 200$ MeV and $\chi\gamma'$ coupling $\alpha' = 0.1$:



PRD95 (2017) 3, 035006



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Heavy neutral leptons



dark matter
neutrino masses/oscillations
short-baseline neutrino anomalies
matter-antimatter asymmetry

Could be explained with additional, sterile neutrinos

$$\Delta\mathcal{L} = i\bar{N}_I \not{\partial} N_I - \left(F_{\alpha I} \bar{L}_\alpha \mathbf{N}_I \tilde{\Phi} + \frac{M_I}{2} \bar{N}_I^c N_I \right)$$

HNL

Majorana mass term

dimensionless Yukawa couplings

left lepton doublet

Higgs doublet



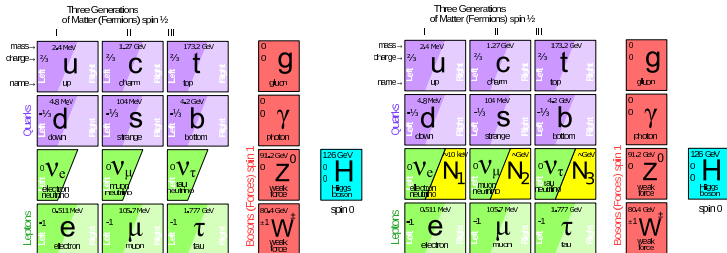
The Majorana mass term induces $\mathcal{L}_{osc} = c_{\alpha\beta} \left(L_{\alpha}^{\dagger} \tilde{\Phi} \right) \left(\tilde{\Phi} L_{\beta} \right) / \Lambda$
 \implies **change flavour** of SM neutrino $\nu_{\alpha} \equiv \tilde{\Phi} L_{\alpha}$

Seesaw mechanism

m_D = Dirac mass term, $(m_D)_{\alpha I} = F_{\alpha I} \langle \Phi \rangle$

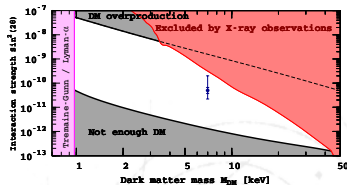
$$(\mathcal{M}_{\nu})_{\alpha\beta} = - \sum_I (m_D)_{\alpha I} \frac{1}{M_I} (m_D)_{\beta I}$$

GeV scale seesaw can generate BAU through HNL oscillations.
Because of $\nu - N$ mixing, HNLs take part in all ν processes with strength reduced by $U_{\alpha I}^2$ and kinematics reflecting m_N .



Suitable values of m_N and U_f^2 allow to simultaneously explain:

- ν oscillations induced by massive states N_2, N_3
- dark matter: N_1 with mass \sim keV
- BAU: leptogenesis due to Majorana mass term

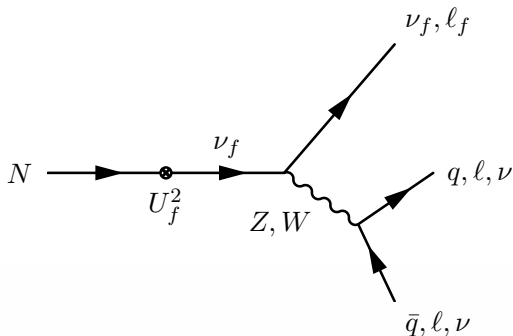


Astrophys. J. 789(2014)13

Phys.Rev.Lett. 113(2014)251301



HNLs can be produced in decays where a ν is replaced by a N (kinetic mixing, **low** \mathcal{BR}). Main neutrino sources in SHiP: c and b mesons.



They can then decay again to SM particles through **mixing** (U^2) with a SM neutrino. This (now **massive**) neutrino can decay to a large amount of final states through emission of a Z^0 or W^\pm boson.



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→ Number of detected HNL events:

$$\Phi(p.o.t) \times \sigma(pp \rightarrow NX) \times \mathcal{P}_{vtx} \times \mathcal{BR}(N \rightarrow visible) \times \mathcal{A}$$

with

$$\sigma(pp \rightarrow NX) \propto \chi_{cc}, \chi_{bb}, U_f^2$$

$$\mathcal{BR}(N \rightarrow visible) \propto U_f^2$$

→ HNL production:

- χ_{cc}, χ_{bb} obtained from simulations (Pythia8)
- $\mathcal{BR}(m_N, U_f^2)$ parametrised according to theory

JHEP 0710 (2007) 015

→ Daughters acceptance (\mathcal{A}):

- HNLs kinematics obtained from simulation
- every decay channel with **detectable** daughters is simulated



Charm mesons are the main source of HNLs in SHiP. Contribution of b mesons for $m_N > 2$ GeV.

- Pythia8 used to retrieve the spectrum of c and b mesons in 400 GeV/ c proton-on-target collisions
- HNL production simulated in kinematically-allowed decay chains:
 - $D \rightarrow K \ell N$
 - $D_s \rightarrow \ell N$
 - $D_s \rightarrow \tau \nu_\tau$ followed by $\tau \rightarrow \mu \nu N$ or $\tau \rightarrow \pi N$
 - $B \rightarrow \ell N$
 - $B \rightarrow D \ell N$
 - $B_s \rightarrow D_s \ell N$
- $\mathcal{BR}(pp \rightarrow NX)$ computed as sum of the BRs of the kinematically-allowed channels

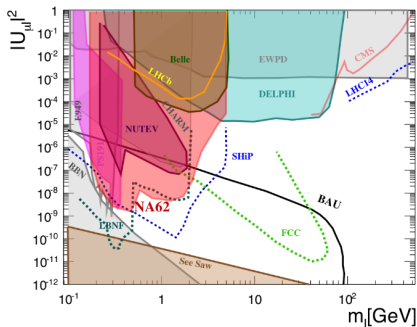
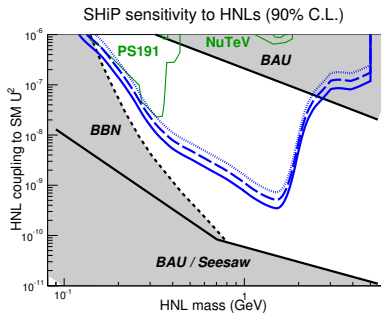


For a given **HNL mass**, its lifetime was computed on the basis of the **widths** of its **kinematically allowed decay channels**:

- $N \rightarrow H^0 \nu$, with $H^0 = \pi^0, \rho^0, \eta, \eta'$
- $N \rightarrow H^\pm \ell^\mp$, with $H = \pi, \rho$
- $N \rightarrow 3\nu$
- $N \rightarrow \ell_i^\pm \ell_j^\mp \nu_j$
- $N \rightarrow \nu_i \ell_j^\pm \ell_j^\mp$
- $N \rightarrow q\bar{q}\nu$ and $N \rightarrow q\bar{q}\ell$ with $q\bar{q}$ hadronising separately

All decay channels into ≥ 2 charged particles considered **detectable**.

SHiP sensitivity to HNLs



- **Best sensitivity** up to the charm kinematic limit
- Significant contribution from B decays
- Nice **synergy with FCC-ee** operated as Z^0 factory
 - very clean signature $Z^0 \rightarrow \nu\bar{\nu}, \nu \rightarrow N \rightarrow \ell^+\ell^-$
- **10 ± 6 background events would not affect sensitivity!**



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- new particles are **light** \implies they must be **singlets** with respect to the gauge group of the SM
- they may couple to different **singlet operators (portals)** of the SM
 - dim 2: hypercharge field, $\epsilon F_{\mu\nu} F'^{\mu\nu}$, vector portal
 - dim 2: Higgs field, $(\alpha_1 \chi + \alpha \chi^2) H^\dagger H$, Higgs/scalar portal
 - dim 2 1/2: Higgs-lepton, $Y H^T \bar{N} L$, neutrino portal
 - dim 4: $AG_{\mu\nu} \epsilon^{\mu\nu\rho\eta} G^{\rho\eta}$, $\partial_\mu A \bar{\psi} \gamma^\mu \gamma^5 \psi$, ..., axion portal
 - SUSY models





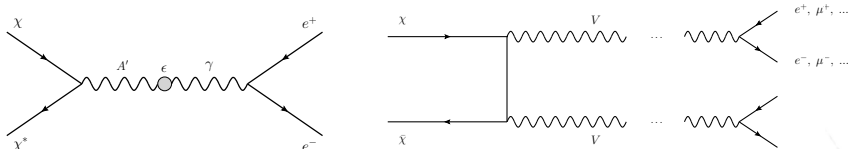
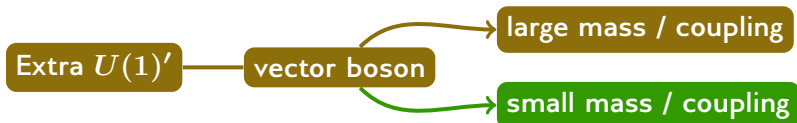
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The vector portal



SM group $SU(3) \times SU(2) \times U(1)$ may descend from a larger group:

$$SU(3) \times SU(2) \times [U(1)]^n$$



- **light** ($m_V \sim \text{MeV-GeV}$) vector particles especially interesting for cosmological reasons

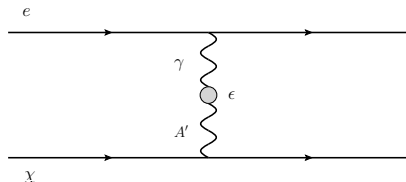
Dark photons and kinetic mixing



$$\mathcal{L} = \underbrace{\mathcal{L}_{\psi,A} + \mathcal{L}_{\chi,A'}}_{\text{QED-like}} - \frac{\epsilon}{2} F_{\mu\nu} F'^{\mu\nu} + \frac{1}{2} m_{A'}^2 (A'_\mu)^2$$

↑ QED fields ↑ $U(1)'$ fields ↑ field strength tensors ↑ mass term

Eq. of motion: $\partial_\mu F^{\mu\nu} = eJ^{(EM)\nu} \implies -\frac{\epsilon}{2} F_{\mu\nu} F'^{\mu\nu} = e\epsilon A'_\mu J^{(EM)\mu}$
 \implies coupling to EM current reduced by ϵ .



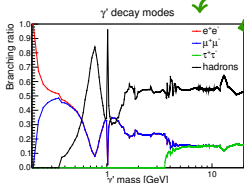
$m_{A'} \rightarrow 0 \implies$
 e.m. charge of $\chi \rightarrow e\epsilon$.

Okun, Sov. Phys. JETP 56 (1982) 502 – Holdom, Phys. Lett. B 166 (1986) 196

Vector portal phenomenology



→ Decay: $\Gamma_{tot} = \Gamma(\ell^+\ell^-) + \Gamma(\text{hadrons}) + \Gamma(\chi\bar{\chi})$



$$\frac{\Gamma_{ll}}{\Gamma_{\chi\bar{\chi}}} \sim \frac{\alpha\epsilon^2}{\alpha_D}, \quad \alpha_D = \text{dark fine structure constant}$$

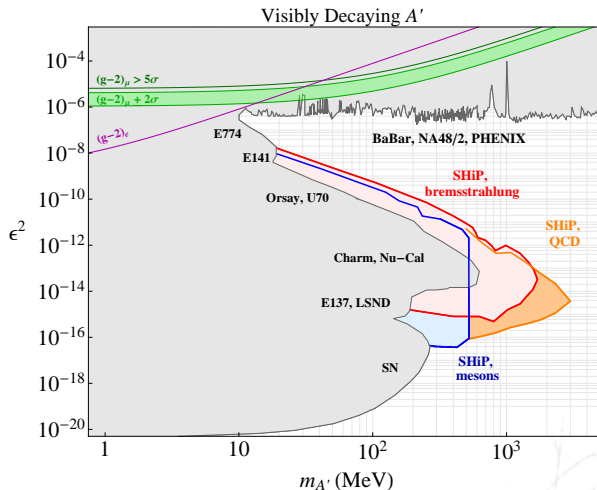
→ Production at SHiP:

- meson decays e.g. $\pi^0 \rightarrow \gamma V$ ($\sim \epsilon^2$) *Phys.Rev. D80(2009)095024*
- p bremsstrahlung on target nuclei $pp \rightarrow ppV$ *Phys.Lett. B731(2014)320-326*
- large $m_V \Rightarrow$ direct QCD production through underlying $q\bar{q} \rightarrow V$,
 $gg \rightarrow V$ (need some more theory work!) *Phys.Rev. D86(2012)035022*

→ Light dark matter at SHiP:

if $\chi\bar{\chi}$ decays dominant $\Rightarrow \chi$ can scatter on electrons $\sim \alpha\alpha_D\epsilon^2$:
dense detector to look for light DM.

SHiP sensitivity: vector portal



Sensitivity studied considering $\Gamma_{tot} = \Gamma(\ell^+\ell^-) + \Gamma(\text{hadrons})$.



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Most general renormalizable \mathcal{L} :

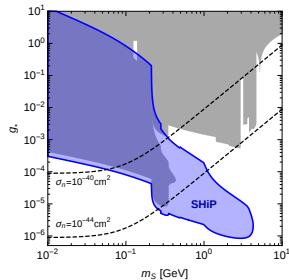
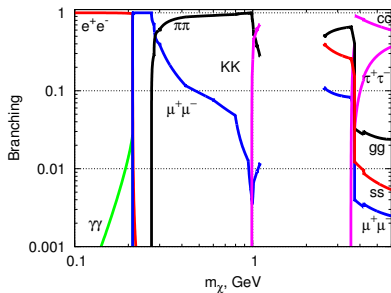
$$\Delta\mathcal{L} = \frac{1}{2}\partial_\mu S\partial^\mu S + (\alpha_1 S + \alpha S^2) (H^\dagger H) + \lambda_2 S^2 + \lambda_3 S^3 + \lambda_4 S^4$$

portal couplings to SM H

lowest-dim gauge
and Lorentz singlet
from SM fields

scalar self-couplings

- $\alpha_1 \neq 0$: S mixes with Higgs after EW symmetry breaking
 \Rightarrow coupling between S and all SM particles
- $\alpha_1 = 0$ (forbidden by exact \mathcal{Z}_2 symmetry): S does not mix with H
 \Rightarrow new particles must be pair-produced



- Existing limits from searches for rare meson decays e.g. $B \rightarrow KS$
- Production: K decays (SHiP efficiency $\approx 0.2\%$) and B decays
- Decay: $S \rightarrow \gamma\gamma, ee, \mu\mu, \pi\pi, KK$

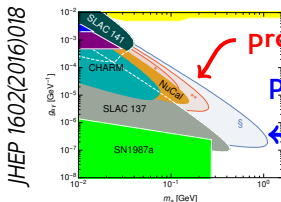


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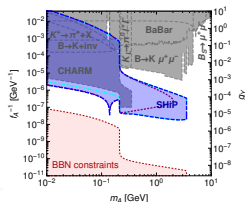
Axion-like particles



- The axion mass m_A is very constrained due to the axial QCD anomaly breaking the PQ symmetry. Other ALPs are not so constrained.
- SHiP can probe ALPs coupled to gauge bosons and to SM fermions:
 - $pp \rightarrow AX$, $A \rightarrow \gamma\gamma$: all neutral, more challenging (left plot)
 - $pp \rightarrow BX$, $B \rightarrow AK$, $A \rightarrow \mu^+\mu^-$ (right plot)



projection NA62
projection SHiP





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- crucial to understand the muon flux and spectrum (2018)
 - muon shield design depends on this
 - simulation validated against CHARM measurements, only low- p_{\perp}
 - use beam test with a NA61/SHINE-like detector and replica SHiP target
- improve current knowledge of inclusive charm production
 - normalisation for HS signal and ν_{τ} cross-section
 - cascade production? Charmed hadron spectrum?
 - use test beam at H4, with smaller target and nuclear emulsions
 - 10^8 p.o.t. \sim 4 weeks data taking at the SPS (LS2)

Scientific Committee Paper	
Report number	CERN-SPSC-2017-020 ; SPSC-EOI-016
Title	Muon-flux measurements for SHiP at H4


Scientific Committee Paper	
Report number	CERN-SPSC-2017-033 ; SPSC-EOI-017
Title	Measurement of associated charm production induced by 400 GeV/c protons

Conclusions



- **General purpose** experiment to look for weakly interacting long lived particles
 - covers previously unexplored regions of the Hidden Sector in several theories
 - covers cosmologically interesting regions
- **Unique** opportunity for ν_τ physics allowing for
 - $\bar{\nu}_\tau$ discovery
 - σ and form factors measurements
 - also dark matter search
- **Complements** LEP/LHC and boosts past experiments sensitivities
 - $\times 10^5$ for HS, $\times 200$ for ν_τ
 - makes best use of existing SPS complex
- **Next steps:**
 - comprehensive design report (2018)
 - μ flux (2018) and χ_{cc} (~ 2019) measurements

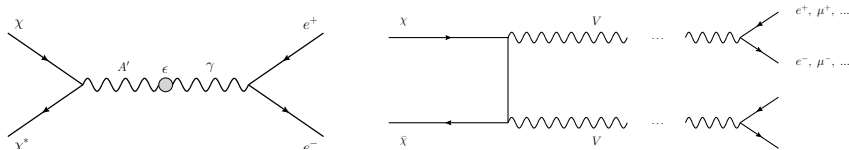




Questions?

- spare slides -

Motivations for light vector particles



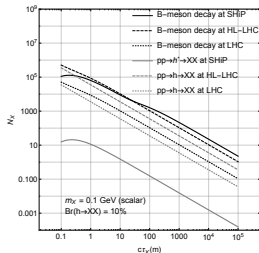
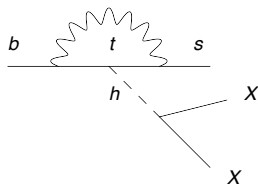
→ Dark matter ($\Omega_{DM} \sim 0.25$):

- light scalar dark matter $m_\chi \sim \text{MeV}$ can solve the positron excess
- WIMP interacting with SM through light mediator ($\chi\bar{\chi} \rightarrow VV \rightarrow \text{SM}$) (hides DM from direct searches)
- non thermal DM (sterile neutrinos)
- DM self-interaction in structure formation ($m_V \sim \text{MeV-GeV}$)

→ Muon $g - 2$:

Light vector particle coupled to muons provides upward correction through one-loop diagram (exchange of A'). *Not minimal model.*

Z_2 scalar portal



→ Higher dimension portals: $\frac{1}{\Lambda} |H|^2 \bar{\psi} \psi$ (dark fermions),
 $\frac{1}{\Lambda^2} m_{Z_D}^2 |H|^2 Z_{D\mu} Z_D^\mu$ (dark gauge boson)

→ decays of the SM Higgs into hidden states

→ at SHiP $E_{CM} \simeq 28 \text{ GeV} < m_H$

Production channels at SHiP:

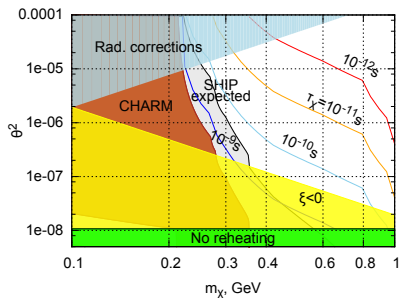
- heavy meson decays (dominant is $B \rightarrow K^{(*)} X X$)
- gluon fusion $pp \rightarrow h^* \rightarrow X X$

→ X decays back to SM with different coupling



- In particle physics, the inflaton is a **scalar field** that couples to SM fields to ensure **re-heating** of the post-inflation Universe (production of particles that thermalize) and transfer of inflaton fluctuations into **adiabatic matter perturbations**.
- $\mathcal{L}_{int} = \alpha S^2 H^\dagger H$, with approx. $10^{-11} < \alpha < 10^{-7}$
 - $\alpha < 10^{-11}$ → inefficient reheating
 - $\alpha > 10^{-7}$ → quantum correction would imply large, scale-dependent density perturbations (\neq observations)

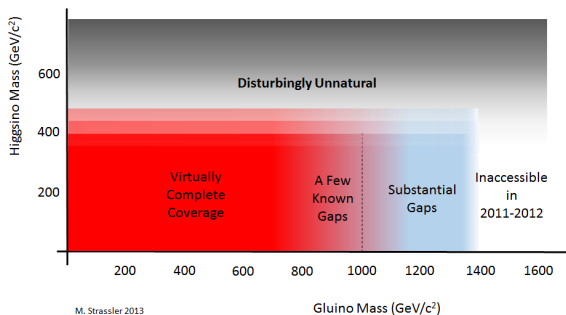
→ **Sensitivity at SHiP** is dominated by the lifetime exponential.



SUSY: where do we stand?



- **SUSY** is one of the most popular options to solve **naturalness**, **grand unification** and **dark matter** (WIMP)
 - $W_{\text{RPC}} = (Y_e)_{ij} L_i H_1 \bar{E}_j + (Y_d)_{ij} Q_i H_1 \bar{D}_j + (Y_u)_{ij} Q_i H_2 \bar{U}_j + \mu H_1 H_2$
- SUSY particles produced in pairs. Accelerator searches significantly constrain “natural” scenarios (e.g. MSSM, fine tuning at $\sim 1\%$).

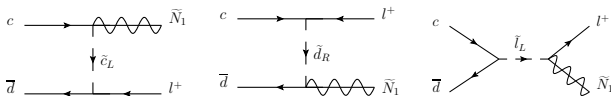


profmatstrassler.com



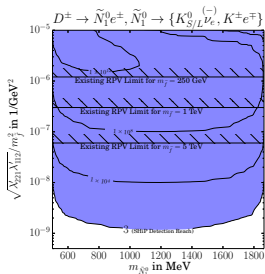
$$\rightarrow \mathbf{W}_{\text{RPV}} = \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k + \kappa_i L_i H_2$$

- The lightest SUSY particle is not anymore stable (no DM)
- Can be searched for at SHiP in D meson decays:

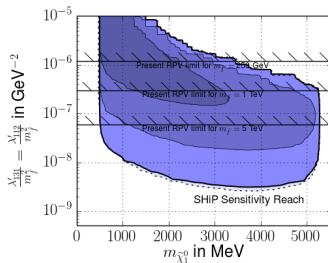


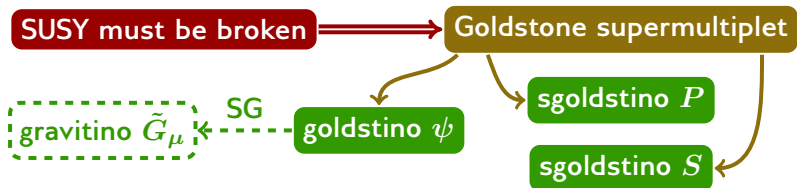
- SHiP sensitivity studied with channels $\tilde{N}_1^0 \rightarrow K^{0(\pm)} \nu$ and $\tilde{N}_1^0 \rightarrow K^\pm \ell^\mp$

SHiP Physics Proposal



arXiv:1511.07436



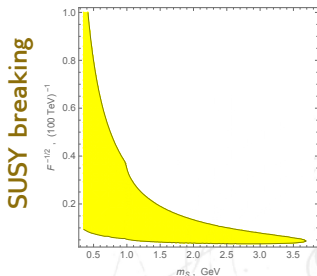


- \tilde{G}_μ (ψ) is R -odd
- P, S are R -even \implies can be singly produced and may decay back to pairs of SM particles
- **at SHiP:**

$$pp \xrightarrow{\text{gluon fusion}} S$$

$$D \rightarrow SX$$

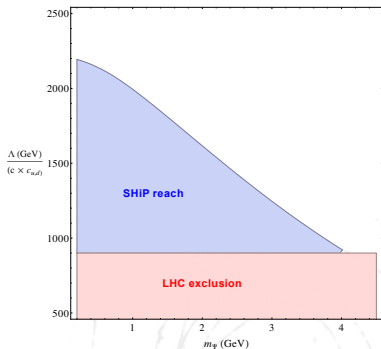
$$S \rightarrow \ell\ell, \pi\pi$$



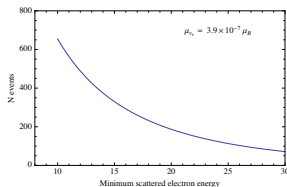
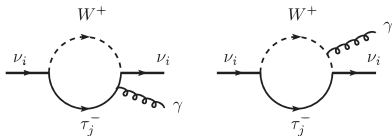
Phys.Rev. D93 (2016) 3



- Dirac fermion (Ψ) split in two Majorana components (χ_1, χ_2)
- interesting dark matter candidate: allows annihilation but appears as Majorana particle for direct and indirect detection purposes
- **Production at SHiP:** $pp \rightarrow \Psi\bar{\Psi}$
- **Decay:** $\chi_2 \rightarrow \ell^+\ell^-\chi_1$



Neutrino magnetic moment



If neutrinos are Dirac particles they can get a magnetic moment:

$$\mu_\nu = \frac{3eG_F m_\nu}{8\pi^2 \sqrt{2}} \simeq (3.2 \times 10^{-19}) \frac{m_\nu}{1 \text{ eV}} \mu_B$$

BSM can enhance μ_ν .

(E.g.: *Shrock, Nucl.Phys. B206 (1982) 359*)

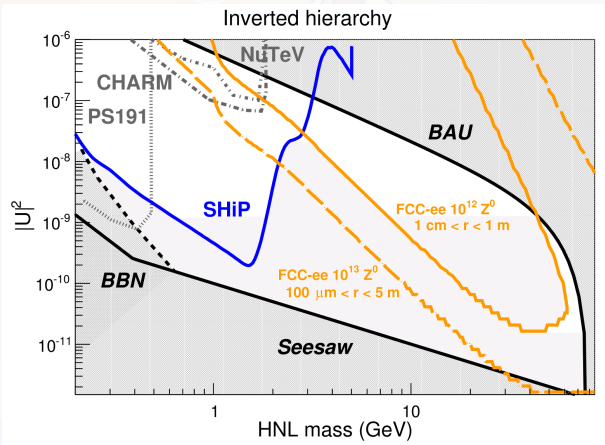
$$e\nu \rightarrow e\nu \implies \left. \frac{dN}{dE_e} \right|_{\mu_\nu} = \frac{\pi\alpha^2 \mu_\nu^2}{m_e^2} \left(\frac{1}{E_e} - \frac{1}{E_\nu} \right)$$

Remove BG from νN scattering: $\theta_{\nu e}^2 < 2m_e/E_e \implies$ sensitivity:

$N_{evt} \sim 4.3 \times 10^{15} \mu_\nu^2 / \mu_B^2$. Prev. limits from 10^{-7} (ν_τ) to 10^{-11} (ν_e).



HNLs at future colliders



<http://arxiv.org/abs/1411.5230>

<http://arxiv.org/abs/1503.08624>

Sensitivity with non-zero background

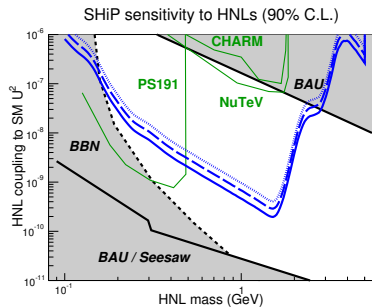
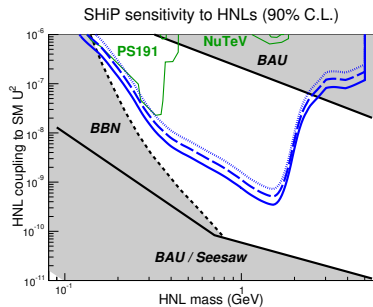


Figure: Variation of the sensitivity contours for scenarios II (left) and IV (right) as a function of the HNLs background estimates. The solid blue curve represents the 90% C.L. upper limit assuming 0.1 background events in 2×10^{20} proton-target collisions. The dashed blue curve assumes 10 background events. The dotted blue curve assumes a systematic uncertainty of 60% on the level of background, i.e. 10 ± 6 background events.

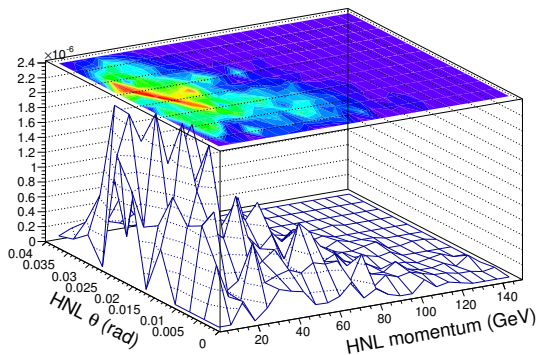
Estimating SHiP's physics reach



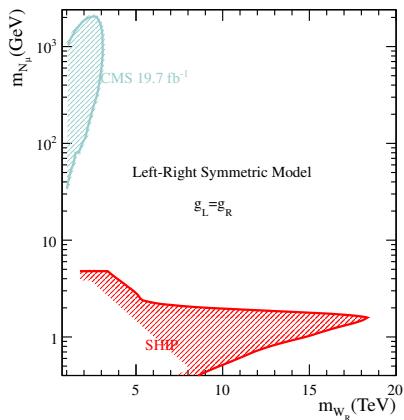
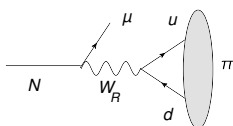
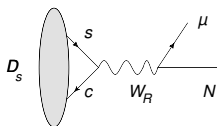
$$\Phi(p.o.t) \times \mathcal{BR}(pp \rightarrow NX) \times \mathcal{P}_{vtx} \times \mathcal{BR}(N \rightarrow \text{visible}) \times A$$

- HNL's momentum and angle are stored in a binned PDF
- HNL spectra are re-weighted by the probability $\mathcal{P}_{vtx}(p, \theta | m_N, U_f^2) \sim \int_V e^{-l/\gamma c\tau} dl$
- Integral of the weighted PDF gives the total probability $\mathcal{P}_{vtx}(m_N, U_f^2)$ that HNLs leave a vertex in SHiP's fiducial volume

Weighted PDF for model 2, $m_N = 1.8$ GeV, $U_\mu = 10^{-9}$



Sensitivity in the Left-Right symmetric model



- SHiP limits on m_{W_R} can be extracted from the HNL limits by $|U_{\mu I}|^2 \rightarrow (m_{W_L}/m_{W_R})^4$
- LHC can perform direct searches on both W_R and N_R
- SHiP can only look for N_R , but in a domain inaccessible to LHC
- based on CMS, *Eur. Phys. J. C* 74 (2014) 3149, and Helo, Hirsch, Kovalenko, *Phys.Rev. D*89 (2014) 073005



- ν **oscillations** provide evidence of LFV in the neutral sector
- **LFV** in charged sector foreseen with $BR \sim \mathcal{O}(10^{-40})!$
- **New physics** models can enhance these BR s
 - in **seesaw** models charged LFV can happen in tree or loop diagrams
 - $\ell \rightarrow 3\ell'$ generally favoured with respect to $\ell \rightarrow \ell' \gamma$ (type 2 and 3 seesaw)
- $\ell \rightarrow 3\ell'$ related by unitarity to $Z^0, h, V \rightarrow \ell^+ \ell'^-$ and $\ell \rightarrow \ell'$ conversion in nuclei (most stringent limits so far by SINDRUM II)
 - $\tau \rightarrow 3\mu$ and $\mu \rightarrow 3e$ can provide better limits than direct searches e.g. for $\phi \rightarrow e\mu, J/\Psi \rightarrow e\mu$
 - $BR(\tau \rightarrow 3\mu) < 1.2 \times 10^{-8}$ (BaBar, Belle, LHCb) *HFAG, arXiv:1412.7515*
- **SHiP** will collect 3×10^{15} τ in the forward region
 - requires **changes to conceptual design** (upgrade):
 - 1 mm W target: $100\times$ less τ , but decaying outside target
 - LHCb VELO + Si tracker + hadron absorber + μ spectrometer
 - **sensitivity** $\sim 10^{-10} / \sqrt{N_{\text{targets}}}$



The Hidden Sector

$$L_{world} = L_{SM} + L_{mediation} + L_{HS}$$

- **Neutrino portal:** new Heavy Neutral Leptons coupling with Yukawa coupling, $L_{NP} = F_{\alpha I}(\bar{L}_\alpha \tilde{\Phi}) N_I$
- **Vector portal:** massive dark photon coupling through loops of particles charged both under $U(1)$ and $U'(1)$: $L_{VP} = \epsilon F'_{\mu\nu} F^{\mu\nu}$
- **Scalar portal:** light scalar mixing with the Higgs $L_{SP} = (\lambda_i S_i^2 + g_i S_i) \bar{\Phi} \Phi$
- **Axion portal:** axion-like particles, $L_{AP} = \frac{A}{4f_A} \epsilon^{\mu\nu\lambda\rho} F_{\mu\nu} F_{\lambda\rho}$
- **SUSY:** neutralino, sgoldstino, gaugino...

Models	Final states
Neutrino portal, SUSY neutralino	$\ell^\pm \pi^\mp, \ell^\pm K^\mp, \ell^\pm \rho^\mp, \rho^\pm \rightarrow \pi^\pm \pi^0$
Vector, scalar, axion portals, SUSY sgoldstino	$\ell^+ \ell^-$
Vector, scalar, axion portals, SUSY sgoldstino	$\pi^+ \pi^-, K^+ K^-$
Neutrino portal, SUSY neutralino, axino	$\ell^+ \ell^- \nu$
Axion portal, SUSY sgoldstino	$\gamma\gamma$
SUSY sgoldstino	$\pi^0 \pi^0$



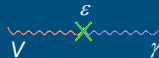
New Physics prospects in Hidden Sector



Standard Model portals:

D = 2: Vector portal

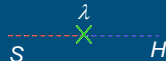
- Kinetic mixing with massive dark/secluded/paraphoton V : $\frac{1}{2}\epsilon F_{\mu\nu}^{SM} F_{HS}^{\mu\nu}$



- Motivated in part by idea of “mirror world” restoring left and right symmetry, constituting dark matter, g-2 anomaly, ...
- Production: proton bremsstrahlung, direct QCD production $q\bar{q} \rightarrow V, qg \rightarrow Vq$, meson decays ($\pi^0, \eta, \omega, \eta', \dots$)

D = 2: Scalar portal

- Mass mixing with dark singlet scalar χ : $(gS + \lambda S^2)H^\dagger H$



- Mass to Higgs boson and right-handed neutrino, inflaton, dark phase transitions BAU, dark matter, “dark naturalness”,...
- Production: Direct $p + target \rightarrow X + S$, meson decays e.g. $B \rightarrow KS, K \rightarrow \pi S$

D = 5/2: Neutrino portal

- Mixing with right-handed neutrino N (Heavy Neutral Lepton): $Y_{i\ell} H^\dagger \bar{N}_i L_\ell$



- Neutrino oscillation, baryon asymmetry, dark matter
- Production: Leptonic, semi-leptonic decays of heavy hadrons

D = 4: Axion portal

- Mixing with Axion Like Particles, pseudo-scalars pNGB, axial vectors a : $\frac{a}{F} G_{\mu\nu} \tilde{G}^{\mu\nu}, \frac{\partial_\mu a}{F} \bar{\psi} \gamma_\mu \gamma_5 \psi$, etc

- Generically light pseudo-scalars arise in spontaneous breaking of approximate symmetries at a high mass scale F
- Extended Higgs, SUSY breaking, dark matter, possibility of inflaton,...
- Production: Primakoff production, mixing with pions and heavy meson decays

And higher dimensional operator portals

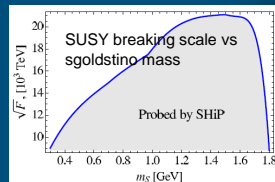
- Chern-Simons portal (vector portal)



New Physics prospects in Hidden Sector

- **SUper-SYmmetric “portals”**

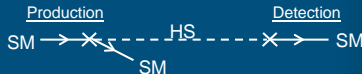
- Some of SUSY low-energy parameter space open to complementary searches
- Sgoldstino S(P) : $\frac{M_{\gamma\gamma}}{F} S F^{\mu\nu} F_{\mu\nu}$
- Neutralino in R-Parity Violating SUSY
- Hidden Photinos, axinos and saxions....



→ A very large variety of models based on these or mixtures thereof

○ Two search methods:

1. “Indirect detection” through portals in (missing mass)
2. “Direct detection” through both portals in and out



→ *SHiP has significant sensitivity to all of these!*

Assumption invisible decay width $\chi\bar{\chi}$ is absent or sub-dominant, $m_\chi > \frac{1}{2}m_{portal}$, where χ hidden sector particle



Sterile Neutrinos

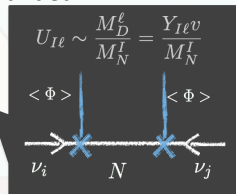
Fermions get mass via the Yukawa couplings:

$$-\mathcal{L}_{\text{Yukawa}} = Y_{ij}^d \overline{Q}_{Li} \phi D_{Rj} + Y_{ij}^u \overline{Q}_{Li} \tilde{\phi} U_{Rj} + Y_{ij}^\ell \overline{L}_{Li} \phi E_{Rj} + \text{h.c.},$$

If we want the same coupling for neutrinos, we need right-handed (sterile) neutrinos... the most generic Lagrangian is

$$\mathcal{L}_N = i \overline{N}_i \partial_\mu \gamma^\mu N_i - \frac{1}{2} M_{ij} \overline{N}_i^c N_j - Y_{ij}^\nu \overline{L}_{Li} \tilde{\phi} N_j$$

Kinetic term
Majorana mass term
Yukawa coupling



Seesaw mechanism:

$$\mathcal{V} = (\nu_{Li}, N_j)$$

$$-\mathcal{L}_{M_\nu} = \frac{1}{2} \overline{\mathcal{V}} M_\nu \mathcal{V} + \text{h.c.}$$

$$M_\nu = \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix}$$

$$\lambda_\pm = \frac{M_N \pm \sqrt{M_N^2 + 4M_D^2}}{2}$$

if $M_N \gg M_D$:

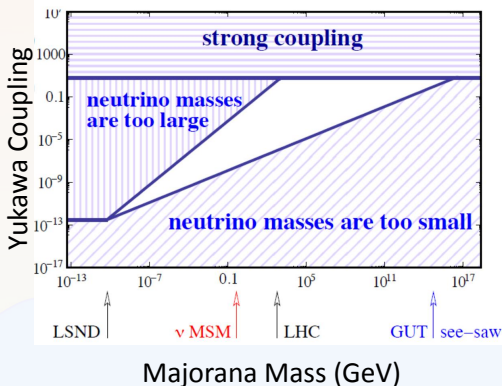
$$\lambda_- \sim \frac{M_D^2}{M_N}$$

$$\lambda_+ \sim M_N$$



Sterile neutrino masses

Seesaw formula $m_D \sim Y_{I\alpha} \langle \phi \rangle$ and $m_\nu = \frac{m_D^2}{M}$



- Assuming $m_\nu = 0.1\text{eV}$
- if $Y \sim 1$ implies $M \sim 10^{14}\text{GeV}$
- if $M_N \sim 1\text{GeV}$ implies $Y_\nu \sim 10^{-7}$

remember $Y_{top} \sim 1$. and $Y_e \sim 10^{-6}$

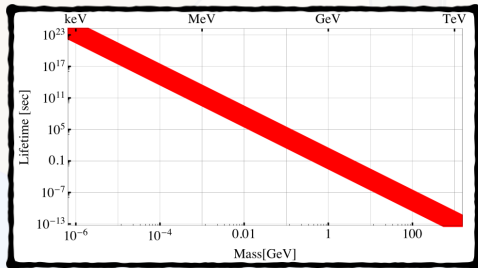
If we want to explain the smallness of neutrino masses (in a natural way) the mass of sterile neutrinos should be at least at the GeV scale



Constraints on N_1

The decay mode $N \rightarrow \nu\nu\nu$ is always present

$$LT = \left(\frac{U^2 G_F^2 M_N^5}{86\pi^3} \right)^{-1} \simeq 0.3 \left(\frac{1\text{GeV}}{M_N} \right)^4 \text{ sec}$$



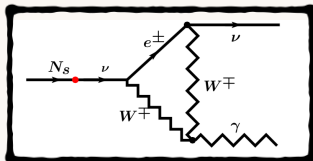
This gives an upper bound for the mass of the sterile neutrino Dark Matter

- $M_N \sim 1\text{KeV} \implies \tau_N \sim 10^{24}\text{sec}$
- $\frac{\text{Age of the Universe}}{\tau_N} \sim 10^{-6}$

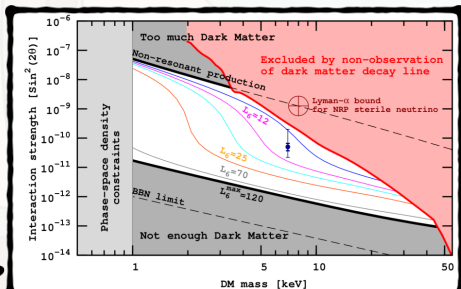


Constraints on N_1

DM sterile neutrinos decay subdominantly as $N_1 \rightarrow \nu \gamma$ with a branching ratio $\mathcal{B}(N_1 \rightarrow \gamma \nu) \sim \frac{1}{123}$



Discussion in the community, not yet clear if this is a “good” signal, needs confirmation



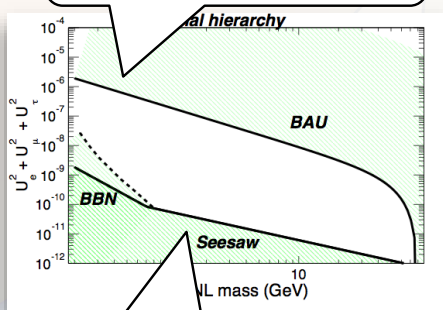
Bulbul et al. 2014 (arXiv:1402.2301)

Boyarisky et al. 2014 (arXiv:1402.4119)



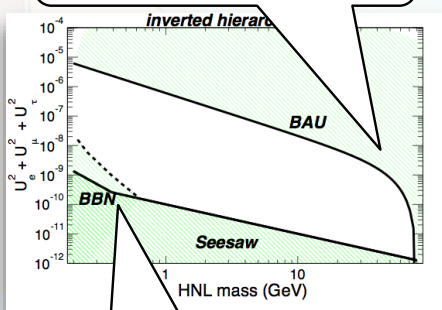
Constraints on N_2, N_3

If U^2 is too large, $N_{2,3}$ are in **thermal equilibrium** during the expansion of the Universe



The **seesaw** limit defines the region where $N_{2,3}$ can explain the observed active neutrino Δm^2

At $M_N \geq M_W$ the rate is **enhanced** by $N \rightarrow Wl$ leading to stronger constraints on U^2



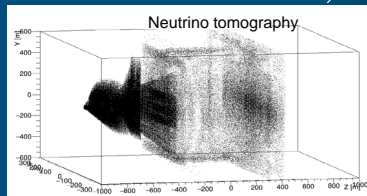
If $\tau(N_2, N_3) < 0.1$ s, they cannot affect the **Big Bang nucleosynthesis**



Backgrounds with TP detector



Background source	Decay modes
ν or μ + nucleon $\rightarrow X + K_L$	$K_L \rightarrow \pi e \nu, \pi \mu \nu, \pi^+ \pi^-, \pi^+ \pi^- \pi^0$
ν or μ + nucleon $\rightarrow X + K_S$	$K_S \rightarrow \pi^0 \pi^0, \pi^+ \pi^-$
ν or μ + nucleon $\rightarrow X + \Lambda$	$\Lambda \rightarrow p \pi^-$
n or p + nucleon $\rightarrow X + K_L$, etc	as above



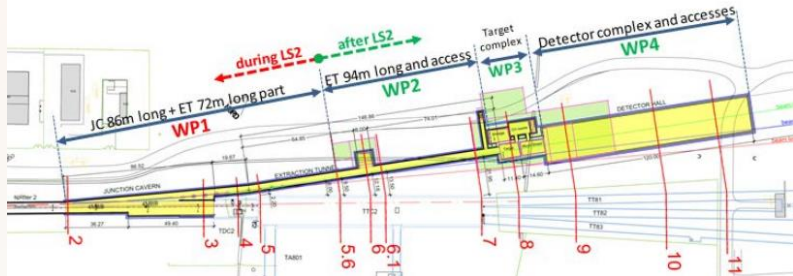
- Background summary: no evidence for any irreducible background**

- No events selected in MC \rightarrow Expected background UL @ 90% CL

Background source	Stat. weight	Expected background (UL 90% CL)
ν-induced		
$2.0 < p < 4.0$ GeV/c	1.4	1.6
$4.0 < p < 10.0$ GeV/c	2.5	0.9
$p > 10$ GeV/c	3.0	0.8
$\bar{\nu}$-induced		
$2.0 < p < 4.0$ GeV/c	2.4	1.0
$4.0 < p < 10.0$ GeV/c	2.8	0.8
$p > 10$ GeV/c	6.8	0.3
Muon inelastic	0.5	4.6
Muon combinatorial	–	<0.1
Cosmics		
$p < 100$ GeV/c	2.0	1.2
$p > 100$ GeV/c	1600	0.002



NA work packages



- Preparation of facility in four well-defined quasi-independent work packages
 - WP1: Junction cavern + 70m beam line for clearance during operation (21 months)
 - WP2 : Rest of beam line (12 months)
 - WP3 : Target complex (12 months)
 - WP4 : Experiment facility (18 months)
- ➔ Only WP1 has to be done during a stop of the North Area only
- ➔ WP1 associated with cool down, removal and re-installation of services and beam line (24-27 months)
- ➔ Construction of facility has no interference with operation of SPS and LHC at any time



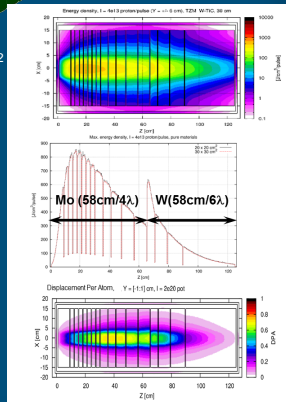
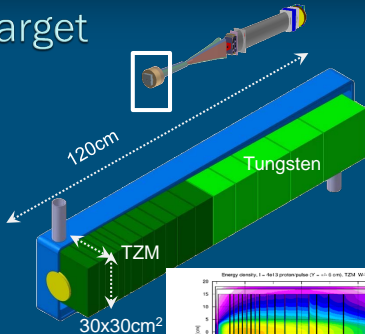
SHiP target



Design considerations with 4×10^{13} p / 7s

→ 355 kW average, 2.56 MW during 1s spill

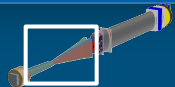
- High temperature
 - Compressive stresses
 - Atomic displacement
 - Erosion/corrosion
 - Material properties as a function of irradiation
 - Remote handling (Initial dose rate of 50 Sv/h...)
- Hybrid solution: Mo allow TZM (4λ) + W(6λ)



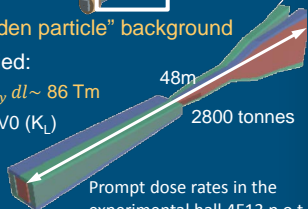
	DONUT ¹⁾	CHARM ²⁾	SHiP
Target material	W-alloy	Cu (variable ρ)	TZM + pure W
Momentum (GeV/c)	800	400	400
Intensity	0.8×10^{13}	1.3×10^{13}	4×10^{13}
Pulse length (s)	20	23×10^{-6}	1
Rep. rate (s)	60	~10	7.2
Beam energy (kJ)	1020	830	2560
Avg. beam power (spill) (kW)	51	3.4×10^7 (fast)	2560
Avg. beam power (SC) (kW)	17	69	355
POT	Few 10^{17}	Few 10^{18}	2×10^{20}



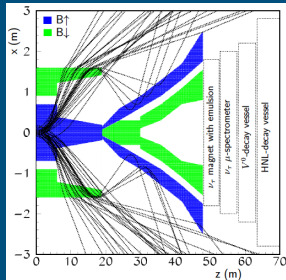
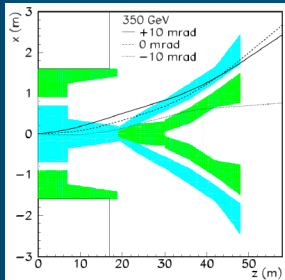
Active muon shield



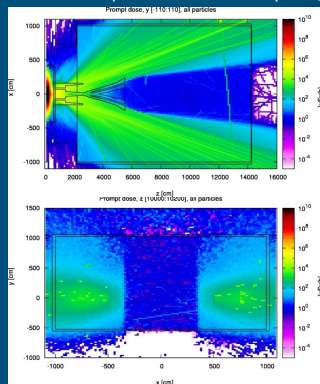
- Muon flux limit driven by emulsion based ν -detector and “hidden particle” dose background
- Passive and magnet sweeper/passive absorber options studied:
 - Conclusion: Shield based entirely on magnetic sweeping with $\int B_y dl \sim 86 \text{ Tm}$
 - $< 7 \times 10^3$ muons / spill ($E_\mu > 3 \text{ GeV}$) which can potentially produce $V0$ (K_L)
 - Negligible occupancy



Prompt dose rates in the experimental hall 4E13 p.o.t. / 7s



→ Challenges: flux leakage, constant field profile, modelling magnet shape





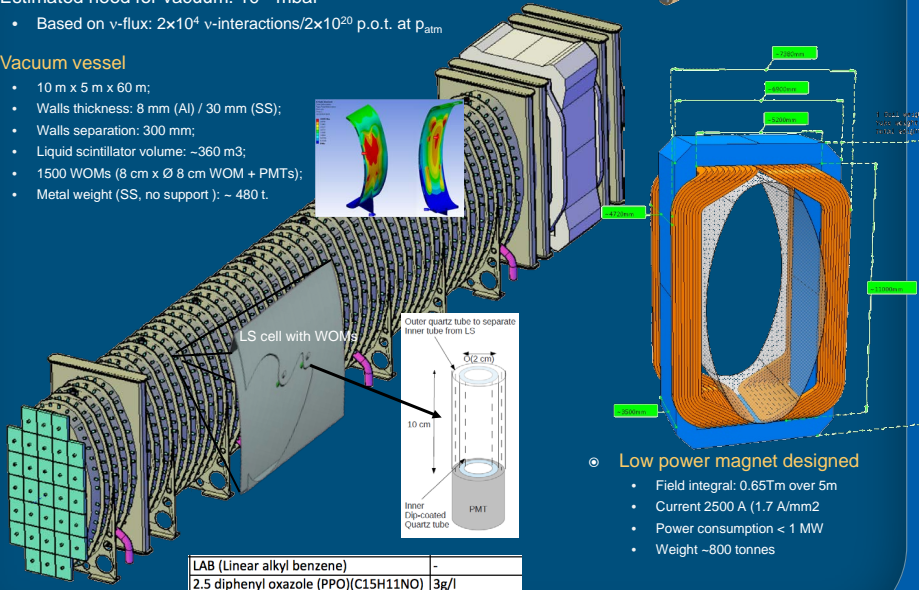
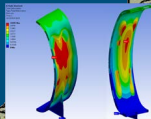
TP: Vessel and spectrometer magnet



- Estimated need for vacuum: 10^{-3} mbar
 - Based on v -flux: 2×10^4 v -interactions/ 2×10^{20} p.o.t. at p_{atm}

Vacuum vessel

- 10 m x 5 m x 60 m;
- Walls thickness: 8 mm (Al) / 30 mm (SS);
- Walls separation: 300 mm;
- Liquid scintillator volume: ~ 360 m³;
- 1500 WOMs (8 cm x \varnothing 8 cm WOM + PMTs);
- Metal weight (SS, no support): ~ 480 t.



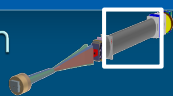
Low power magnet designed

- Field integral: 0.65Tm over 5m
- Current 2500 A (1.7 A/mm²)
- Power consumption < 1 MW
- Weight ~ 800 tonnes

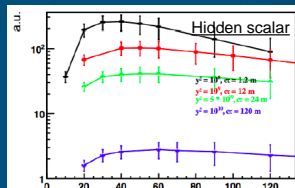
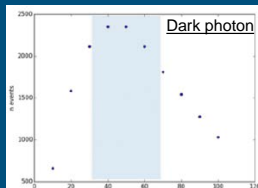
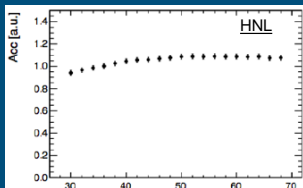
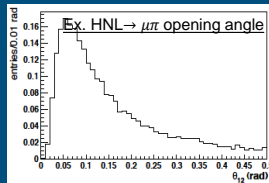
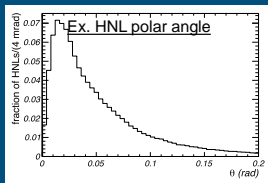
LAB (Linear alkyl benzene)	-
2.5 diphenyl oxazole (PPO)(C ₁₅ H ₁₁ NO)	3g/l



HS detector optimization



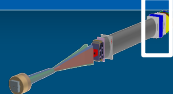
- Optimization of geometrical acceptance for a given E_{beam} and Φ_{beam}
 - Hidden particle **lifetime** (~flat for longlived)
 - Hidden particle **production angles** (~distance and transversal size)
 - Hidden particle **decay opening angle** (~length and transversal size)
 - Muon flux** (~distance and acceptable occupancy)
 - Background** (~detector time and spatial resolution)
 - Evacuation** in decay volume / **technically feasible** size ~ W:5m x H:10m



→ Acceptance saturates ~40m – 50m



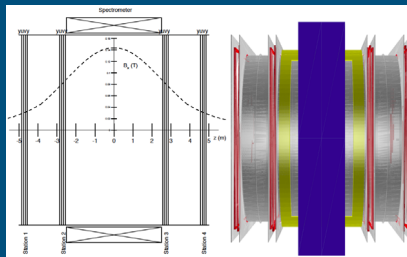
HS tracking system



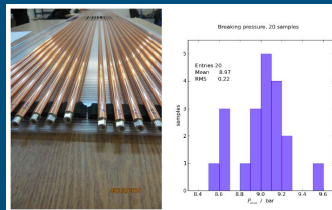
NA62-like straw detector

Parameter	Value
Straw	
Length of a straw	5 m
Outer straw diameter	9.83 mm
Straw wall (PET, Cu, Au)	
PET foil thickness	36 μm
Cu coating thickness	50 mm
Au coating thickness	20 mm
Wire (Au-plated Tungsten) diameter	30 μm
Straw arrangement	
Number of straws in one layer	568
Number of layers per plane	2
Straw pitch in one layer	17.6 mm
Y extent of one plane	~ 10 m
Y offset between straws of layer 1&2	8.8 mm
Z shift from layer 1 to 2	11 mm
Number of planes per view	2
Y offset between plane 1&2	4.4 mm
Z shift from plane 1 to 2	26 mm
Z shift from view to view	100 mm
Straw station	
Number of views per station	4 (Y-U-V-Y)
Stereo angle of layers in a view Y,U,V	0, 5, -5 degrees
Z envelope of one station	~ 34 cm
Number of straws in one station	9088
Straw tracker	
Number of stations	4
Z shift from station 1 to 2 (3 to 4)	2 m
Z shift from station 2 to 3	5 m
Number of straws in total	36352

Horizontal orientation of 5m straws



First production of 5m straws at JINR



Straws in test beam 2016

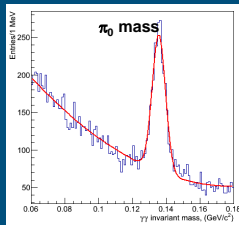
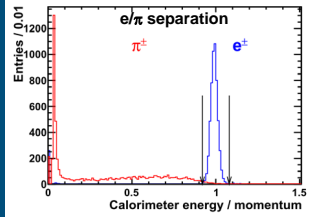
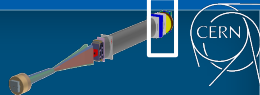
- Study sagging effects and compensation
- Read out of signal, attenuation / two-sided readout

Upstream straw veto may be based on same technology

JINR Dubna (NA62, SHiP): Straws
St Petersburg (CMS, SHiP): Infra



PID performance



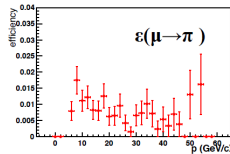
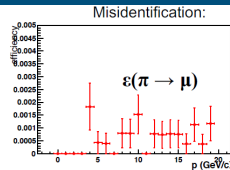
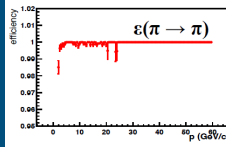
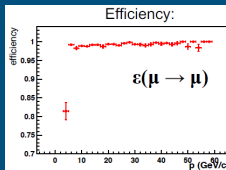
Electron efficiency >98%
 Pion contamination: <2%
 Neutral pion mass resolution: 5 MeV

Muon misid with ECAL+HCAL

Rejection factor for $\epsilon_\mu = 95\%$

Energy, GeV	E+H1+H2
1.0	23
1.5	32
2.0	50
2.7	120
3.0	160
5.0	210
10.0	250

2/07/2015



→ ECAL (July), HCAL (September), MUON (October) in test beam 2015 on PS and SPS